

Electronic Supplementary Material

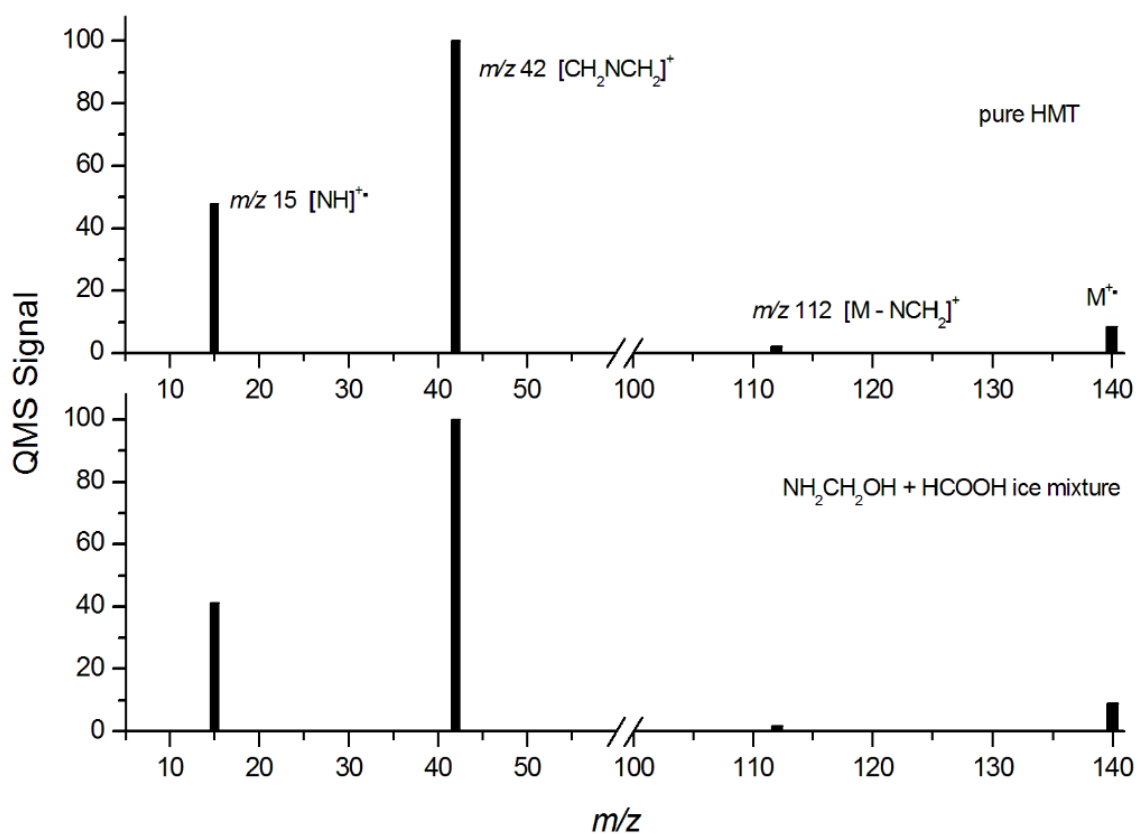
**Title:** Mechanism of hexamethylenetetramine (HMT) formation in interstellar conditions

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