

## Electronic Supplementary Information

### **Abnormal spatial heterogeneity governing charge-carrier mechanism in efficient Ruddlesden–Popper perovskite solar cells**

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## Part 1. Fitting results in main text

### 1. The radiative recombination rate ( $k_{\text{rad}}$ ) and nonradiative recombination rates ( $k_{\text{nonrad}}$ )

The PL quantum yield (PLQY) is defined as the ratio of the radiative recombination rate ( $k_{\text{rad}}$ ) and the total recombination rate ( $k_{\text{tot}}$ )<sup>1</sup>,

$$\text{PLQY} = k_{\text{rad}} / k_{\text{tot}} = k_{\text{rad}} / (k_{\text{rad}} + k_{\text{nonrad}}), \quad (1)$$

where  $k_{\text{nonrad}}$  is the nonradiative recombination rate. All PLQY results are summarized in Table S3, while  $k_{\text{tot}}$  values were identified by the exponential fitting of transient PL spectra (TRPL) (Fig. S13).

### 2. The diffusion lengths ( $L$ ) and mobilities ( $\mu$ ) of charge carriers

We analyzed  $L$  and  $\mu$  using the 1D diffusion models and Einstein equations<sup>2</sup>,

$$L \approx 2d / \pi * (2 (\tau / \tau_{\text{quench}} - 1))^{1/2}, \quad (2)$$

$$\mu = L^2 q / (kT \tau) \quad (3)$$

where  $d$  is the film thickness,  $\tau$  is the carrier lifetimes of perovskites,  $\tau_{\text{quench}}$  is the quenched carrier lifetimes of perovskites in transporting layers,  $q$  is the elementary charge,  $k$  is Boltzmann constant, and  $T$  is temperature. All  $d$  values are shown in Fig. S15. Additionally,  $\tau$  and  $\tau_{\text{quench}}$  were estimated from the time-resolved photoluminescence (TRPL) of perovskite layers (Fig. S13), and perovskite layers on transporting layers (Fig. S16). The estimated  $d$  can be found in Fig. S15.

## Part 2. Supplementary tables

Table S1 Different elemental ratios of typical 3D-like (large  $n$ , here  $n = \infty$ ) and 2D-like (small  $n$ , here  $n = 3$ ) phases.

Element	Ratio (%) in 2D	Ratio (%) in 3D
Cs	0.145	0.255
C	9.675	1.645
I	50.6	54.0
Pb	28.2	33.3

Table S2 Statistical distribution of photovoltaic parameters for RPP solar cells using CsFMRP and MRP, respectively.

N	RPP	$V_{oc}$ (V)	$J_{sc}$ (mA cm <sup>-2</sup> )	FF (%)	PCE (%)
2	MRP	0.83±0.12	2.83±0.15	46.83±4.81	1.08±0.11
	CsFMRP	1.01±0.03	5.38±0.73	38.19±1.46	2.09±0.41
3	MRP	0.71±0.05	4.72±0.25	57.43±6.19	1.93±0.36
	CsFMRP	1.17±0.01	14.33±0.26	61.69±1.61	10.32±0.45
4	MRP	1.00±0.03	8.45±0.85	51.42±2.43	4.36±0.46
	CsFMRP	1.16±0.01	18.32±0.25	73.37±0.80	15.60±0.33

Table S3 PLQY of CsFMRP and MRP films, respectively.

PLQY (%)	CsFMRP	MRP
n = 2	0.706	0.17
n = 3	10.17	3.76
n = 4	23.6	8.37

Table S4 Dynamics of different ground state bleach (GSB) peaks of ultrafast transient absorption (TAS) spectra (Fig. 5c) for MRP and CsFMRP.

MRP		
GSB peak	First rise ( $\text{ps}^{-1}$ )	Second decay ( $\text{ps}^{-1}$ )
610 nm (n = 3)	/	0.039
645 nm (n = 4)	2.457	0.014
675 nm (n = 5)	2.231	0.022
CsFMRP		
GSB peak	First rise ( $\text{ps}^{-1}$ )	Second decay ( $\text{ps}^{-1}$ )
620 nm (n = 3)	/	0.031
750 nm (3D-like)	1.243	0.007

Table S5 Composition details of precursor solutions for MRP and CsFMRP, respectively.

MRP			
n	PbI <sub>2</sub> (mM)	MAI (mM)	PEAI (mM)
2	0.85	0.425	0.85
3	0.85	0.567	0.567
4	0.85	0.6375	0.425

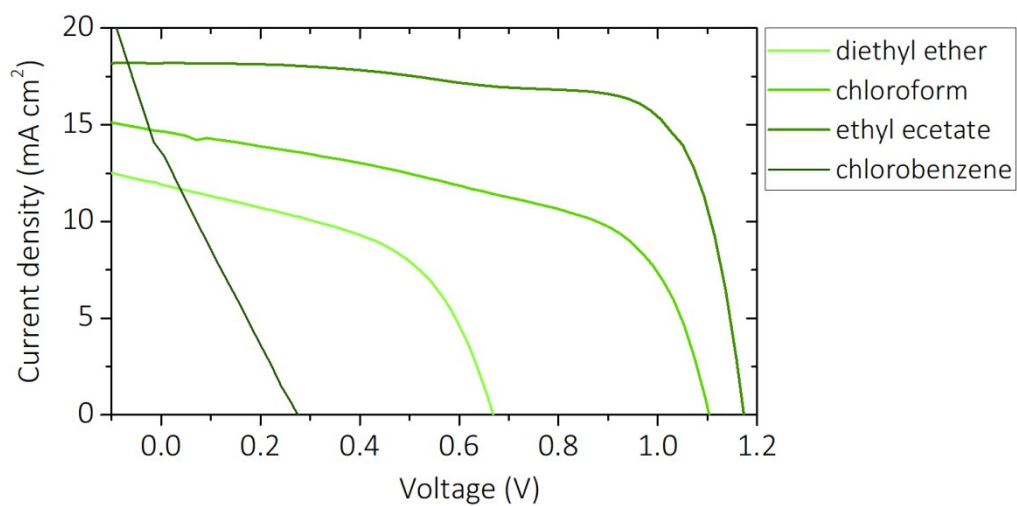
  

CsFMRP						
n	PbI <sub>2</sub> (mM)	PbBr <sub>2</sub> (mM)	FAI (mM)	MABr (mM)	PEAI (mM)	CsI ( $\mu$ L/1 mL) <sup>1</sup>
2	0.836	0.051	0.374	0.051	0.85	28.34
3	0.832	0.068	0.4987	0.068	0.567	37.76
4	0.83	0.0765	0.561	0.0765	0.425	42.5

<sup>1</sup> Volume of CsI stock solutions for 1 mL CsFMRP precursor solutions. The concentration of CsI stock solutions (in DMSO) is 1.5M.

