

Supporting information

Identifying high performance and durable methylammonium-free lead halide perovskites through high throughput synthesis and characterization

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22 I. Experimental materials and methods

23

24 **Materials**

25 Lead iodide (PbI_2) and lead bromide (PbBr_2) were purchased from TCI America. Formamidinium
26 iodide (FAI) and formamidinium bromide (FABr) were purchased from Dyenamo. 2,2',7,7'-
27 Tetrakis[N,N-di(4-methoxyphenyl)amino]-9,9'-spirobifluorene (spiroOMeTAD) was purchased
28 from 1-Material. Dimethylformamide (DMF), dimethyl sulfoxide (DMSO) and chlorobenzene
29 (CB) were purchased from Acros Organics. Cesium iodide (CsI), cesium bromide (CsBr), titanium
30 diisopropoxide bis(acetylacetonate) (75 wt. % in isopropanol), acetylacetone, anhydrous ethanol
31 (EtOH), titanium dioxide paste (TiO_2 , 30 NR-D), lithium bistrifluorosulfonyl imide (Li-TFSI),
32 tris(2-(1H-pyrazol-1-yl)-4-tert-butylpyridine)cobalt(III) tri[bis(trifluoromethane)sulfonimide]
33 (FK 209 Co(III) TFSI), 4-tertbutyl pyridine (tBP), acetonitrile were bought from Sigma-Aldrich.
34 Au pellets were purchased from Kurt J. Lesker. All the chemicals were used as received without
35 further purification. Conductive glass, fluorine-tin-oxide (FTO) coated (10 ohms/sq, 1-inch \times 1-
36 inch), was purchased from Kintec.

37

38 **Perovskite active layer deposition**

39 Preparation of 0.4 M perovskite precursor solutions: Due to the poor solubility of CsBr in DMF
40 and DMSO, 0.4 M perovskite precursor solutions were prepared for all 49 compositions. Four
41 master solutions were used to prepare a total of 49 precursor solutions with varying ratios of Cs/FA
42 and Br/I. The master solutions were 0.4 M CsPbI_3 , FAPbI_3 , CsPbBr_3 , and FAPbBr_3 , which were
43 synthesized by dissolving stoichiometric CsI and PbI_2 , FAI and PbI_2 , CsBr and PbBr_2 , FABr and
44 PbBr_2 in DMSO, respectively. The perovskite precursor solutions were prepared in a N_2 glovebox
45 with < 2 ppm of O_2 and H_2O under the mild heating condition at $\sim 65^\circ\text{C}$ for 1 h to assist dissolving.
46 These four solutions were mixed in the right molar ratios to yield the other 45 stoichiometric
47 solutions with the desired $\text{Cs}_y\text{FA}_{1-y}\text{Pb}(\text{Br}_x\text{I}_{1-x})_3$ compositions.

48 Fabrication of perovskite thin-films from 0.4 M perovskite precursor solutions: The 0.4 M
49 precursor perovskite solution was freshly spin-coated on a substrate pre-heated at 65°C by a two-
50 step spin-coating process, dynamically adding CB as the antisolvent. The first step was carried out
51 at 500 rpm with an acceleration rate of 250 rpm/s for 10 s. The second step followed at 2000 rpm
52 with an acceleration rate of 1000 rpm/s for 70 s. CB (250 μL) was gently dripped at 5 s before the

53 second step end. All films were then annealed at 65°C for 10 min with the exception of CsPbI₃
54 (65°C for 1 min).

55 Preparation of 1.0 M perovskite precursor solutions: Four compositions (i.e. Cs_{2/6}FA_{4/6}PbI₃,
56 Cs_{2/6}FA_{4/6}PbBr_{5/2}I_{1/2}, Cs_{1/6}FA_{5/6}PbI₃ and Cs_{1/6}FA_{5/6}PbBr_{5/2}I_{1/2}) were selected to increase the
57 concentration from 0.4 M to 1.0 M to improve the photovoltaic performance. 1 M perovskite
58 precursor solutions of Cs_{2/6}FA_{4/6}PbI₃ and Cs_{1/6}FA_{5/6}PbI₃ were prepared by dissolving a mixture of
59 PbI₂, FAI, and CsI with 5% excess stoichiometry of PbI₂ in a mixed solvent of DMF and DMSO
60 [DMF (v) : DMSO (v) = 4:1]. 1 M perovskite precursor solutions of Cs_{2/6}FA_{4/6}PbBr_{5/2}I_{1/2} and
61 Cs_{1/6}FA_{5/6}PbBr_{5/2}I_{1/2} were prepared by dissolving a mixture of PbI₂, PbBr₂, FAI, and CsI with 5%
62 excess stoichiometry of PbI₂ and PbBr₂ in a mixed solvent of DMF and DMSO [DMF (v) : DMSO
63 (v) = 4:1]. The perovskite precursor solutions were prepared in a N₂ glovebox with < 2 ppm of O₂
64 and H₂O under the mild heating condition at ~70 °C for 1 h to assist dissolving.

65 Fabrication of perovskite thin-films from 1.0 M perovskite precursor solutions: The 1.0 M
66 precursor perovskite solution was spin-coated on the substrate. Spin-coating parameters are two
67 steps, 1000rpm 10 seconds (1000 rpm/s) followed by 6000rpm 20 seconds (2000 rpm/s). CB was
68 gently added 5 s before the end of the second step. The films were then annealed at 150 °C for 10
69 min.

70

71 **Device fabrication**

72 Etched FTO-coated glass was sequentially cleaned by sonication in a detergent solution (2%
73 Hellmanex in deionized water), deionized water, acetone, and isopropanol for 15 min to each
74 solvent. After drying with N₂, the substrates were treated under ultraviolet-ozone for 15 min to
75 remove the last traces of organic residues before the deposition of the electron-transporting layer.
76 A compact TiO₂ (c-TiO₂) layer was deposited on the top of FTO-coated glass by spray pyrolysis
77 of a precursor solution of 480 μL acetylacetone and 720 μL titanium diisopropoxide
78 bis(acetylacetonate) (75 wt. % in isopropanol) diluted in 10.8 mL of anhydrous ethanol. The
79 solution was sprayed on the substrates preheated at 450 °C using O₂ as the carrier gas with a 30 s
80 interval between each cycle. The FTO/c-TiO₂ substrates were kept at 450 °C for 30 minutes after
81 finishing the spray procedure then left cool down to ambient temperature. Next, a mesoporous
82 TiO₂ (m-TiO₂) layer was coated on the top of the FTO/c-TiO₂ substrate by spin-coating a 150
83 mg/mL diluted TiO₂ paste (30 NR-D) in anhydrous ethanol for 10 s at 4000 rpm with a ramp rate

84 of 4000 rpm s⁻¹. After drying at 100 °C, the as-prepared FTO/c-TiO₂/m-TiO₂ stack was sintered at
85 450 °C for 30 min. The perovskite layer was deposited on the substrate right after completing the
86 annealing process. A perovskite active layer was deposited by spin-coating a perovskite precursor
87 solution in a N₂ glovebox as mentioned in the “Perovskite active layer deposition” section. The
88 hole-transporting layer was deposited by spin-coating a solution consisting of 91.4 mg of Spiro-
89 OMeTAD, 36.1 μL of tBP, 20.7 μL Li-TFSI (1.8 M in acetonitrile), 9.0 μL of Co(III) TFSI solution
90 (0.25 M in acetonitrile), and 1 mL CB. The concentration of Spiro-OMeTAD solution was 70 mM
91 and the molar ratio of Spiro-OMeTAD : Co(III) TFSI : Li-TFSI : tBP was 1 : 0.03 : 0.5 : 3.3. The
92 80 μL doped Spiro-OMeTAD solution was dynamically spin-coated on the surface of the
93 perovskite layer at 3000 rpm for 30 s with the ramp of 3000 rpm/s. Finally, 80 nm of Au electrode
94 was thermally evaporated on the top of the Spiro-OMeTAD layer. The evaporation rate was
95 adjusted to 0.01 Å s⁻¹ at the first 1 nm, 0.05 Å s⁻¹ for the thickness between 1 nm and 5 nm, 0.1 nm
96 s⁻¹ for the thickness between 5 nm and 20 nm, and 0.5 Å s⁻¹ for the rest 60 nm.