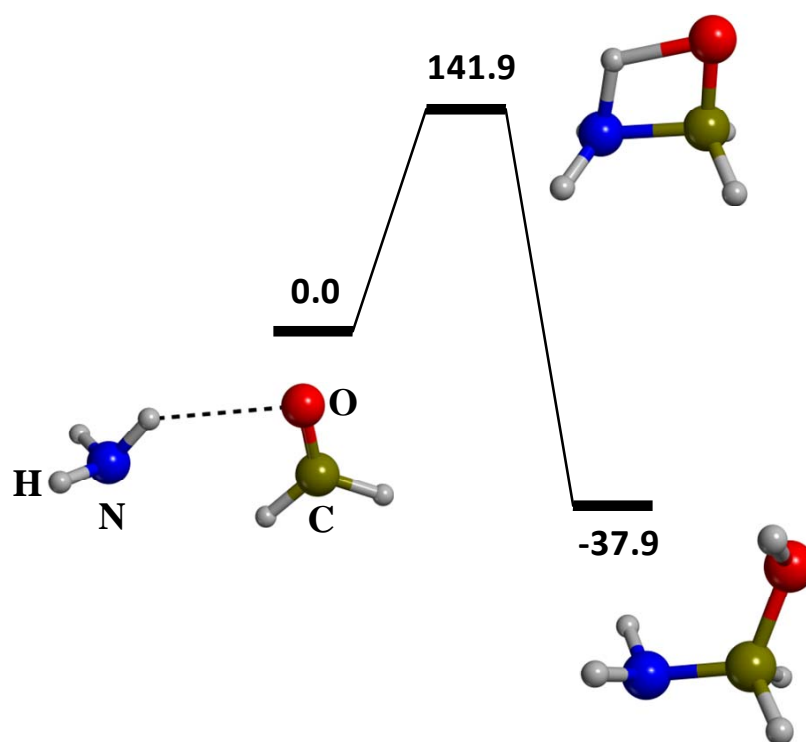
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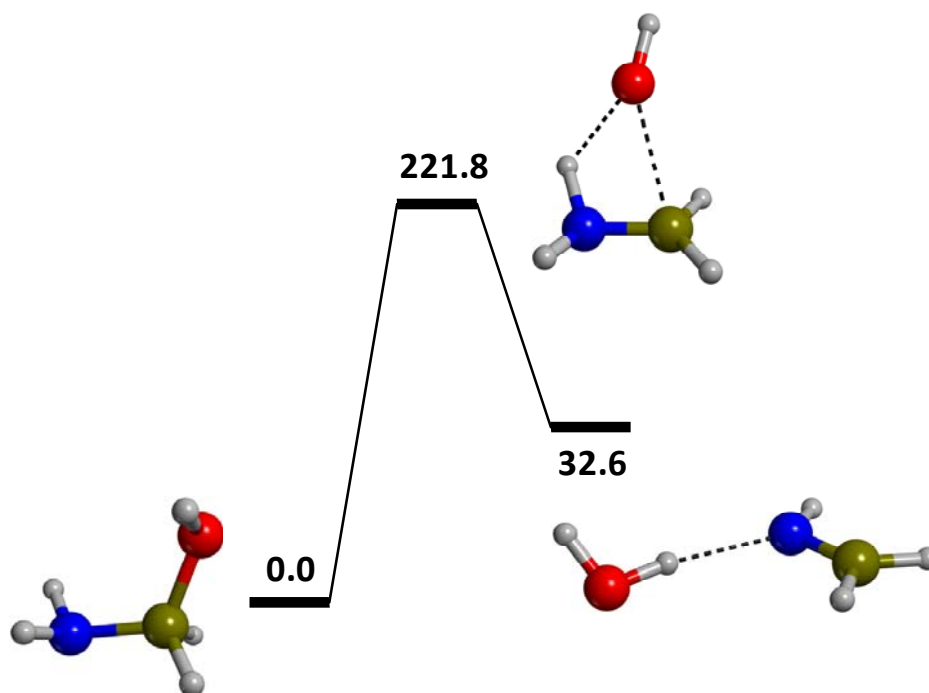
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