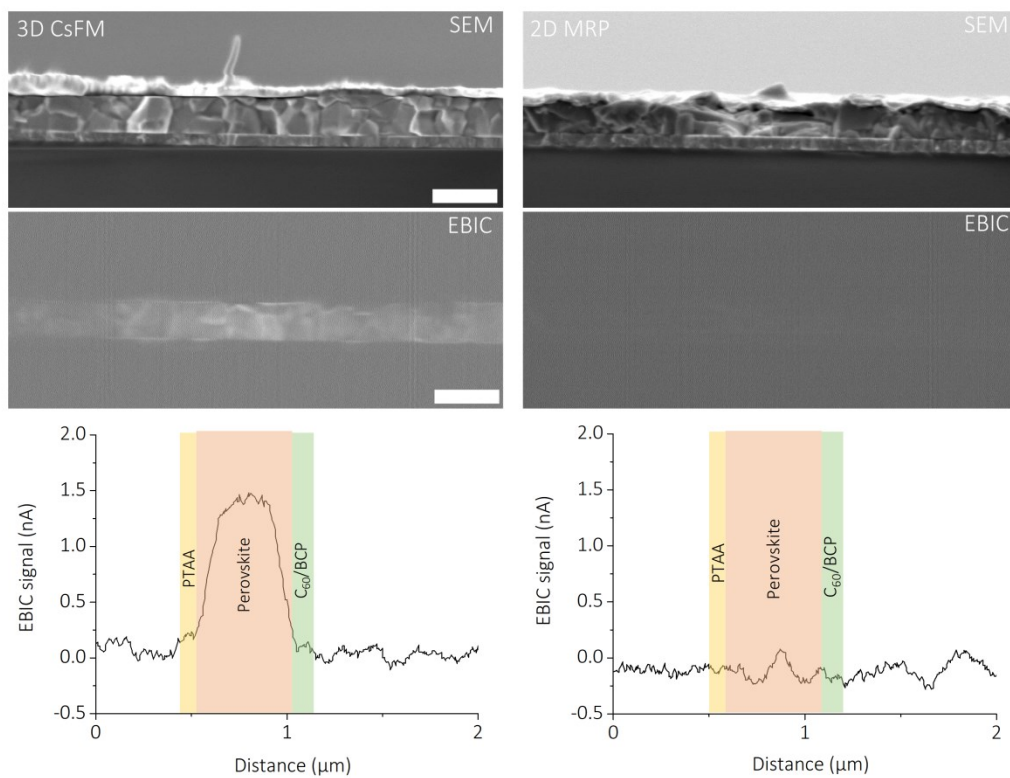


### Part 3. Supplementary figures

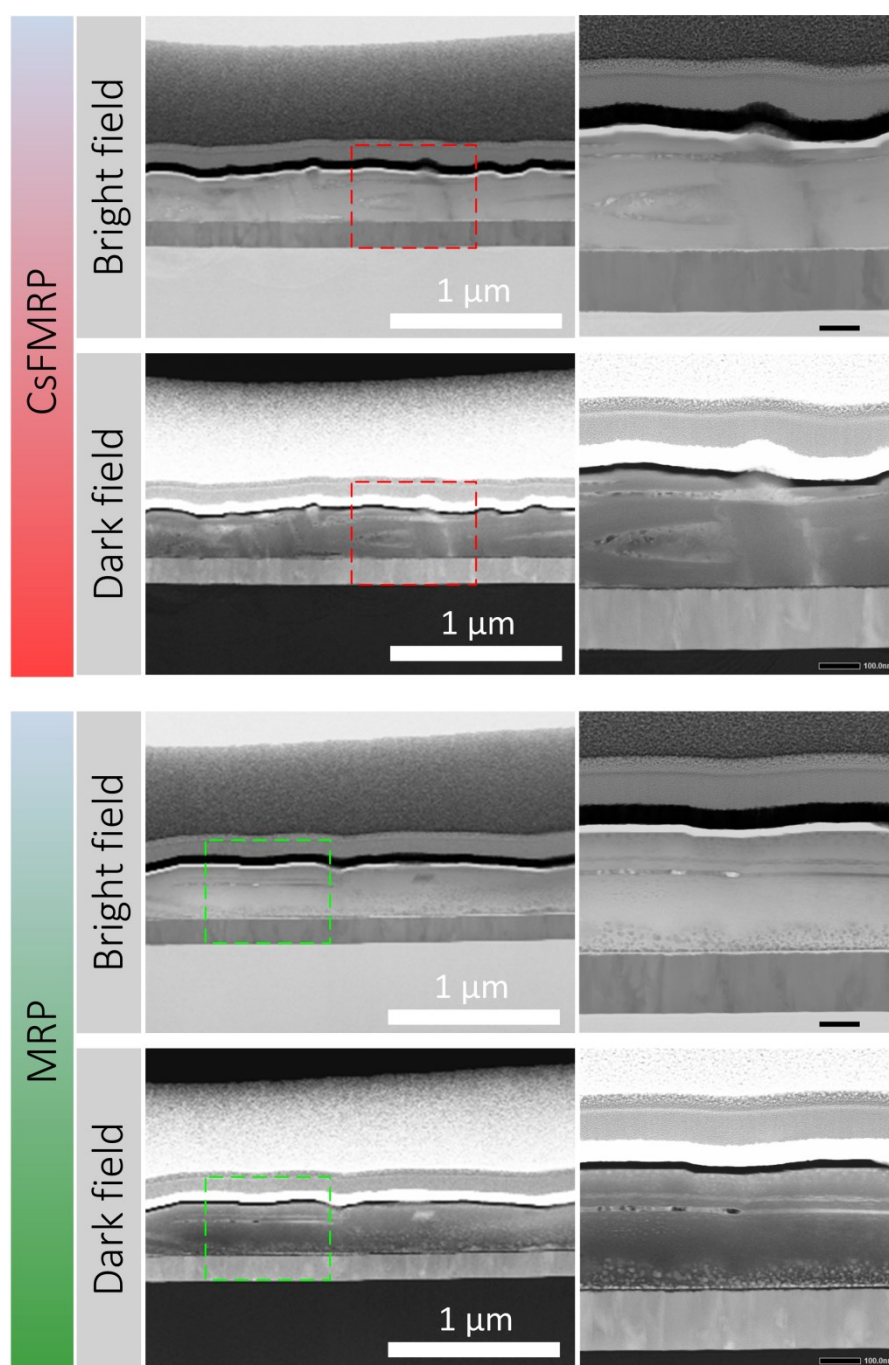


**Fig. S1** J-V curves of typical CsFMRP ( $n = 4$ ) perovskite solar cells driven by selective antisolvent.

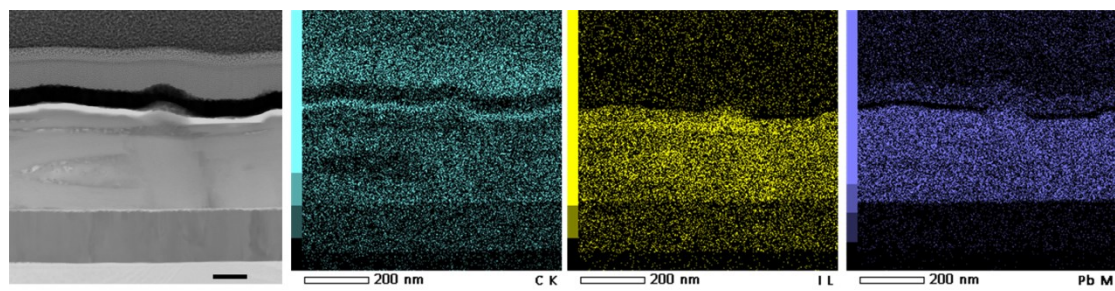
Traditional antisolvent chlorobenzene (CB) fails to work due to the bottom thin PTAA layer can partially dissolved by CB.



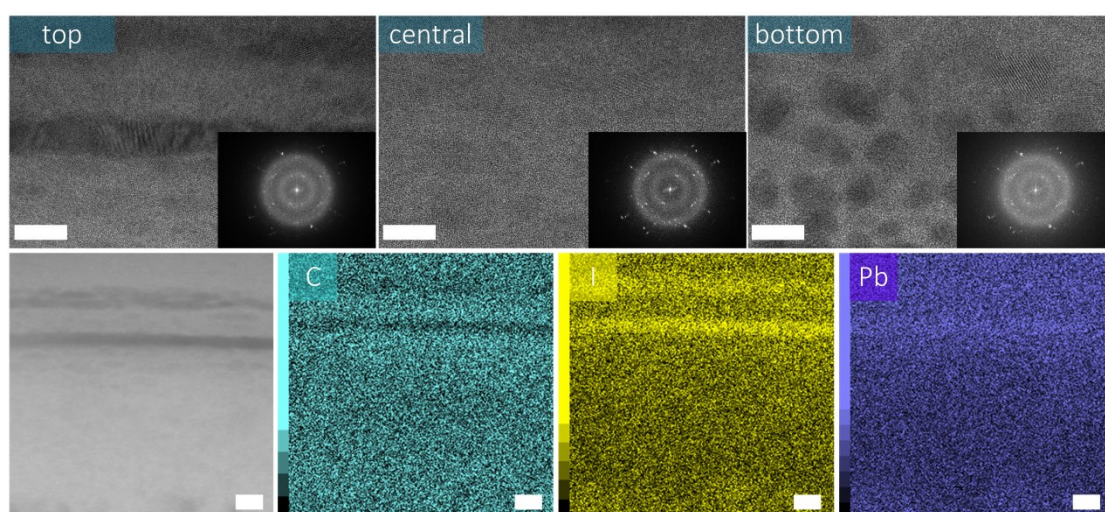