97

98 **Characterization**

99 UV-vis absorption measurements were performed on a Varian Cary 5000 UV-Vis/NIR
100 spectrometer with a dual-beam. Steady-state photoluminescence was measured using a Horiba
101 Jobin Yvon Fluorolog 3-2iHR1 fluorometer with a 450 W xenon lamp as the excitation source.
102 The films were oriented at 60° relative to the excitation source to decrease interference from
103 reflected light. Perovskite films for UV-vis absorption and photoluminescence measurements were
104 encapsulated with epoxy resin to prevent degradation due to the exposure to ambient atmosphere.
105 The surface morphology of perovskite films was characterized by a high-resolution field emission
106 scanning electron microscope (SEM, Hitachi SU8010) operating at 5 keV and 10 μA. X-ray
107 diffraction (XRD) patterns were recorded in ambient using Malvern PANalytical Empyrean with
108 Cu_{Kα} radiation (1.54 Å wavelength). The XRD patterns were collected in Bragg-Brentano
109 geometry from 2θ = 5-65°. Grazing incidence wide-angle X-ray scattering (GIWAXS)
110 measurements were performed at beamline 11-BM of the National Synchrotron Light Source II at
111 Brookhaven national laboratory. The X-ray beam had an energy of 13.5 keV, 0.2 mm (height) x
112 0.05 mm (width) size, 1 mrad divergence and an energy resolution of 0.7%. The perovskite films
113 deposited on glass with a size of about 0.5 cm × 0.5 cm were measured at an incidence angle of

0.5° with respect to the substrate plane and an exposure time of 10 s. Data were analyzed based on the SciAnalysis package provided by the beamline. Current-voltage characteristics were acquired using the Litos Lite characterization system (FLUXiM AG, Switzerland) using the Sinus 70 solar simulator as light source, which is state-of-the-art LED solar simulator with AM1.5G spectrum. The solar cells made from 0.4 M perovskite precursor solutions were encapsulated with epoxy resin while the solar cells made from 1.0 M perovskite precursor solutions were not encapsulated. The solar cells were masked with a black metal mask with a pixel area of 0.0625 cm² and were measured in a forward and reverse bias at a scan rate of 10 mV/s under ambient conditions. External quantum efficiency was recorded with a home-built setup, composed of a tungsten-halogen lamp and of an Oriel Apex monochromator illuminator, which were coupled with a Keithley 2300 and calibrated with a Newport silicon photodiode (UV818). The maximum power point (MPP) tracking also was collected for 60 s after measuring current-voltage characteristics under AM1.5G illumination. The stability measurements of the solar cells were performed by Litos system (FLUXiM AG) under a white light-emitting diode with 100 W/m² light intensity. Aging was implemented under a continuous flow of N₂ in a closed box at 25 °C and no additional encapsulation was used. The active area of device for aging was 0.128 cm².

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132 **Computational Approach**

The first-principles density-functional theory (DFT) calculations were carried out within the Perdew-Burke-Ehrenzof exchange-correlation functional revised for solids (PBEsol)¹ and projector augmented wave (PAW) pseudopotentials^{2,3} as implemented in the Vienna Ab initio Simulation Package (VASP) package. Due to importance of van der Waals interactions in the hybrid perovskites⁴, we applied Tkatchenko-Scheffler scheme^{5,6}. Computational setup including planewave cutoff, kgrid and energy tolerances are properly chosen as tested before⁷. The pure and mixed perovskite phases are modeled using a supercell approach, namely employing a 2×3×1 for the cubic and tetragonal phase and 1×3×1 for the orthorhombic phase. These supercells allow to realize the different compositions synthesized experimentally. Starting from the CsPbI₃, the mixed phase is realized using appropriate Cs for FA or I for Br random substitutions. The organic molecule is inserted with a random orientation in the supercell. Hybrid perovskites are known to have a complex potential energy surface with multiple minima.⁸ Thus, in our structural

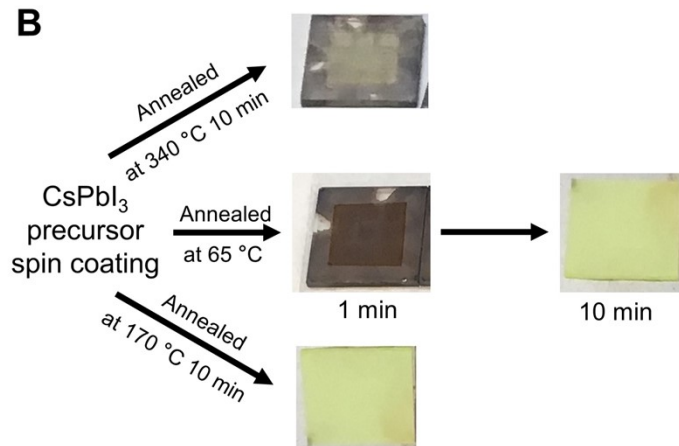
145 optimization, we relaxed the atomic positions and the lattice cell while constraining the system to
146 maintain the initial lattice symmetry.