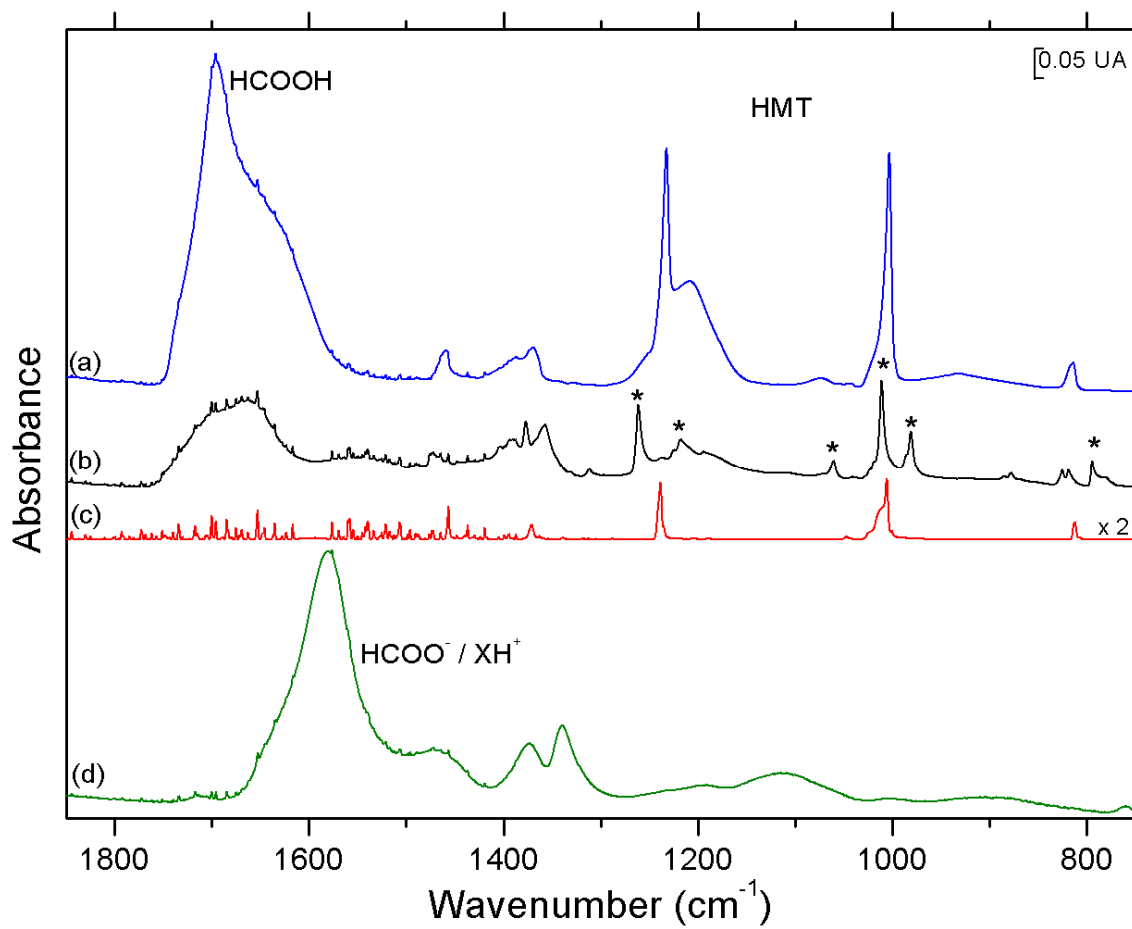
**Figure S1.** QMS signal recorded for pure HMT at 260 K (top panel) and from the warming of NH<sub>2</sub>CH<sub>2</sub>OH : HCOOH ice mixture at 360 K (bottom panel).