**Fig. S2** Cross sectional SEM and real-time EBIC real time images of 3D triple-cation perovskites (3D CsFM) and 2D MRP ( $n = 4$ ) based solar cells (scale bar: 1  $\mu\text{m}$ ). The bottom images show the depth profile of EBIC signals in corresponding devices.



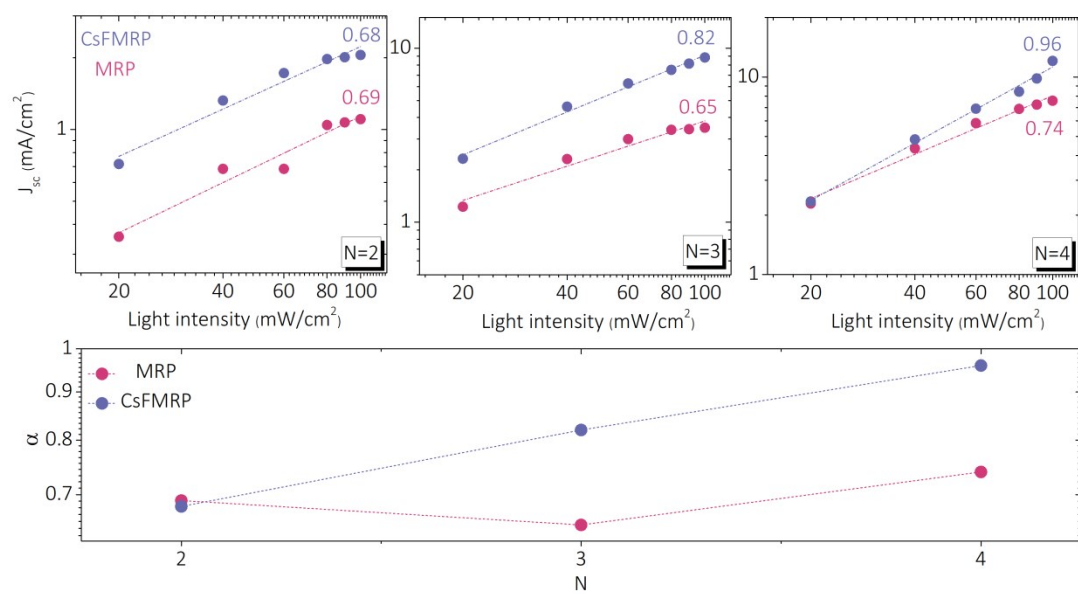
**Fig. S3** Cross sectional TEM images of CsFMRP and MRP based solar cells ((n = 4), scale bar: 100 nm).



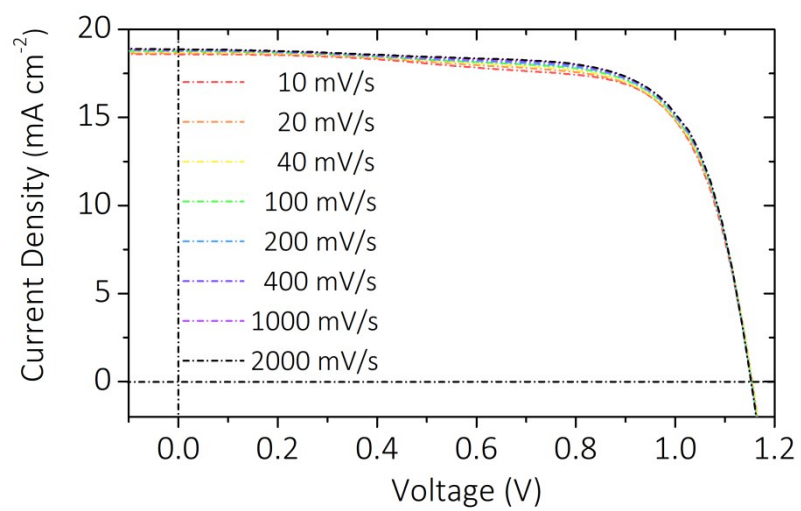
**Fig. S4** Low magnified TEM image and corresponding EDS elemental maps of CsFMRP ( $n = 4$ ).