147 **II. Supplementary figures and tables**

A



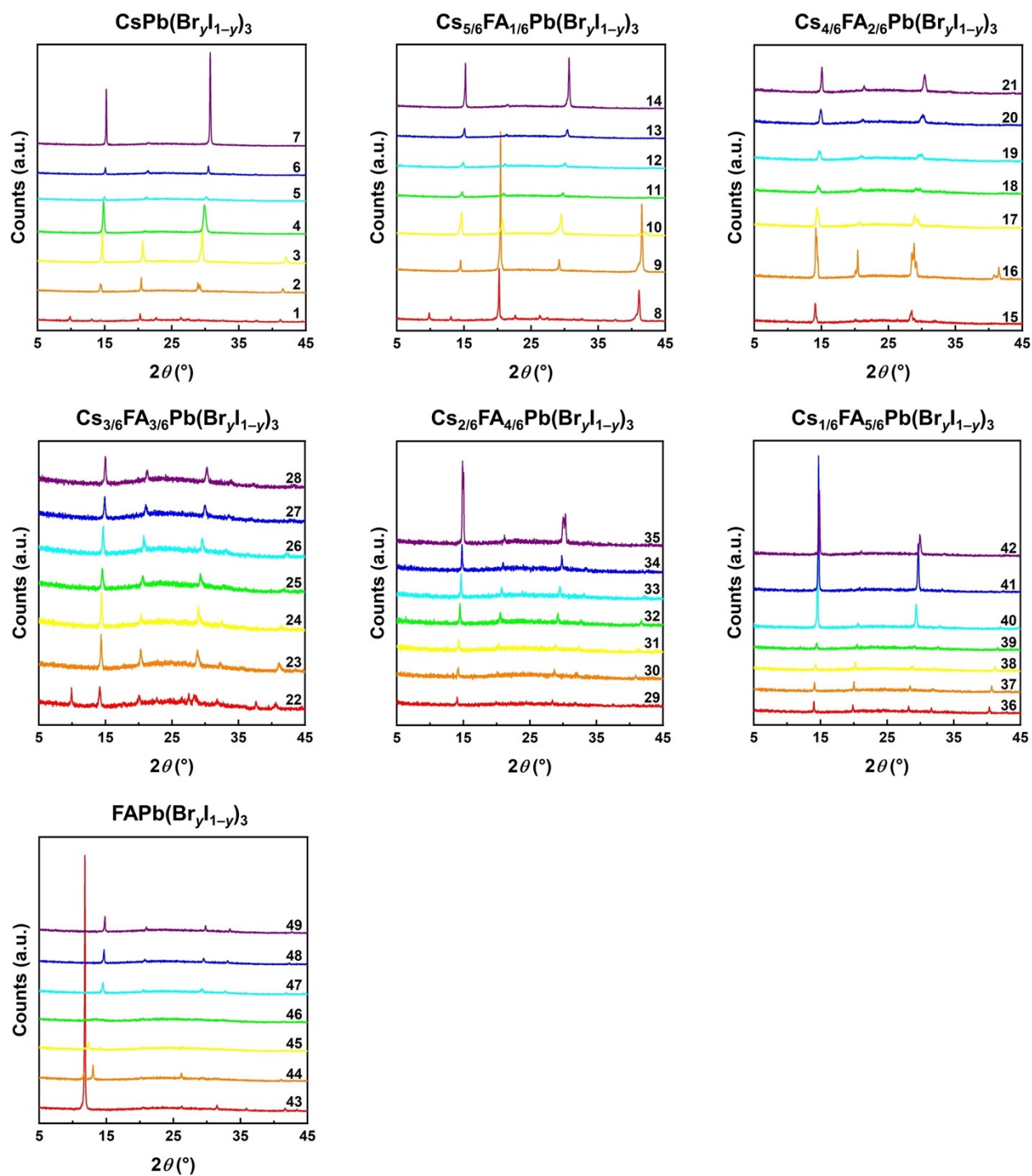
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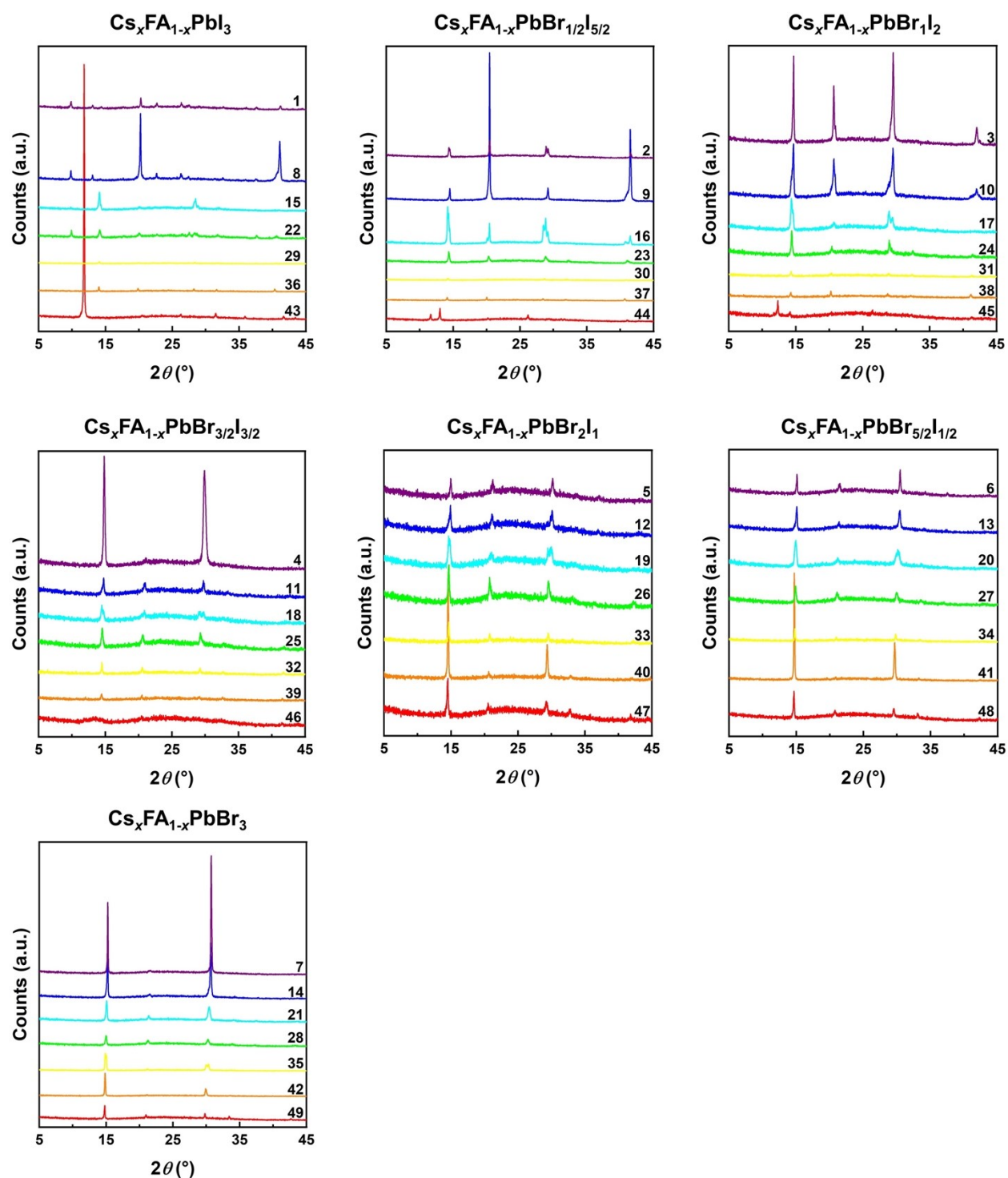
149 Fig. S1 (A) Photograph of the perovskite films (No.8-No.49) annealed at 170 °C for 10 min
150 arranged in the same order as in the composition matrix of Fig. 1A. (B) The different colors of
151 CsPbI₃ films annealed at different temperatures.

153 Crystallographic properties



154
155 Fig. S2 XRD pattern of 49 compositions compared for the halide ratio. Each figure represents a
156 fixed Cs/FA ratio and a changing Br/I ratio, i.e. every row of the composition matrix.