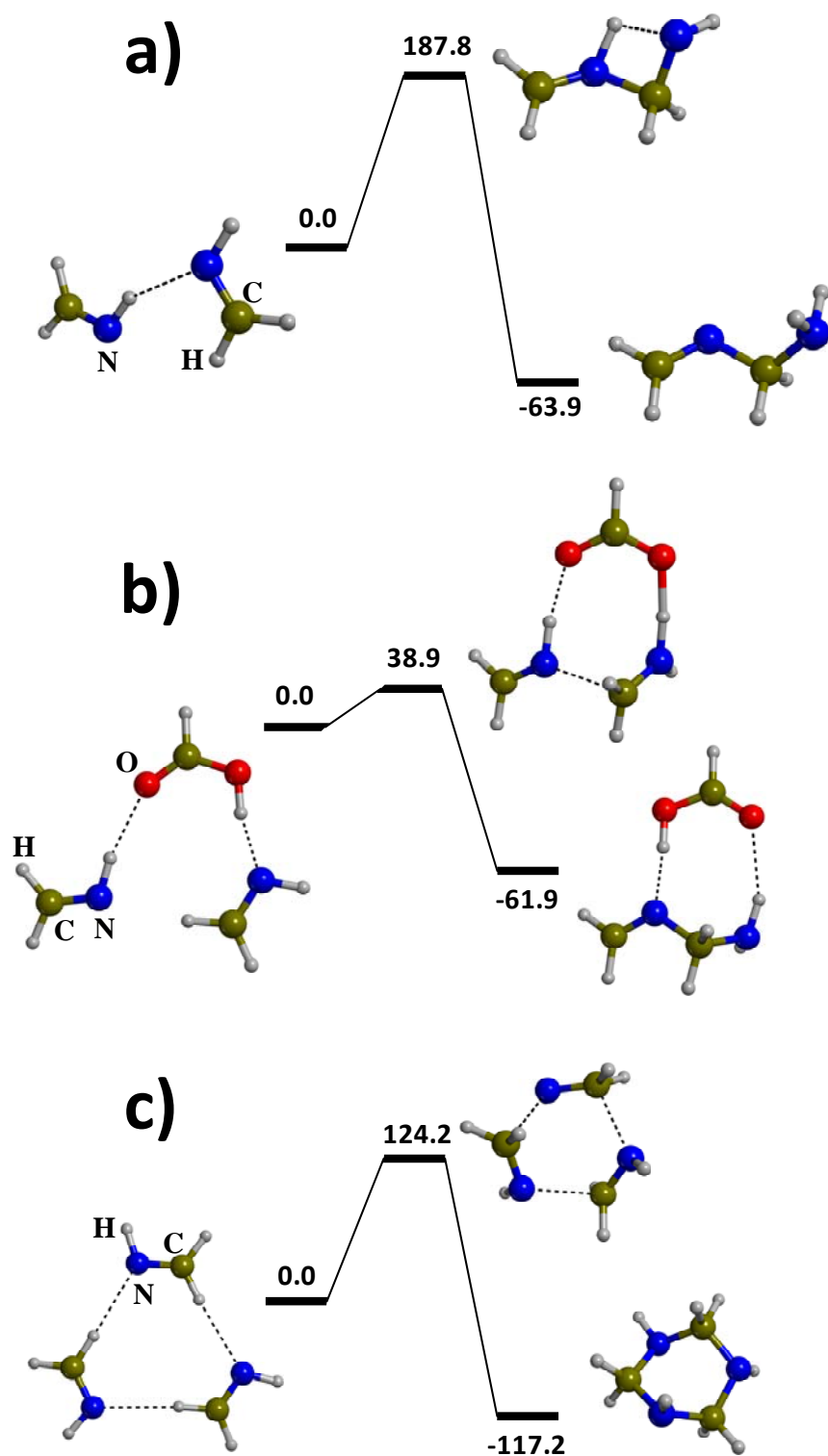
**Figure S2.** B3LYP/6-311++G(d,p) potential energy profile for the gas phase formation of aminomethanol adopting the reaction of  $\text{NH}_3 + \text{H}_2\text{C}=\text{O} \rightarrow \text{NH}_2\text{CH}_2\text{OH}$ . Units in  $\text{kJ mol}^{-1}$ .