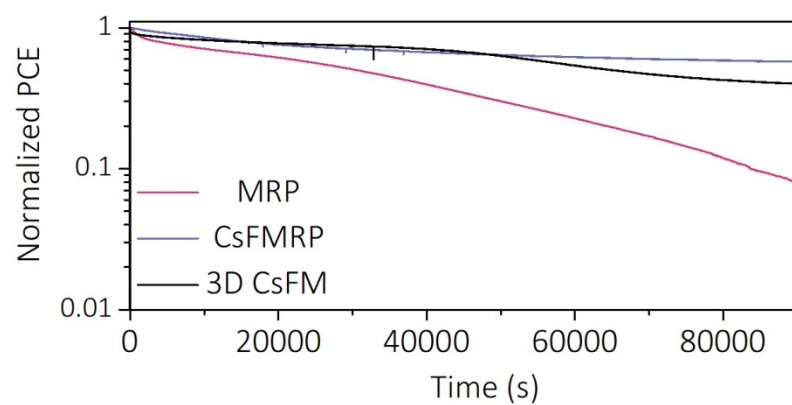
**Fig. S5** High-magnification FIB-HRTEM cross sectional images and corresponding EDS elemental mappings of MRP ( $n = 4$ ) perovskites based solar cells (scale bar: 20 nm).



**Fig. S6** Power law dependence factor  $\alpha$  estimated as the slope of linear fitting of  $J_{sc}$  versus  $I_{int}$  (light-intensity).