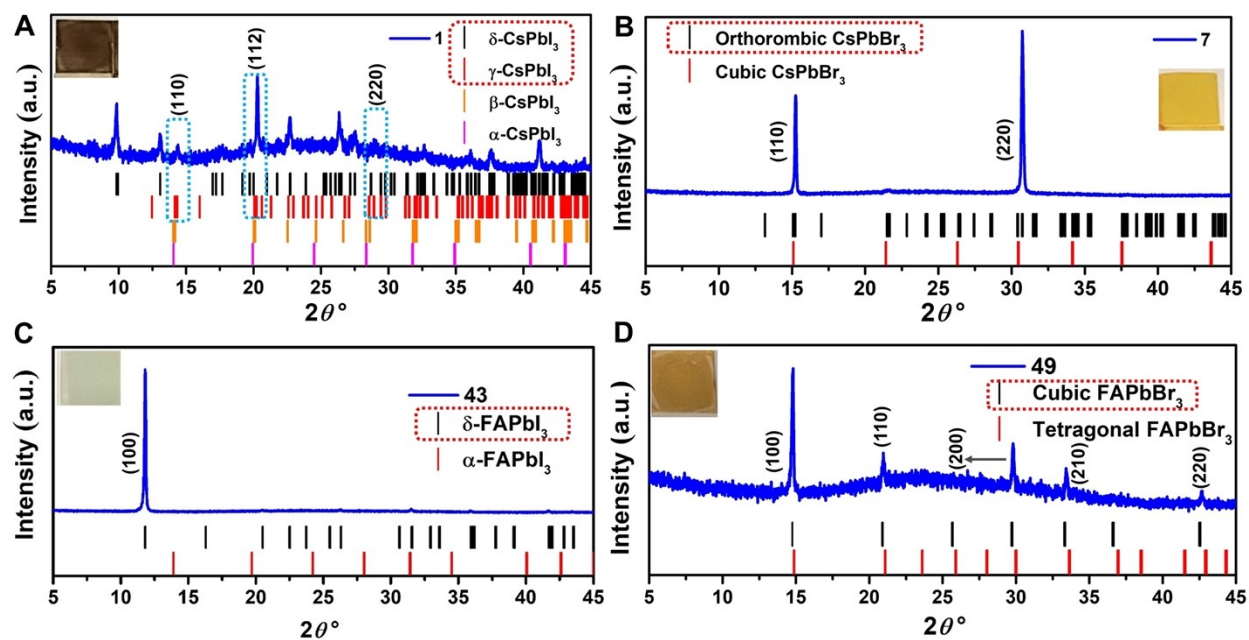
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159 Fig. S3 XRD pattern of 49 compositions compared for the Cs/FA ratio. Each figure represents a
 160 fixed Br/I ratio and a changing Cs/FA ratio, i.e. every column of the composition matrix.

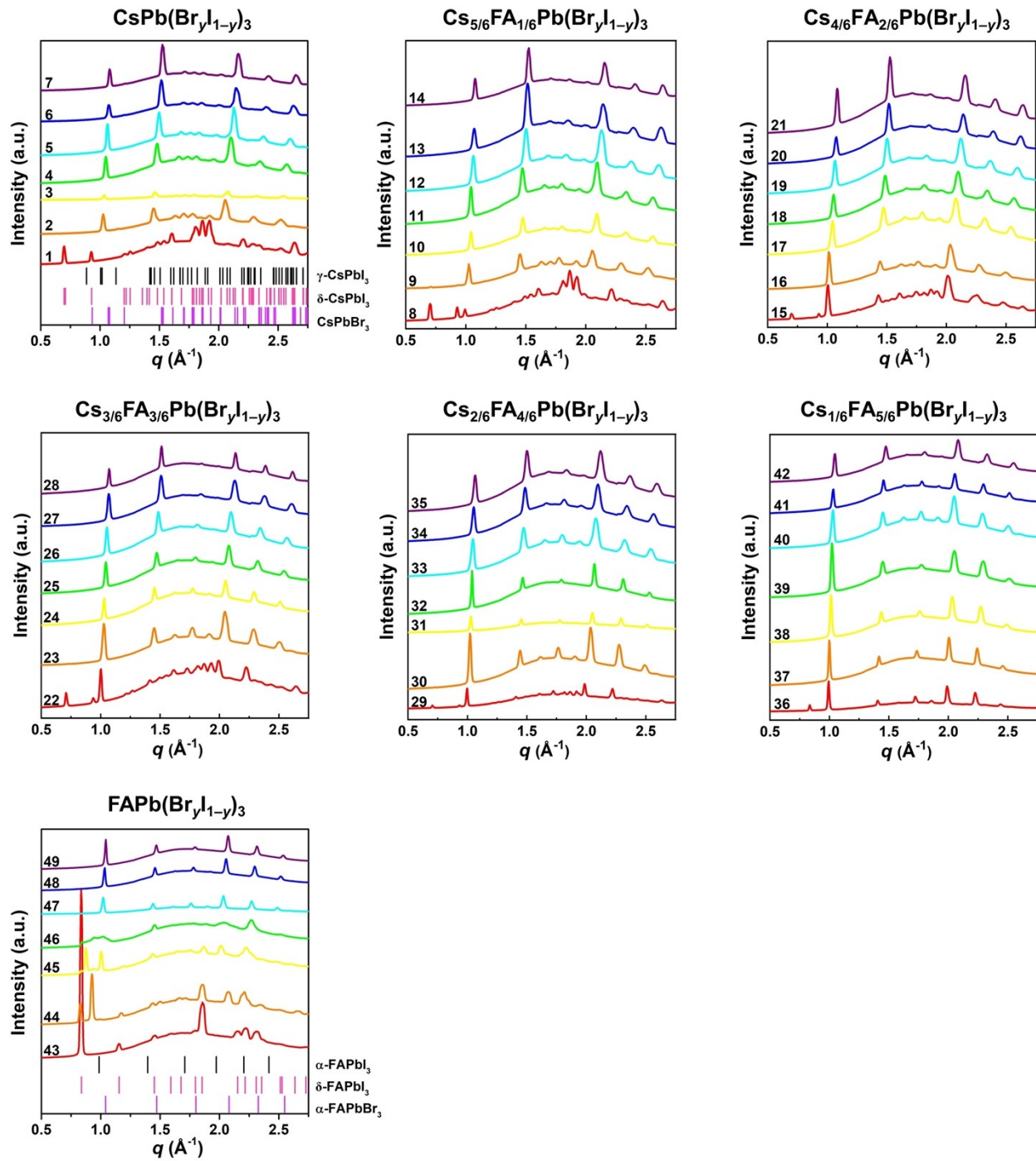
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163 Fig. S4 XRD patterns of four corner compounds, i.e., CsPbI₃ (A), CsPbBr₃ (B), FAPbI₃ (C), and
 164 FAPbBr₃ (D), compared with the calculated XRD peaks from Vesta software. The crystallographic
 165 information files were retrieved from references ^{9–13}.