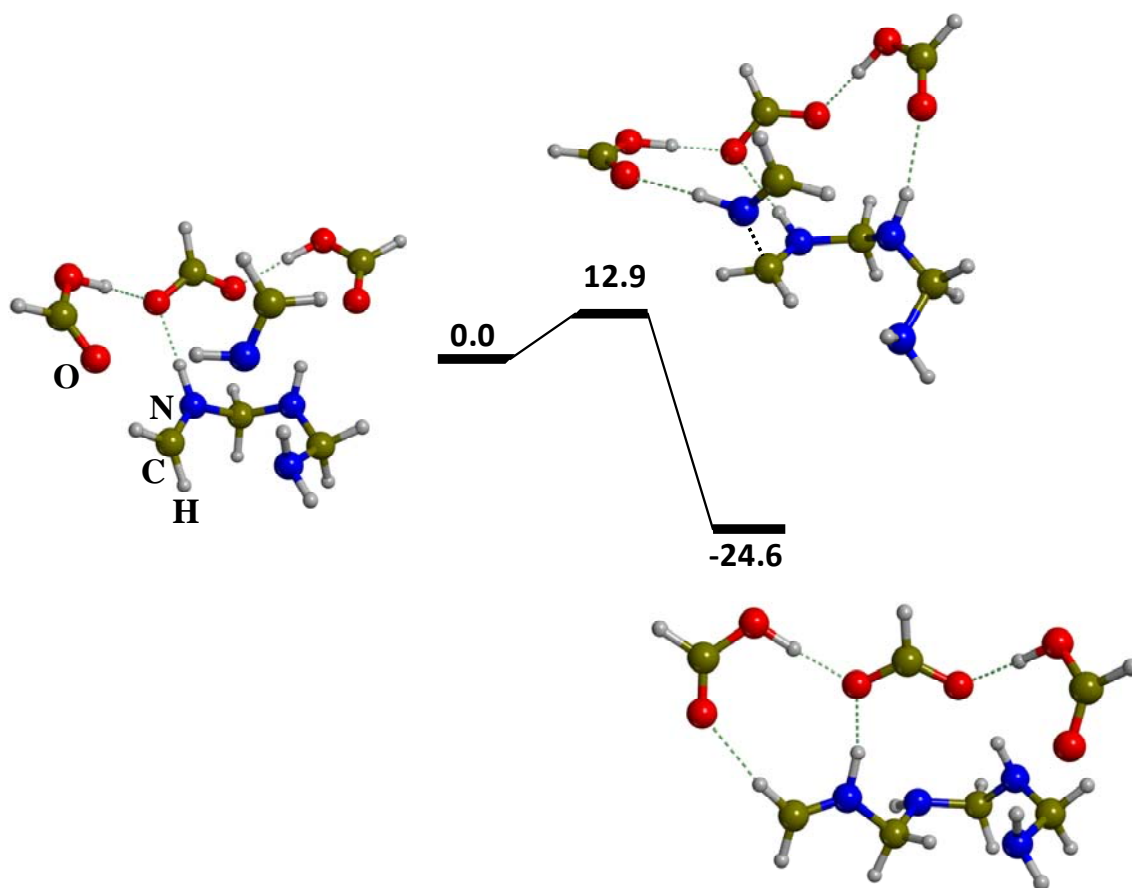
**Figure S3.** B3LYP potential energy profile for the gas-phase dehydration of aminomethanol to give methylenimine adopting the reaction of  $\text{NH}_2\text{CH}_2\text{OH} \rightarrow \text{CH}_2\text{NH} + \text{H}_2\text{O}$ . Units in  $\text{kJ mol}^{-1}$ .