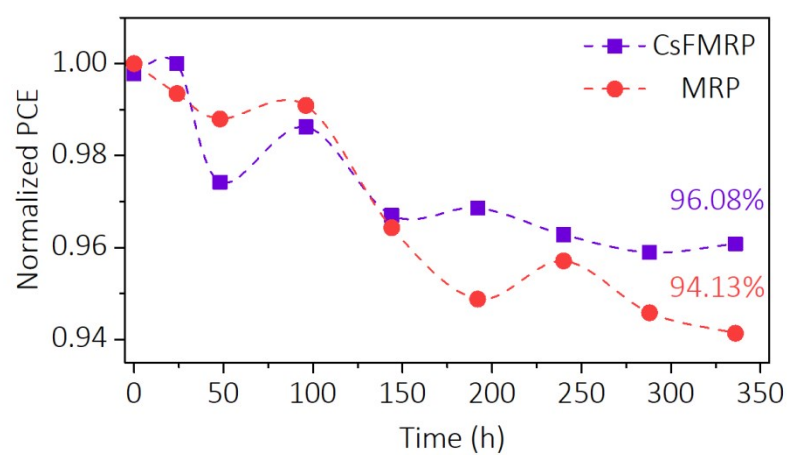


**Fig. S7** J-V curves depending on different scan rates of the champion CsFMRP (n = 4) solar cells.

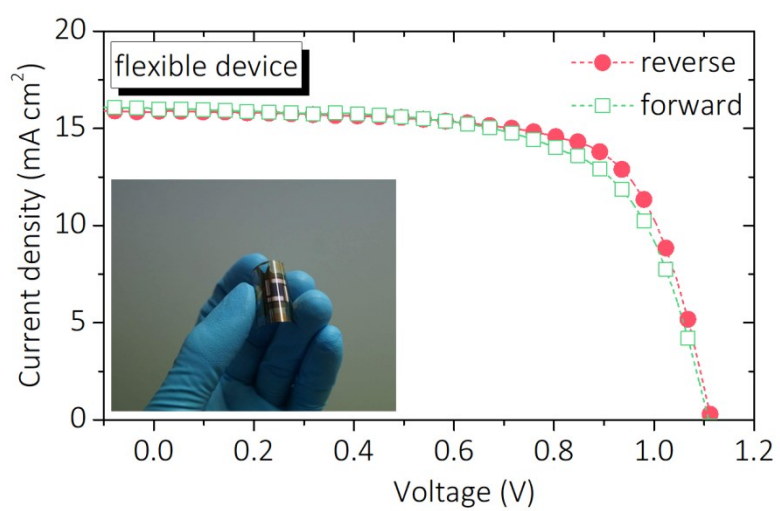


**Fig. S8** Light soaking stabilities of the best devices using 2D CsFMRP, MRP ( $n = 4$ ) and 3D CsFM, respectively.



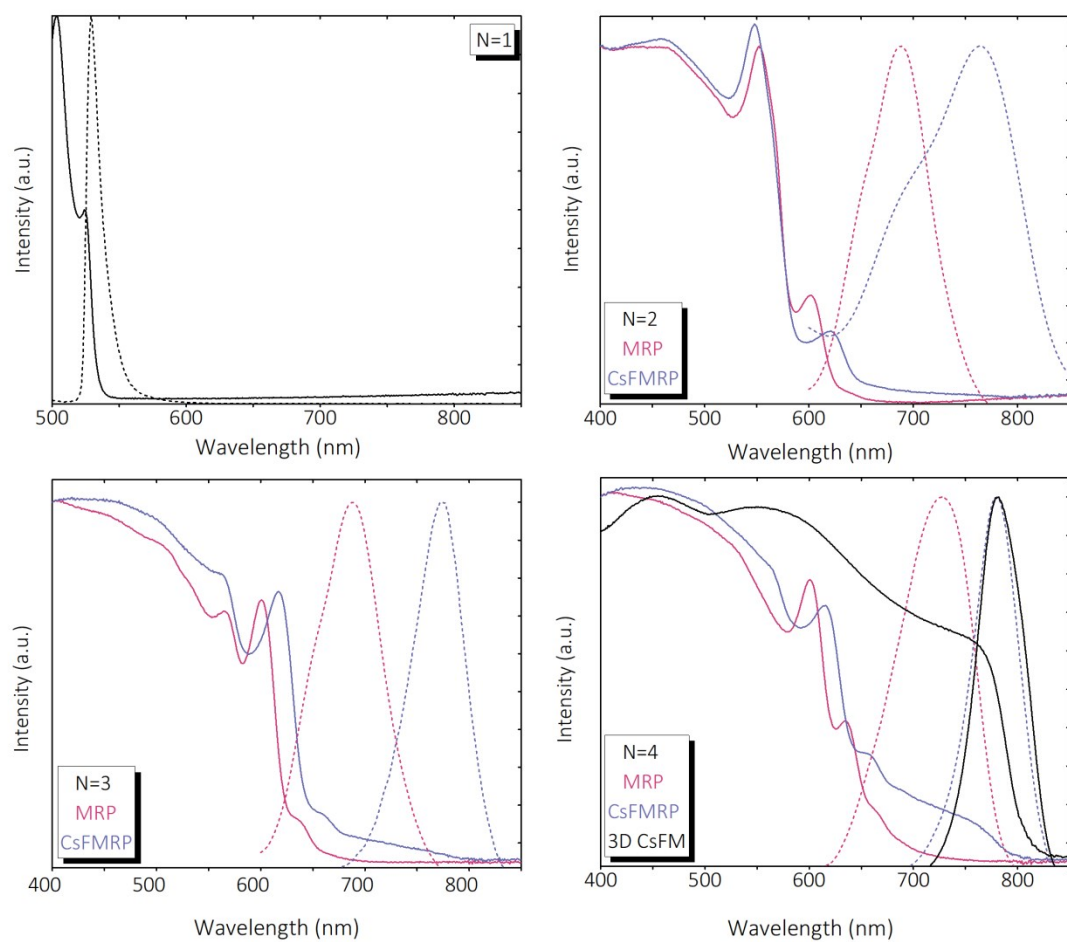


**Fig. S9** Moisture stability of the non-encapsulated devices using CsFMRP and MRP stored in ambient air (~50% humidity).

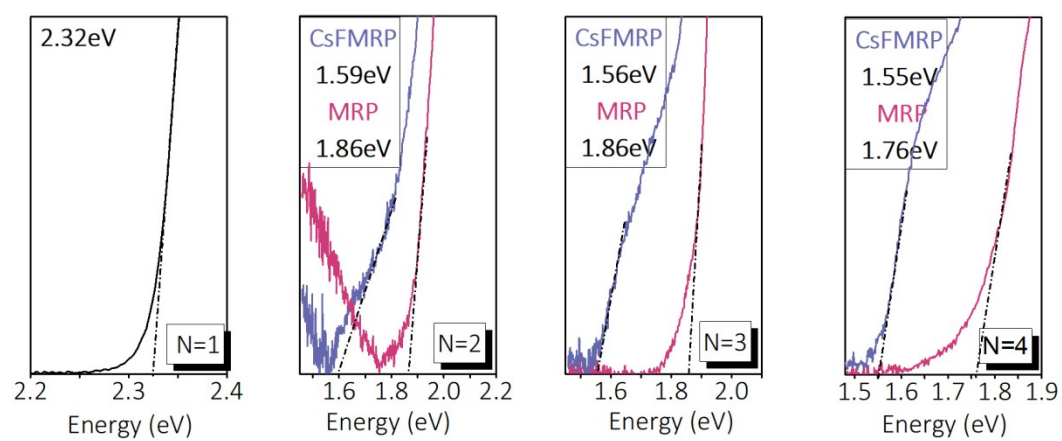


scan	$V_{oc}$ (V)	$J_{sc}$ ( $\text{mA/cm}^2$ )	FF (%)	PCE (%)
reverse	1.11	14.8	69.7	11.5
forward	1.11	15.0	65.0	10.8

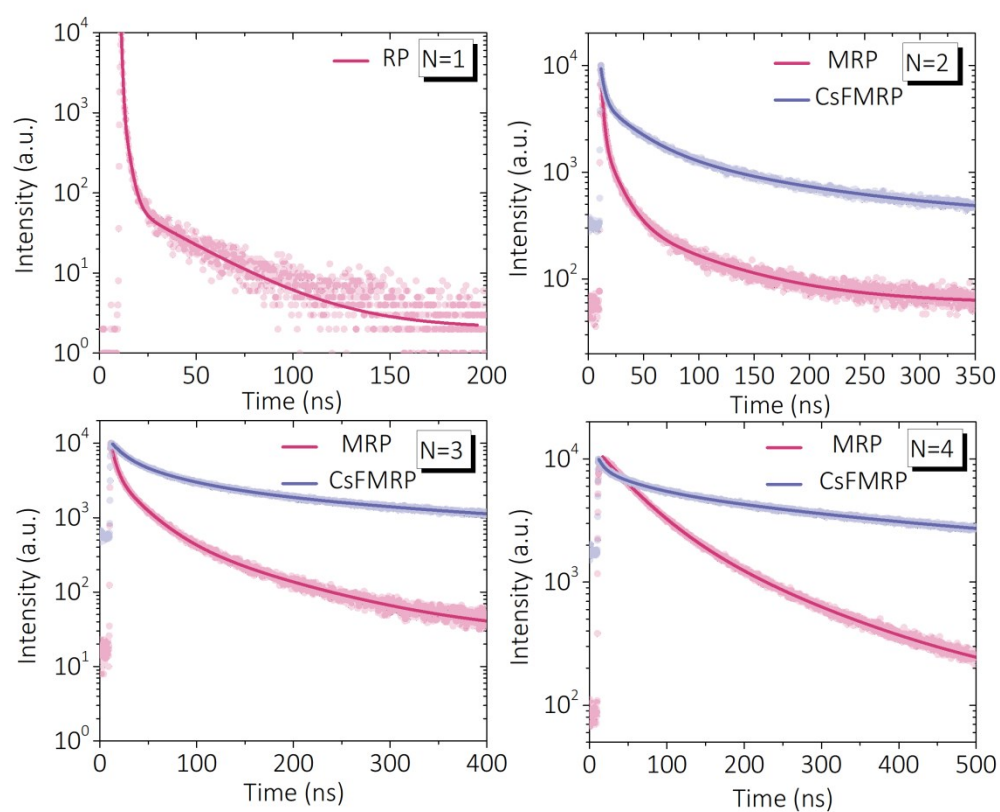
**Fig. S10** Champion flexible 2D CsFMRP ( $n = 4$ ) solar cell. Little hysteresis can be found between reverse and forward scan.