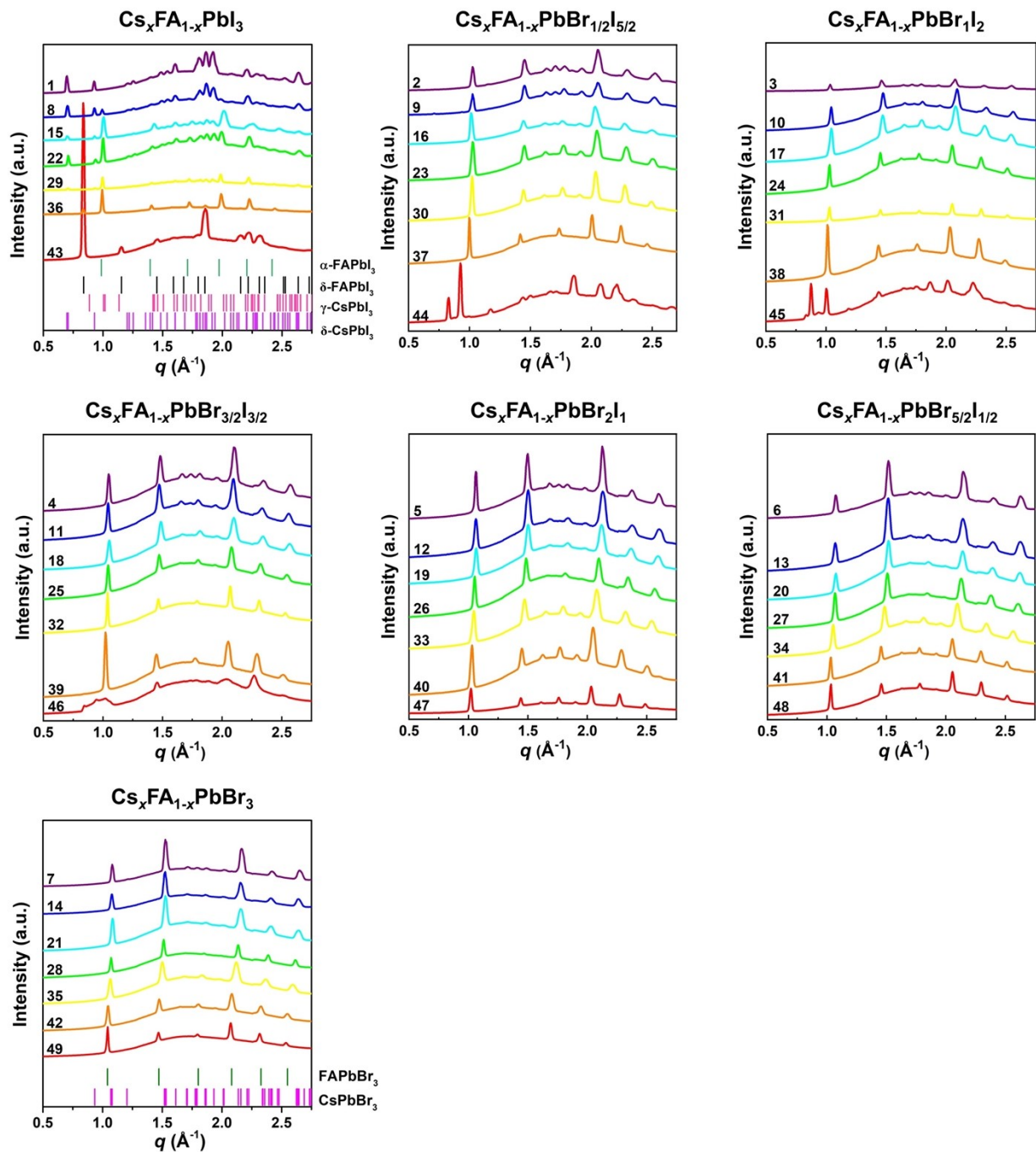
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168 Fig. S5 Circular average of GIWAXS patterns compared for the halide ratio. Each figure represents
 169 a fixed Cs/FA ratio and a changing Br/I ratio, i.e. every row of the composition matrix.

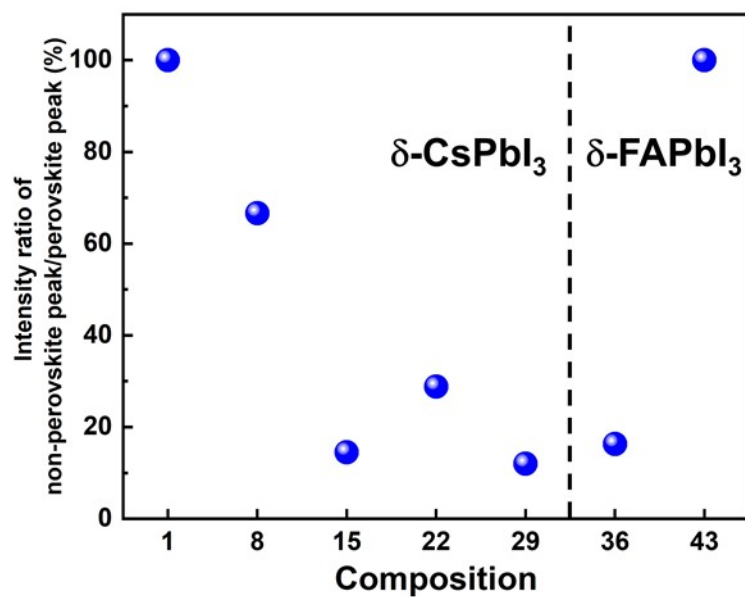
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171

172 Fig. S6 Circular average of GIWAXS patterns compared for the Cs/FA ratio. Each figure
 173 represents a fixed Br/I ratio and a changing Cs/FA ratio, i.e. every column of the composition
 174 matrix.

175

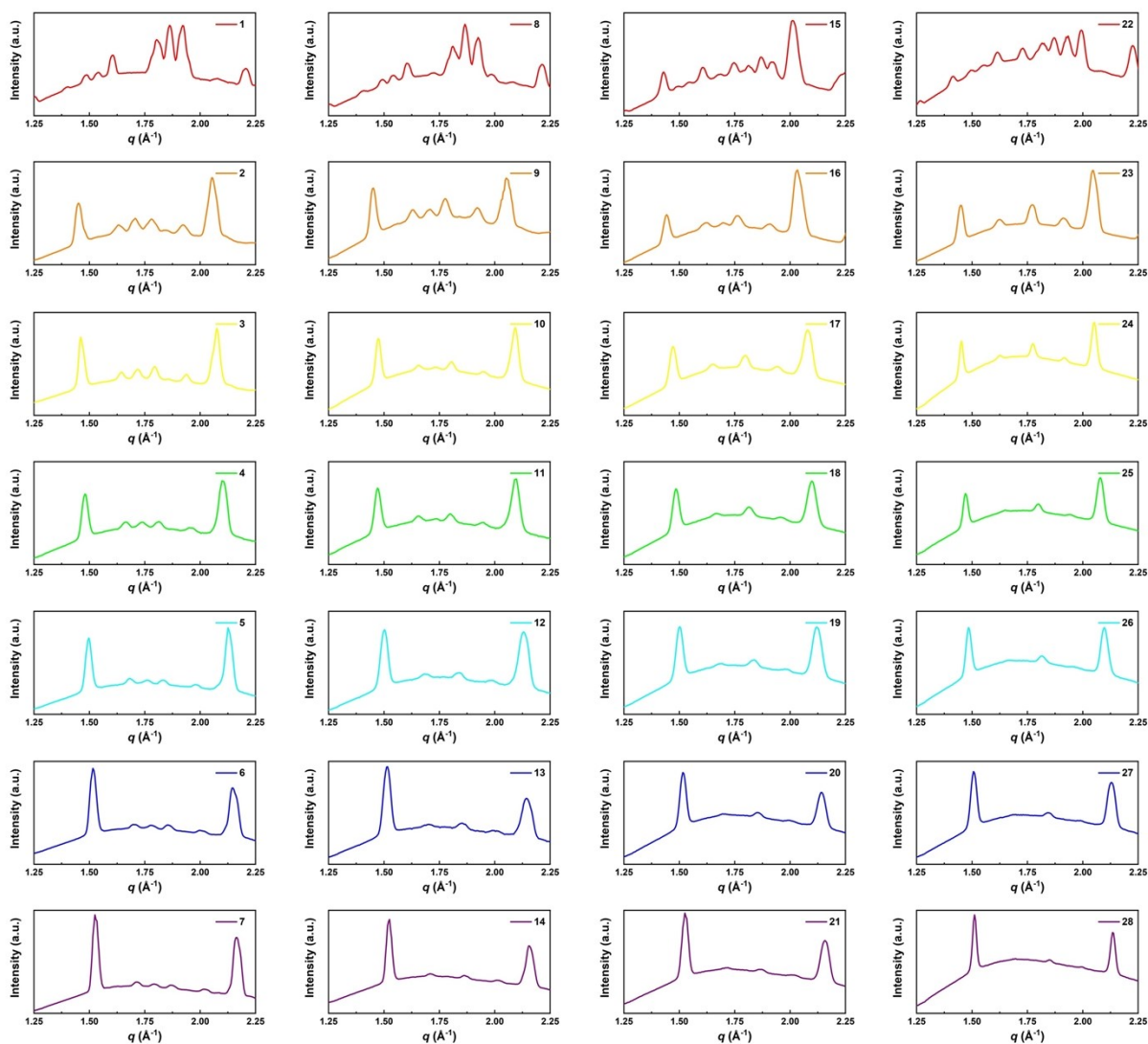


176

177 Fig. S7 The calculated intensity ratio of the characteristic peak at $q = 0.70 \text{ \AA}^{-1}$ ($\delta\text{-CsPbI}_3$) or $q =$
 178 0.83 \AA^{-1} ($\delta\text{-FAPbI}_3$) for non-perovskite phase over characteristic peak at q around 1.00 \AA^{-1} of
 179 perovskite phase for $\text{Cs}_y\text{FA}_{1-y}\text{PbI}_3$ compositions based on GIWAXS data.