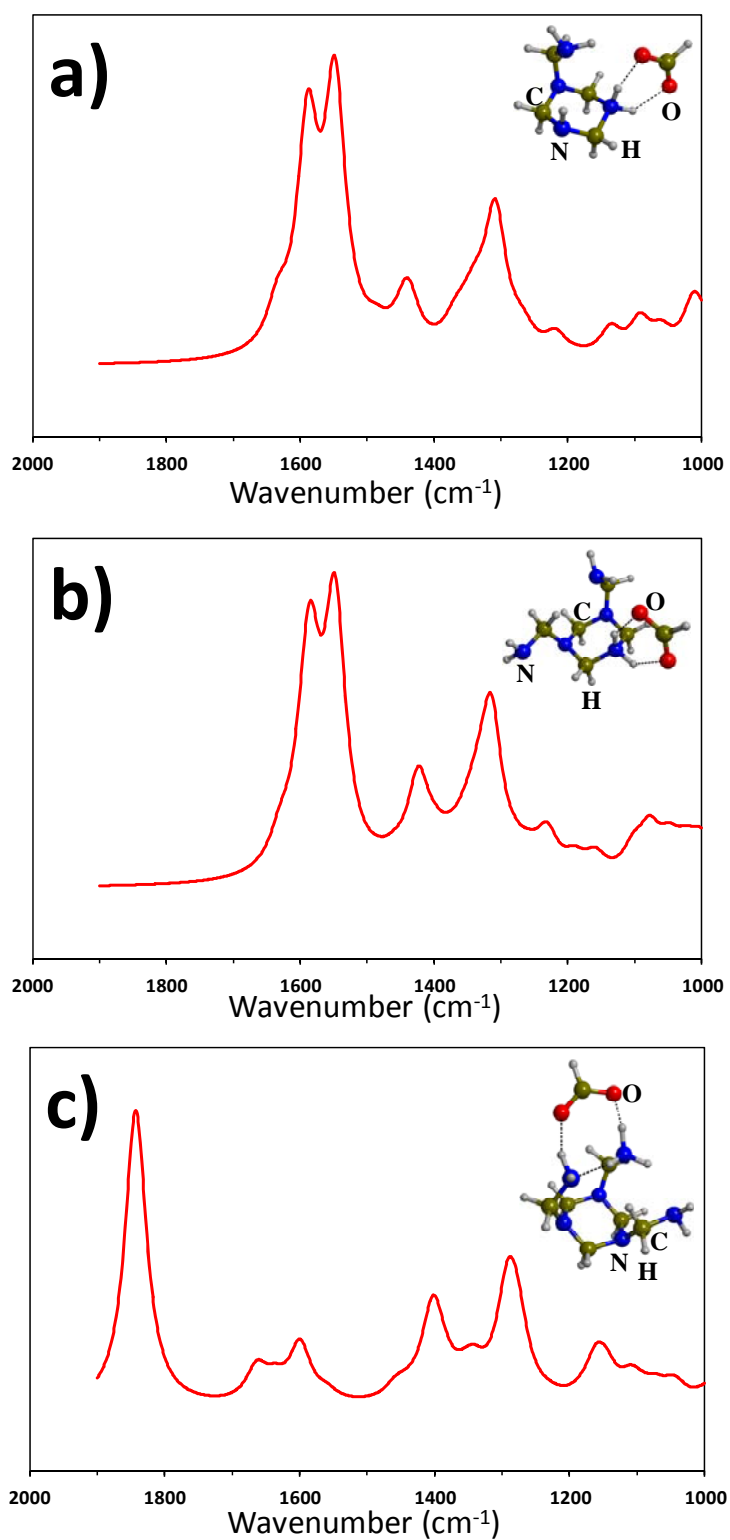
**Figure S4.** Infra-Red spectra of a HMT:HCOOH ice mixture at 100 K (a), 240 K (b), 250 K (c), and comparison with the HCOO<sup>-</sup> XH<sup>+</sup> salt species at 250 K (d).



**Figure S5.** B3LYP-potential energy profiles for the reaction between two  $\text{CH}_2\text{NH}$  molecules to form  $\text{CH}_2\text{NCH}_2\text{NH}_2$  in gas phase (a) and in the presence of  $\text{HCOOH}$  (b). Formation of trimethylentriamine from condensation of three  $\text{CH}_2\text{NH}$  molecules (c). Units in  $\text{kJ mol}^{-1}$ .

