Electronic Supplementary Information

Tubular perovskite solar cell: an improvement of charge separation at perovskite/HTM interfaces

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List of Contents

1. Experimental	S1
2. Morphologies of Tubular and Planar Perovskite Layers	S3
3. Nquist Plots of IMPS/IMVS Responses	S4
4. UV-vis Spectra of Perovskite Films	S4
5. Normalized Fluorescence (FL) Spectra of Perovskite Films	
References	S5

1. Experimental

Perovskite precursor preparation. Methyl ammonium iodide (CH₃NH₃I) was synthesized and purified according to literature.¹ The precursor in the fabrication of tubular perovskite was prepared by dissolving 0.1 mmol of CH₃NH₃I and 0.1 mmol of PbI₂ (99%, Aldrich) into 2.80 mL of DMF (99.8%, Aldrich). After stirred at 70 °C for 12 h under N₂ protection, CH₃NH₃PbI₃ solution with the concentration of 19 wt% was obtained.² Similar procedures were employed in the preparation of CH₃NH₃PbI₃ precursor forming planar perovskite layer. Instead of DMF, a mixture of γ -butyrolactone (GBL) and DMSO (7:3 v/v) was employed as the solvent. The final concentration of CH₃NH₃PbI₃ in the mixed solvent is 40 wt%.³

Solar cell fabrication. FTO-coated glass substrates (15 Ω /square, Nippon Sheet Glass) were cleaned with de-ionized water, acetone, and ethanol sequentially. A 30-50 nm dense blocking TiO₂ layer was fabricated by spin-coating an isopropoxide solution (350 µL Titanium (IV) isopropoxide (97 %, Alfa Aesar) in 5 mL of ethanol with 0.013 M HCl) onto precleaned FTO substrate at 2000 rpm for 30s and sintering the resultant substrate at 500 °C for 30 min.⁴ Afterwards, TiO₂ colloidal solution (20 nm TiO₂: terpineol: ethanol = 1: 1: 3, weight ratio) was spin-coated onto the blocking TiO₂ (bl-TiO₂) layer at 2000 rpm for 30 s. The resultant colloid-deposited FTO substrate was sintered at 500 °C for 30 min for attaining a mesoporous TiO₂ (mp-TiO₂) layer with the thickness of 200~300 nm. The mp-TiO₂ film was further treated with 50 mM TiCl₄ solution at 70 °C and subsequently annealed at 500 °C for 30 min.

The deposition of tubular perovskite layer was carried out by spin-coating the perovskite/DMF precursor onto the surface of mp-TiO₂ film at 3000 rpm for 60 s and heated at 100°C for 10 min under N₂ atmosphere. This protocol was also applied to form planar perovskite layer using the as-prepared perovskite/GBL/DMSO precursor. A slight difference is that the substrate was drop-casted with toluene before the step of CH₃NH₃PbI₃ solution casting. Afterwards, spiro-OMeTAD solution (72.3 mg of Spiro-OMeTAD, 17.5 μ L of lithium bis(trifluoromethanesulfonyl)-imide (LiTFSI)

solution (520 mg of LiTFSI in 1 mL of acetonitrile) and 28.8 μ L of 4-tertbutylpyridine all dissolved in 1 mL of chlorobenzene) was spin-coated onto the tubular and planar CH₃NH₃PI₃/mp-TiO₂/bl-TiO₂/FTO substrates at 2000 rpm for 30 s under N₂ atmosphere.⁵ Finally, spiro-OMeTAD deposited devices were covered with a thin graphite film with an active area of 0.09 cm⁻². This graphite film was fabricated by coating a graphite paste onto a precleaned FTO substrate with a doctor-blade technique and then annealed at 300°C for 30 min. The graphite paste was prepared by stirring a mixture of graphite nanosheets (20 nm, Chenyang Graphite, China), terpineol and ethanol (1:5:1, wt%) vigorously for 30 min at room temperature.

Device characterization. The morphologies of tubular and planar peroviskite films were recorded by using a Nova Nano SEM 230 field-emission scanning electron microscopy (FE-SEM, FEI). WAXS measurements were conducted by a D/max-2500 X-ray diffractometer (Rigaku, Japan) with Cu-K a radiation at a scanning speed of 1° /min. Photoluminesence spectra were determined by a HITACHI F-4500 fluorescence spectrophotometer. The UV-vis absorption spectra were recorded by using a SHIMADZU UV-2600 spectrophotometer. Current-voltage (J–V) characteristics were measured by a computer controlled Keithley-2400 digital source meter (Keithley, USA). Cell illumination was performed by a standard solar simulator with a 300 W xenon lamp and an AM 1.5 filter (Oriel 91160-1000, Newport, USA). Light intensity was adjusted to 1 sun (100 mW cm⁻²) by using a NREL-calibrated Si solar cell. The IPCE spectra were determined by a home-made IPCE measurement system, in which a DK240 monochromator (Spectral Products, USA) in combination with a 300-W halogen lamp was used to generate a light spectrum with a wavelength interval of 10 nm and a Keithley-2400 digital source meter was employed to record photocurrent responses. A Hamamatsu S1337-1010BQ silicon diode for determining incident light power was calibrated in National Institute of Metrology, China. Intensity modulated photovoltage and photocurrent spectroscopes (IMVS/IMPS) were measured by a Zennium electrochemical workstation (ZAHNER, Germany), in which a green light emitting diode (LED, 532 nm) was modulated to yield light frequency ranging from 100 mHz to 10 kHz and illumination intensity from 1 to 40 mW cm⁻². Electron densities were measured by charge extraction technique as reported.⁶ Open circuit voltage decay measurements were carried out by LK2005A Electrochemical System (Lanlike, China) for monitoring voltage signals. The MATLAB software was employed to convert the voltage decay data to lifetimes. All the measurements were performed in air at room temperature on the devices without sealing.

2. Morphologies of Tubular and Planar Perovskite Layers



Fig. S1 (a) SEM image of $CH_3NH_3PbI_3$ perovskite tubes coating onto mesoporous TiO_2 layer. Yellow hexagonal frame highlights the geometry of perovskite tube. (b) An enlargement of the green box in image a. (c) Top SEM view of planar $CH_3NH_3PbI_3$ layer.

3. Nyquist Plots of IMPS/IMVS Responses



Fig. S2 Typical Nyquist plots of (a) IMPS and (b) IMVS responses for tubular perovskite solar cells under illumination intensity of 10 mW cm⁻².

4. UV-vis Spectra of Perovskite Films



Fig. S3 UV-vis spectra of the devices with tubular and planar perovskite layers.

5. Normalized Fluorescence (FL) Spectra of Perovskite Films



Fig. S4 Normalized fluorescence (FL) spectra of tubular and planar perovskite films.

References

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