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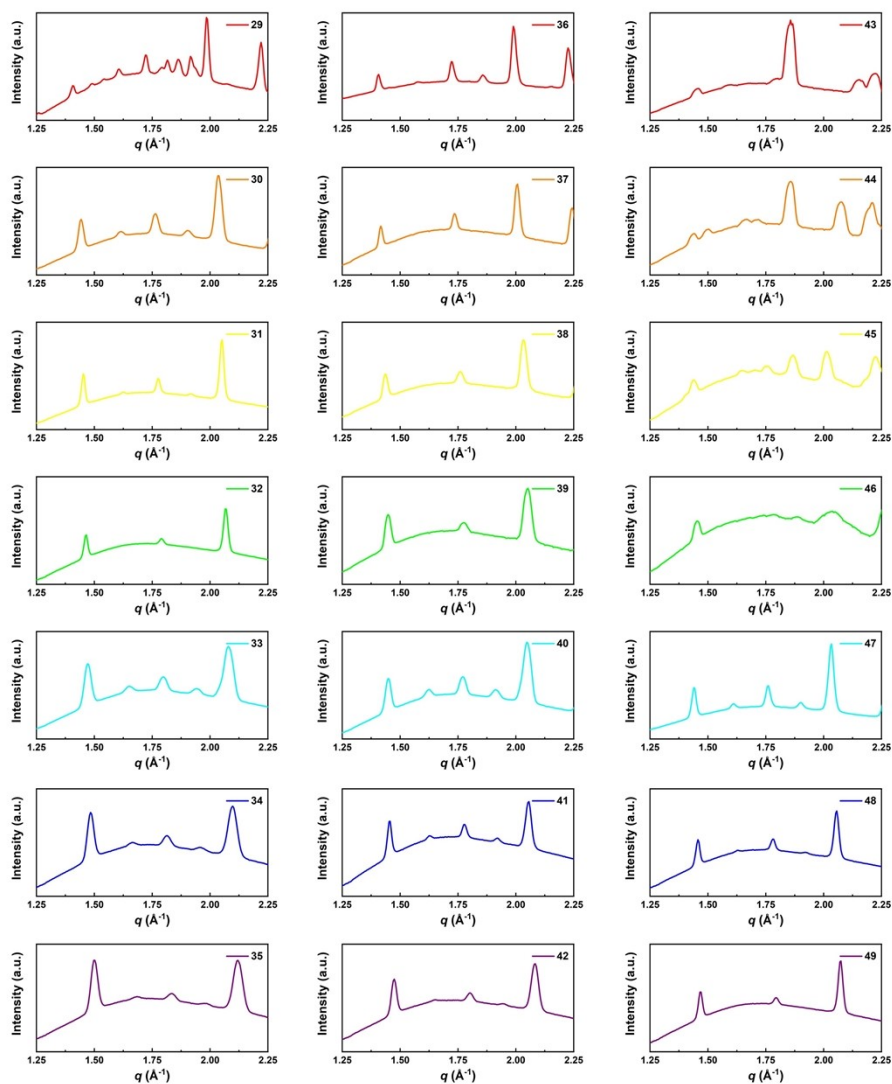
182

183 Fig. S8-1 Enlarged circular average of GIWAXS patterns from $q = 1.25$ - 2.25 \AA^{-1} for composition

184 No.1-28.

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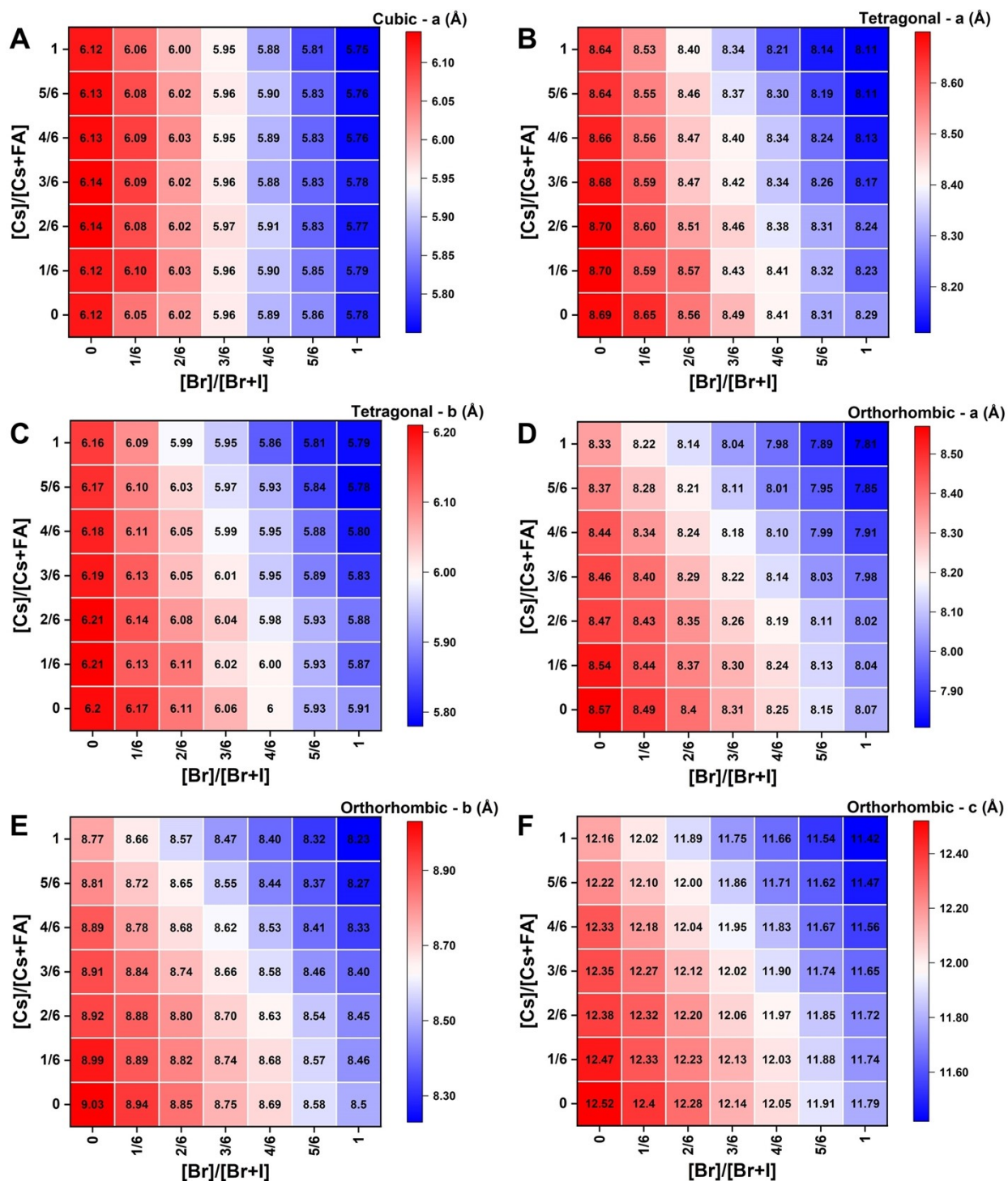