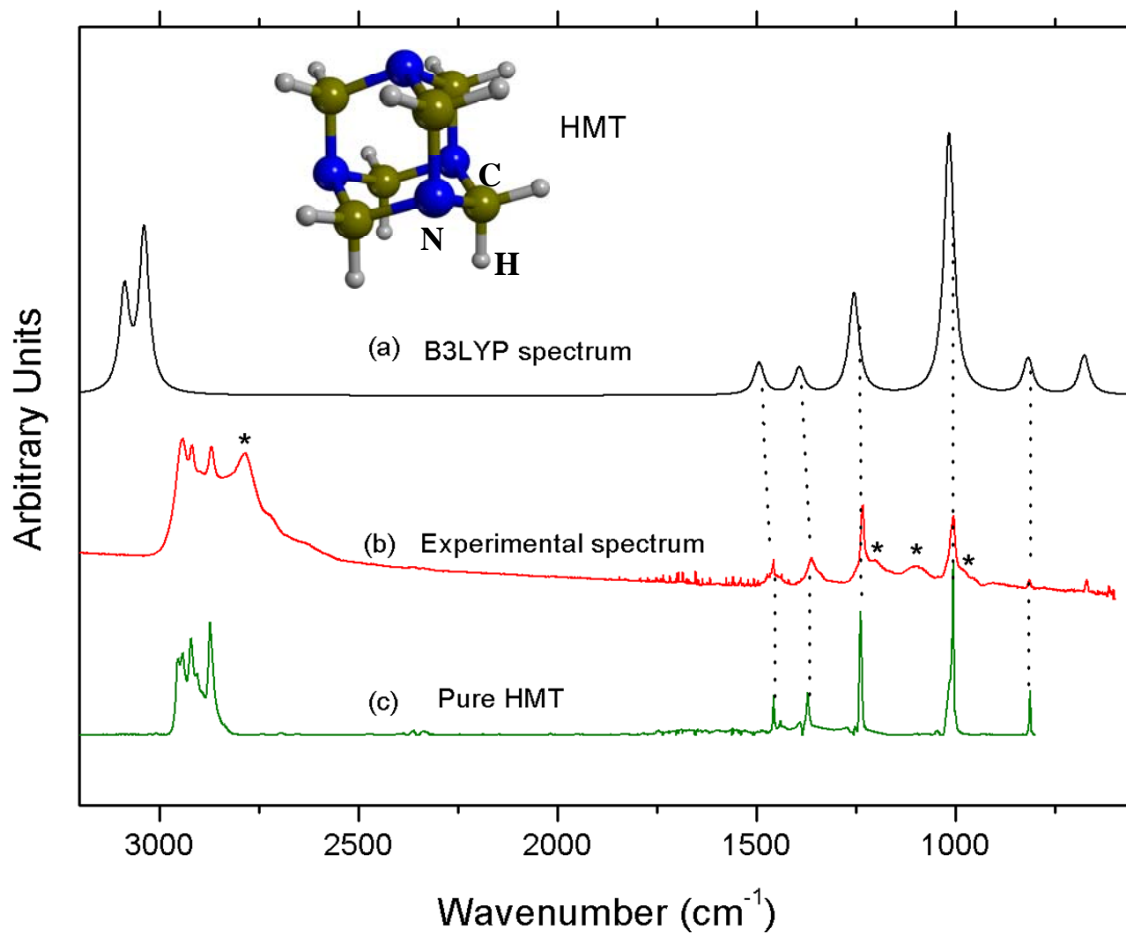


**Figure S6.** B3LYP-potential energy profile relative to the addition of  $\text{CH}_2\text{NH}$  on  $\text{NH}_2\text{CH}_2\text{NHCH}_2\text{NHCH}_2^+$ , as a further step for the elongation of polymethylenimine (PMI). Units in  $\text{kJ mol}^{-1}$ .



**Figure S7.** B3LYP-simulated IR spectra of possible candidates to be the  $\text{HCOO}^- \text{XH}^+$  salt observed experimentally: a), b) and c) refers to  $\text{TMTH}^+$  upon addition of one, two and three  $\text{CH}_2\text{NH}$  molecules, respectively.





**Figure S8.** Comparison between the IR spectra of HMT simulated at B3LYP (a), the experimental spectrum of HMT formed at 330 K from a CH<sub>2</sub>NH:HCOOH ice mixture (stars are assigned to PMI) (b), and pure HMT at 250 K (c).