**Fig. S11** UV-Vis absorption and steady-state PL spectra of different RPP films.

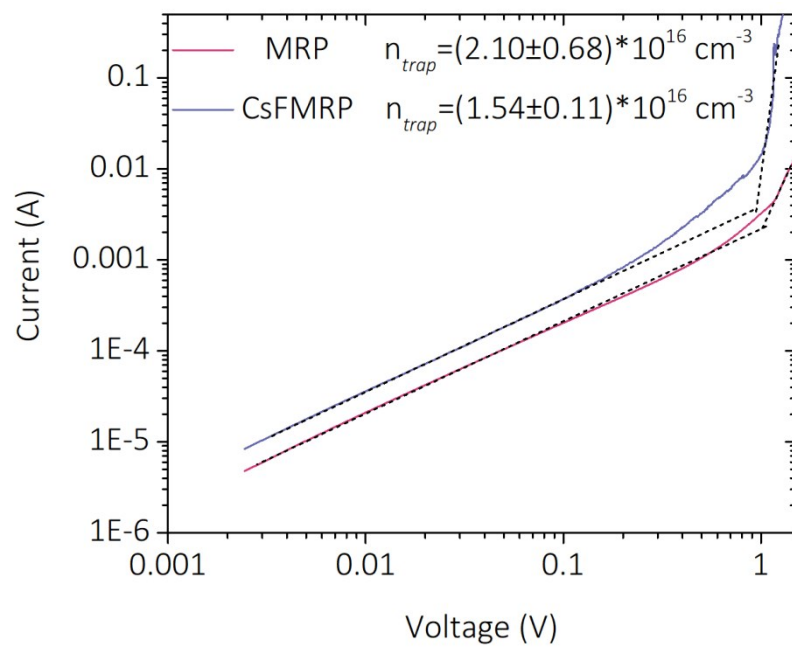


**Fig. S12** Bandgap ( $E_g$ ) fitting from UV-Vis absorption spectra.

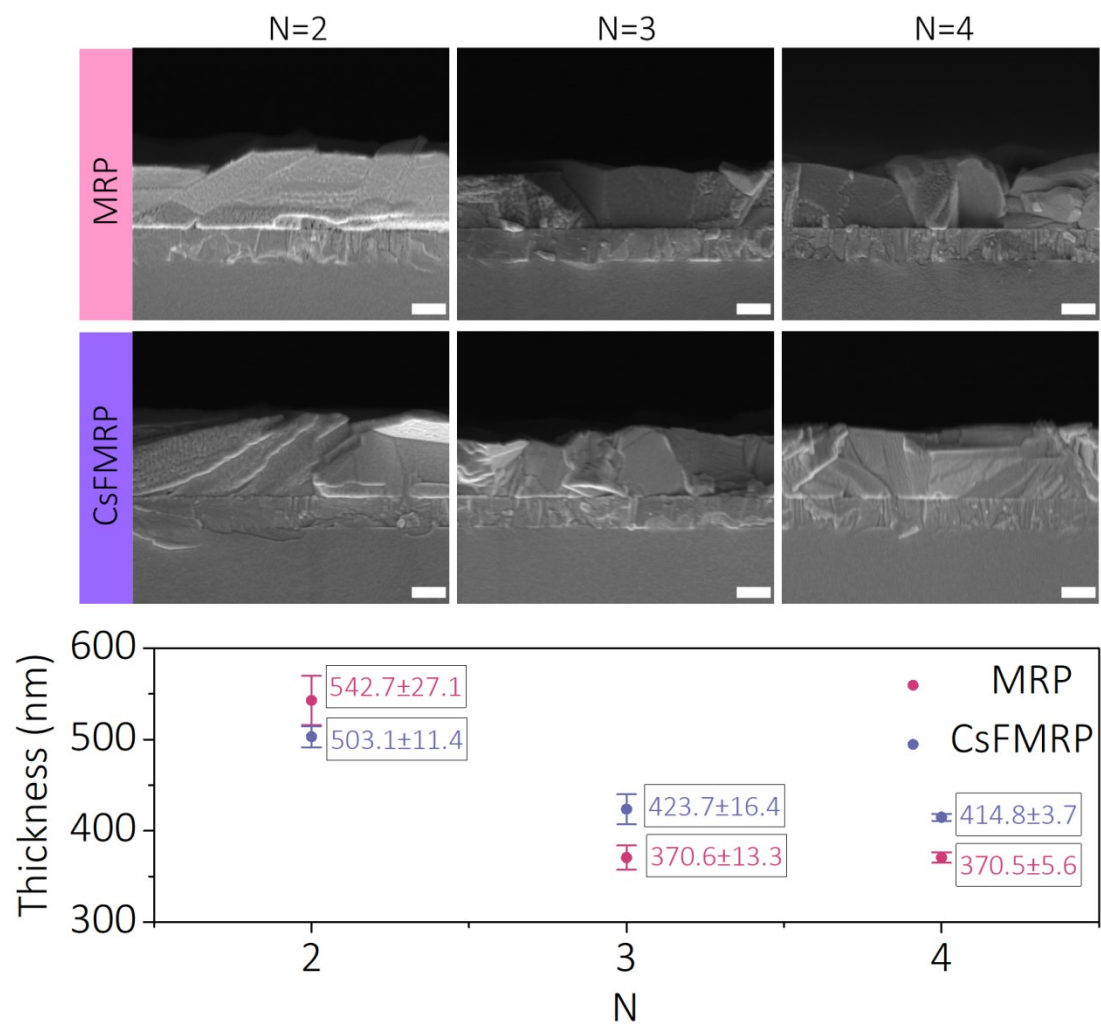


**Fig. S13** Transient PL (TRPL) spectra of different RPP films on quartz. Carrier lifetime ( $\tau$ ) by exponential fitting of corresponding film is summarized as follows,

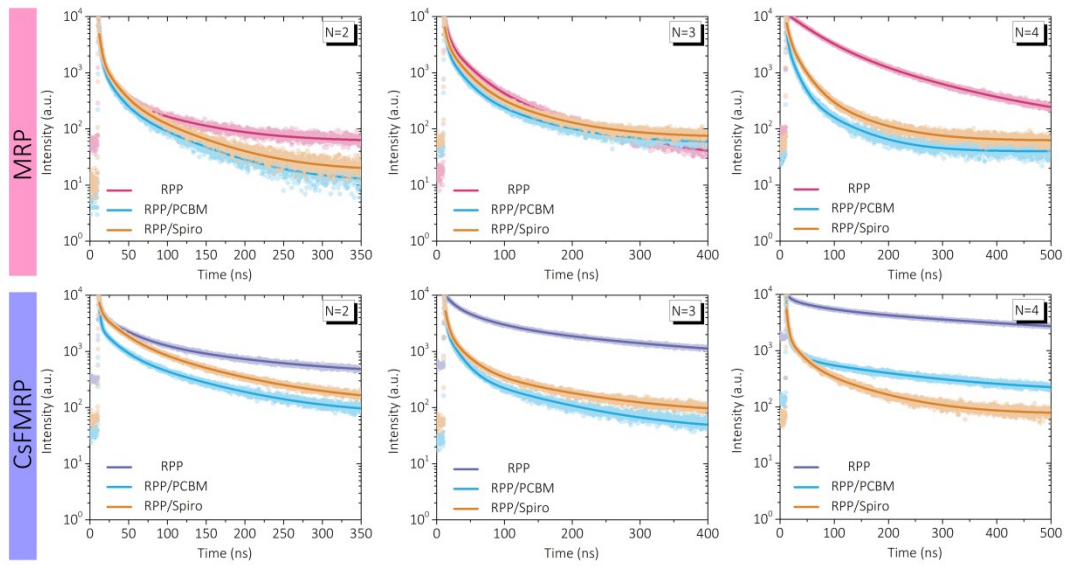
$\tau$ (ns)	CsFMRP	MRP
n = 1	1.06	
n = 2	38.60	10.29
n = 3	98.63	50.38
n = 4	229.40	101.04