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188 Fig. S8-2 Enlarged circular average of GIWAXS patterns from $q = 1.25$ - 2.25 \AA^{-1} for composition

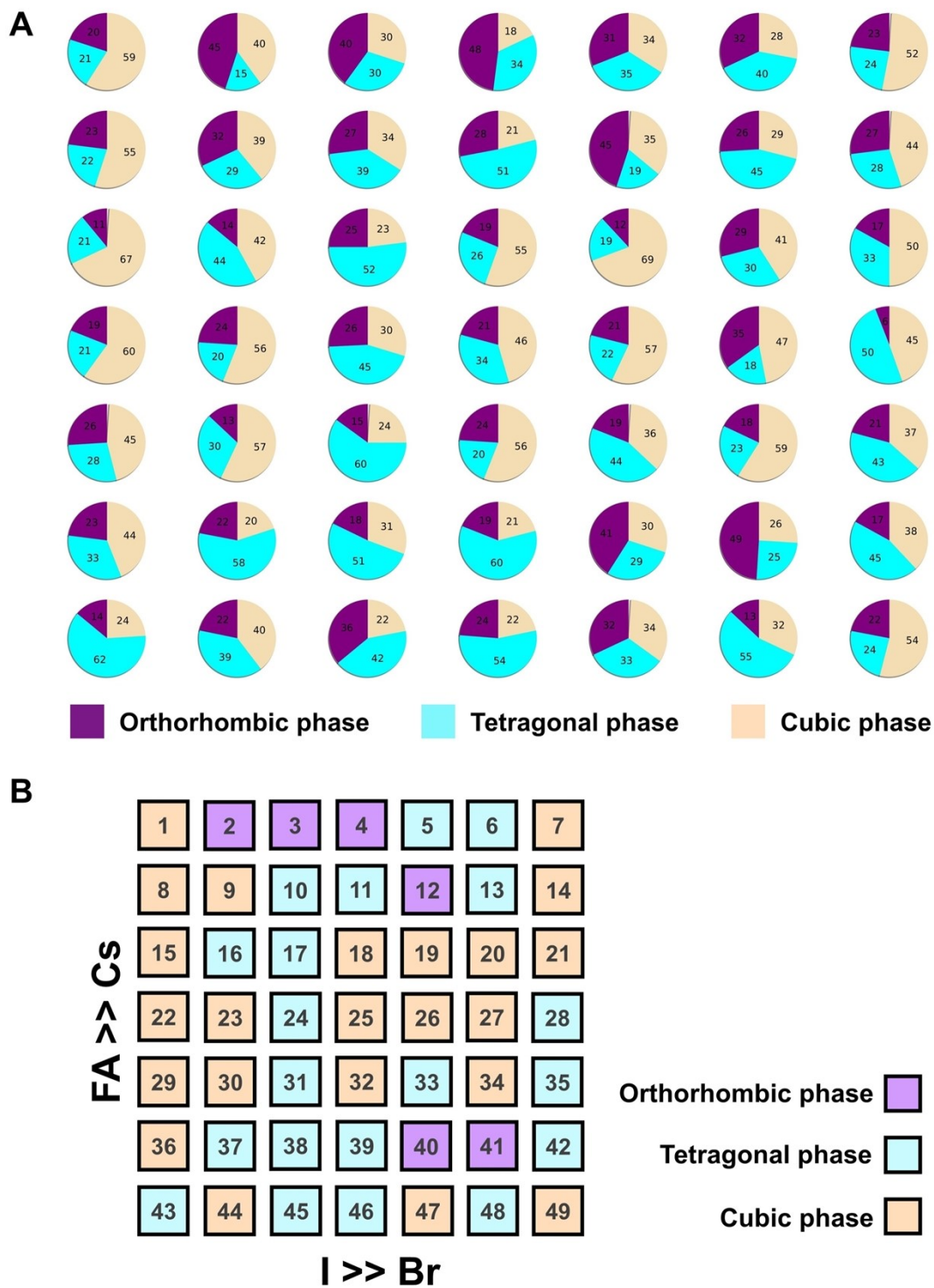
189 No.29-49.

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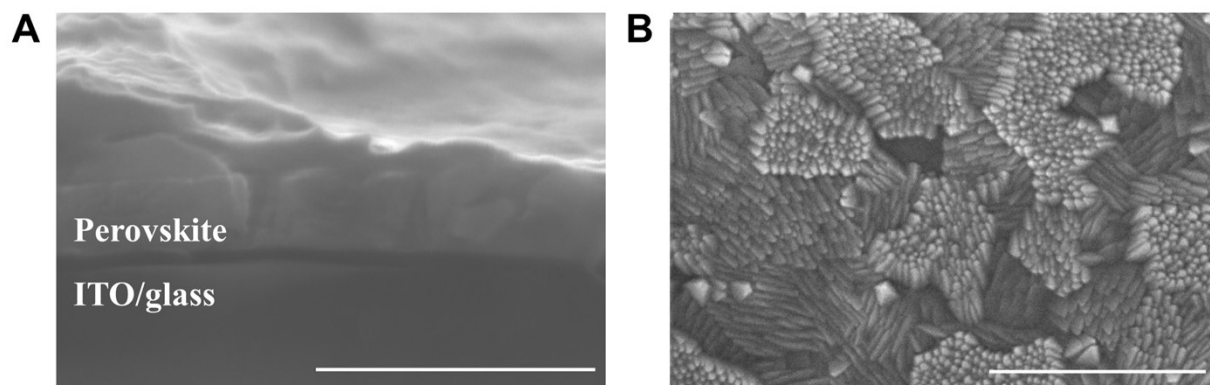
193 Fig. S9 The optimized lattice constants for all 49 compositions and for the three phases cubic (A),
 194 tetragonal (B and C), and orthorhombic (D-E) calculated by DFT.



195

196 Fig. S10 (A) Thermodynamic stability based on Boltzmann distribution at $T = 300^\circ\text{C}$ among cubic,
 197 tetragonal and orthorhombic phases for each composition arranged in the same order as in the
 198 composition matrix of Fig. 1A. (B) Color map of thermodynamically stable phase at room
 199 temperature summarized based on Fig. S9A.