**Fig. S14** Space charge limited charge (SCLC) methods to evaluate the densities of trap states.



**Fig. S15** Film thicknesses of different RPP films (scale bar: 200 nm).



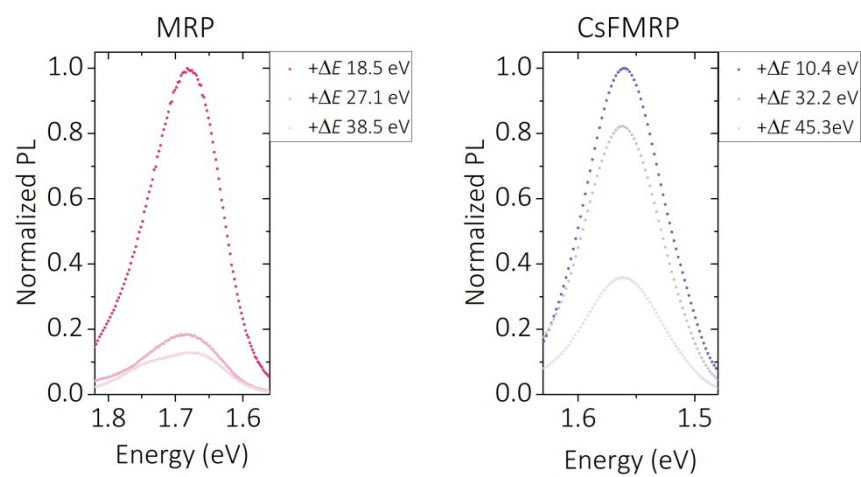
**Fig. S16** TRPL spectra of different perovskites themselves and perovskites on electrons (PCBM) /holes (Spiro-MeOTAD) transporting layers. Quenched carrier lifetime ( $\tau_{\text{quench}}$ ) by exponential fitting of corresponding film on PCBM/Spiro-MeOTAD is summarized as follows,

$\tau_{\text{quench}}$ on PCBM (ns)	CsFMRP	MRP
n = 1	1.03	
n = 2	22.12	9.23
n = 3	10.70	15.81
n = 4	16.52	13.86

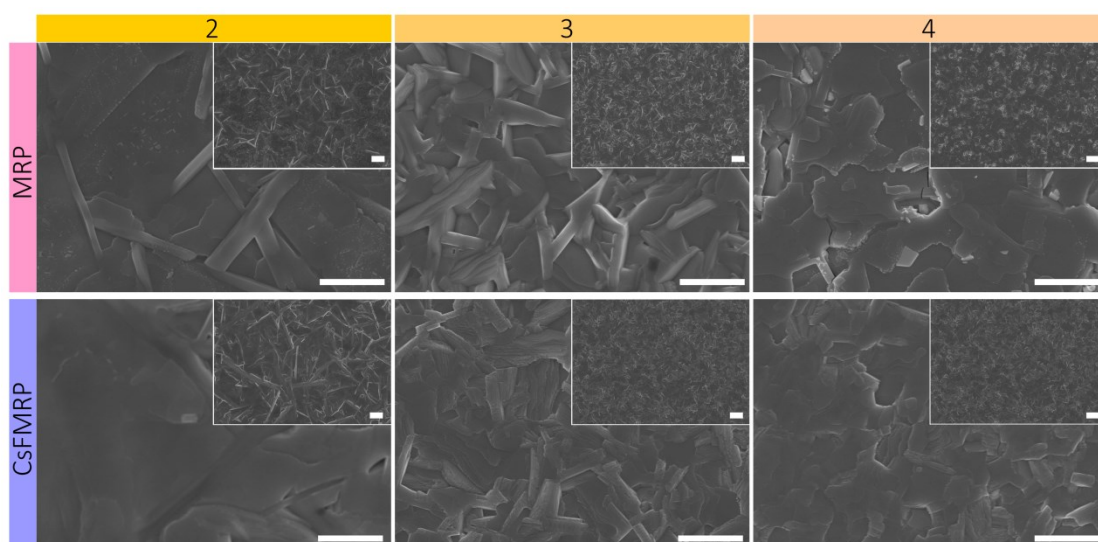
  

$\tau_{\text{quench}}$ on Spiro (ns)	CsFMRP	MRP
n = 1	1.01	
n = 2	30.8	9.56
n = 3	15.3	18.89
n = 4	19.13	19.16

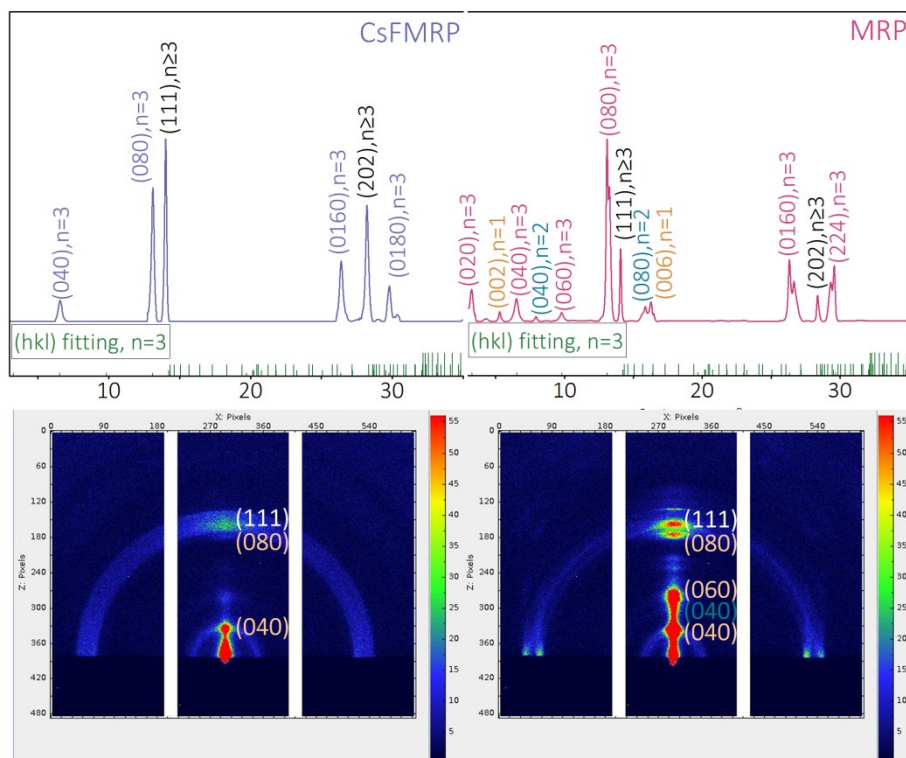




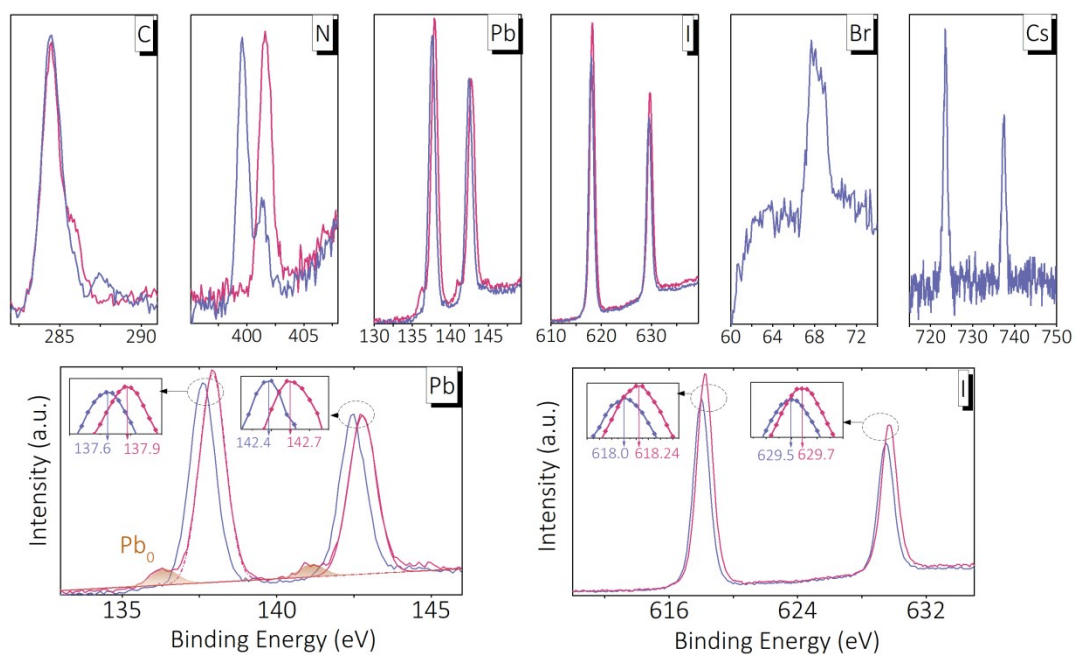
**Fig. S17** Steady state PL spectra of different RPP ( $n = 4$ ) using excess excitation energy beyond electronic band edges.



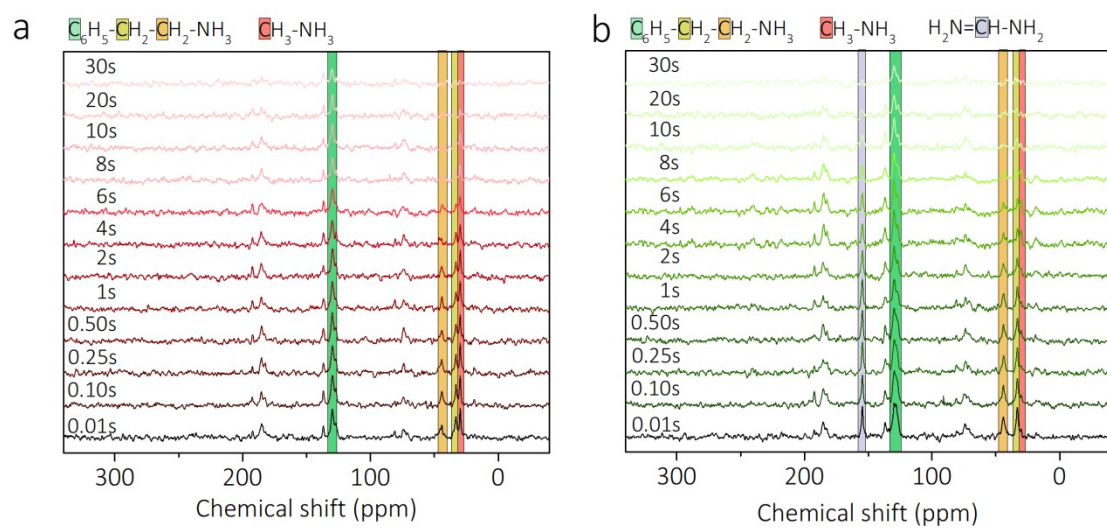
**Fig. S18** Top-view SEM images of corresponding RPP films (scale bar: 1  $\mu\text{m}$  in high magnification, 2  $\mu\text{m}$  in low magnification).



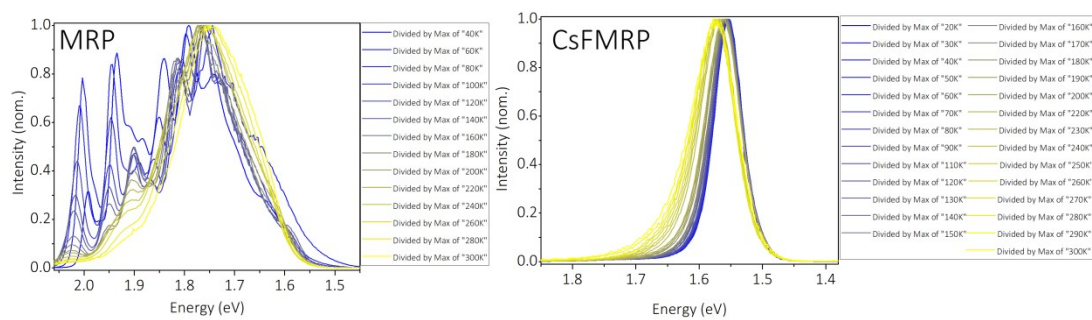
**Fig. S19** XRD and related grazing-incidence wide-angle X-ray scattering (GIWAXS) images of corresponding RPP ( $n = 4$ ) films. In CsFMRP, crystal phase is totally composed of ordered  $n = 3$  phase and 3D-like (large- $n$ ) phase. In contrast, in MRP, crystal phase seems more complicated, composed of  $n = 1, 2, 3$  phases and 3D-like (large- $n$ ) phase.



**Fig. S20** XPS spectra of corresponding RPP films ( $n = 4$ ), and detailed  $\text{PbI}_3^-$  cage interaction (red is MRP, blue is CsFMRP).



**Fig. S21** Details of  $^{13}\text{C}$  resonances in SSNMR dependent on relaxation time for (a) MRP and (b) CsFMRP.



**Fig. S22** Detailed normalized TPL spectra of corresponding RPP ( $n = 4$ ) films.

## References

1. Gong, X. et al. (2018). Electron-phonon interaction in efficient perovskite blue emitters. *Nat. Mater.* *17*, 550–556.
2. Proppe, A. H. et al. (2018). Synthetic control over quantum well width distribution and carrier migration in low-dimensional perovskite photovoltaics. *J. Am. Chem. Soc.* *140*, 2890–2896.