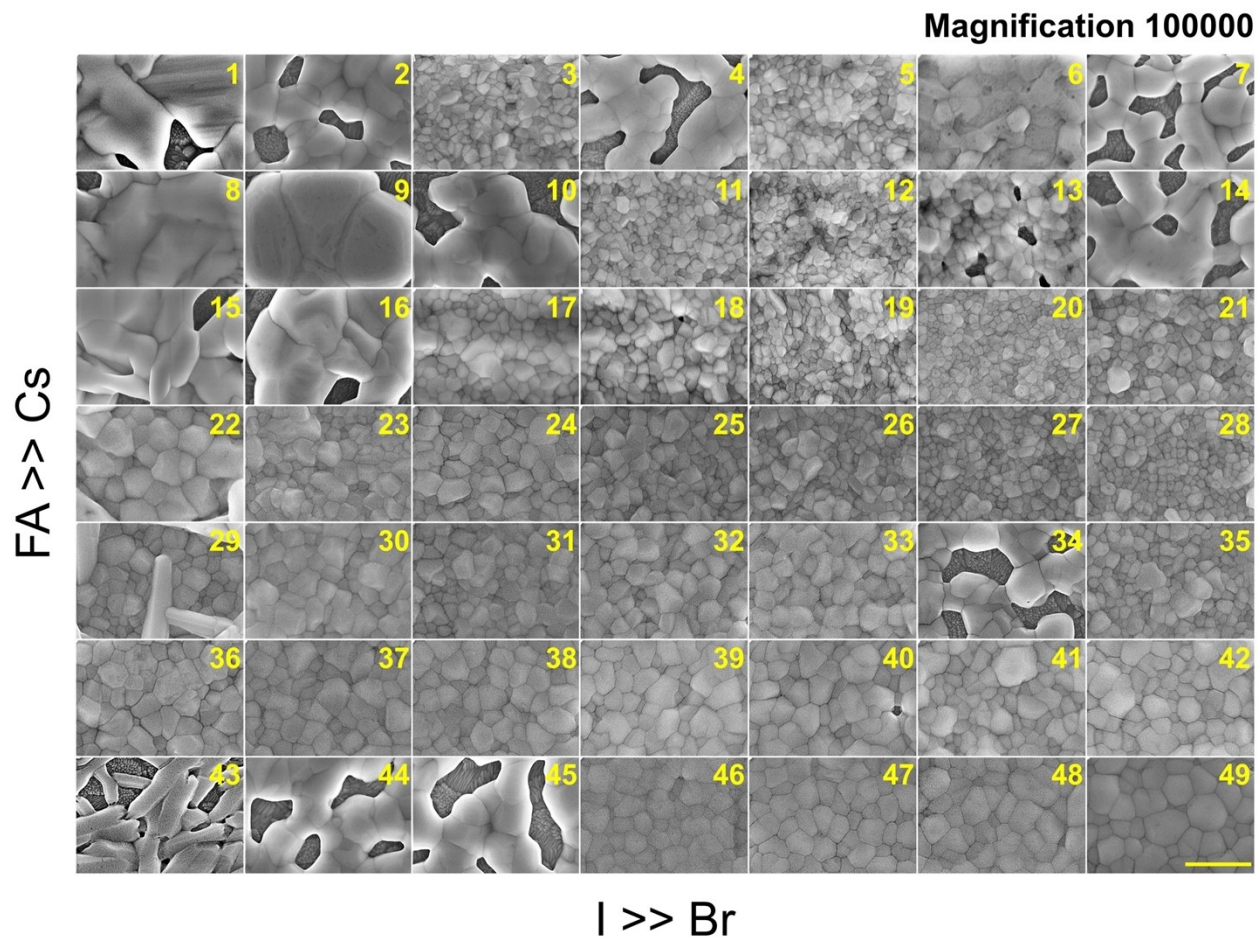
200 Surface morphology



201

202 Fig. S11 (A) SEM image of cross-section view of ITO/perovskite layer. The composition of the
203 perovskite layer is $\text{Cs}_{1/6}\text{FA}_{5/6}\text{PbI}_3$. (B) SEM image of the top view of the bare ITO substrate. The
204 white scale bar in these two SEM images are 500 nm.

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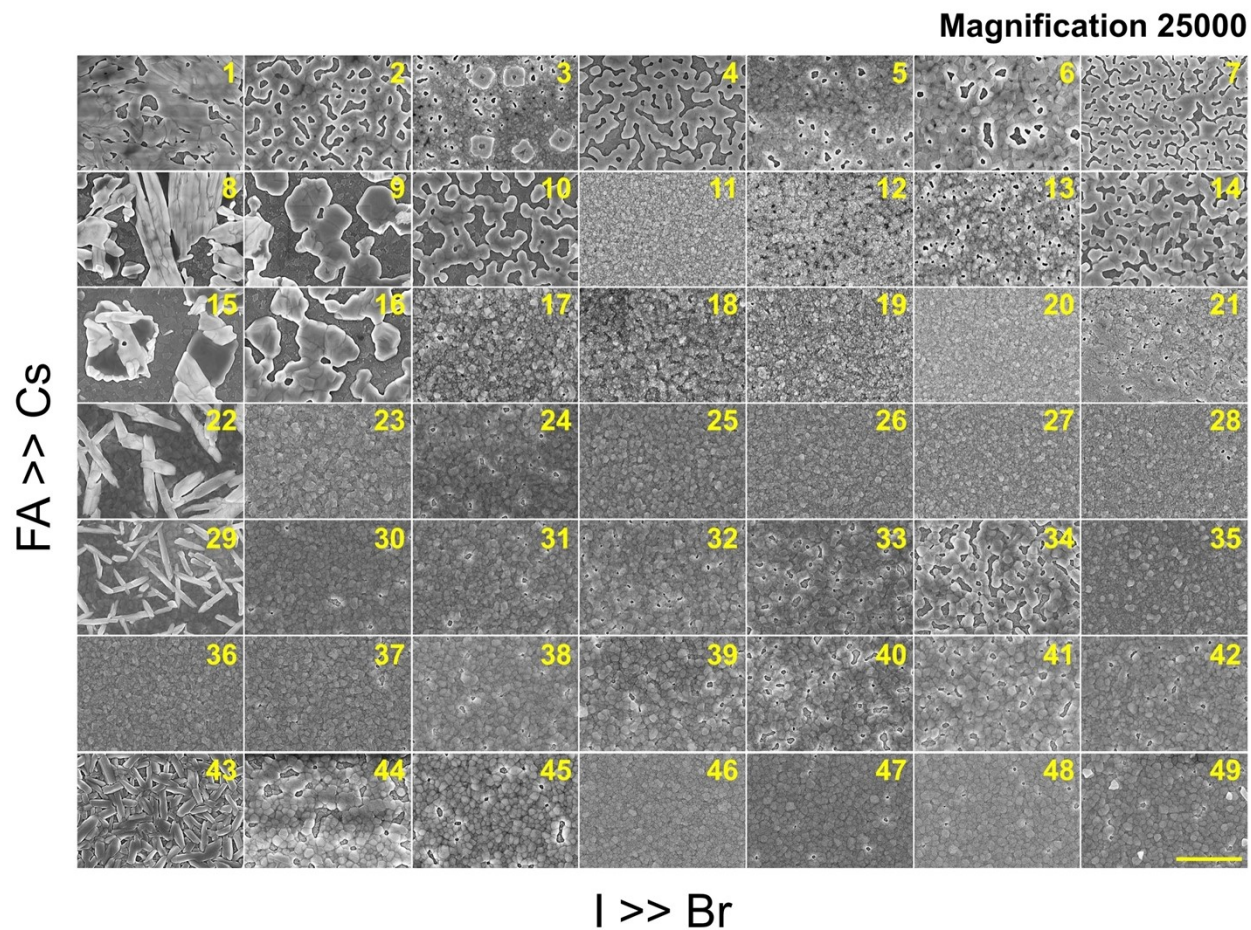
207 Fig. S12 SEM images of the perovskite films in the $\text{Cs}_y\text{FA}_{1-y}\text{Pb}(\text{Br}_x\text{I}_{1-x})_3$ compositional space. The

208 panels are arranged in the same order as in the composition matrix of Fig. 1A in the main text. The

209 yellow scale bar on the right bottom is 500 nm.

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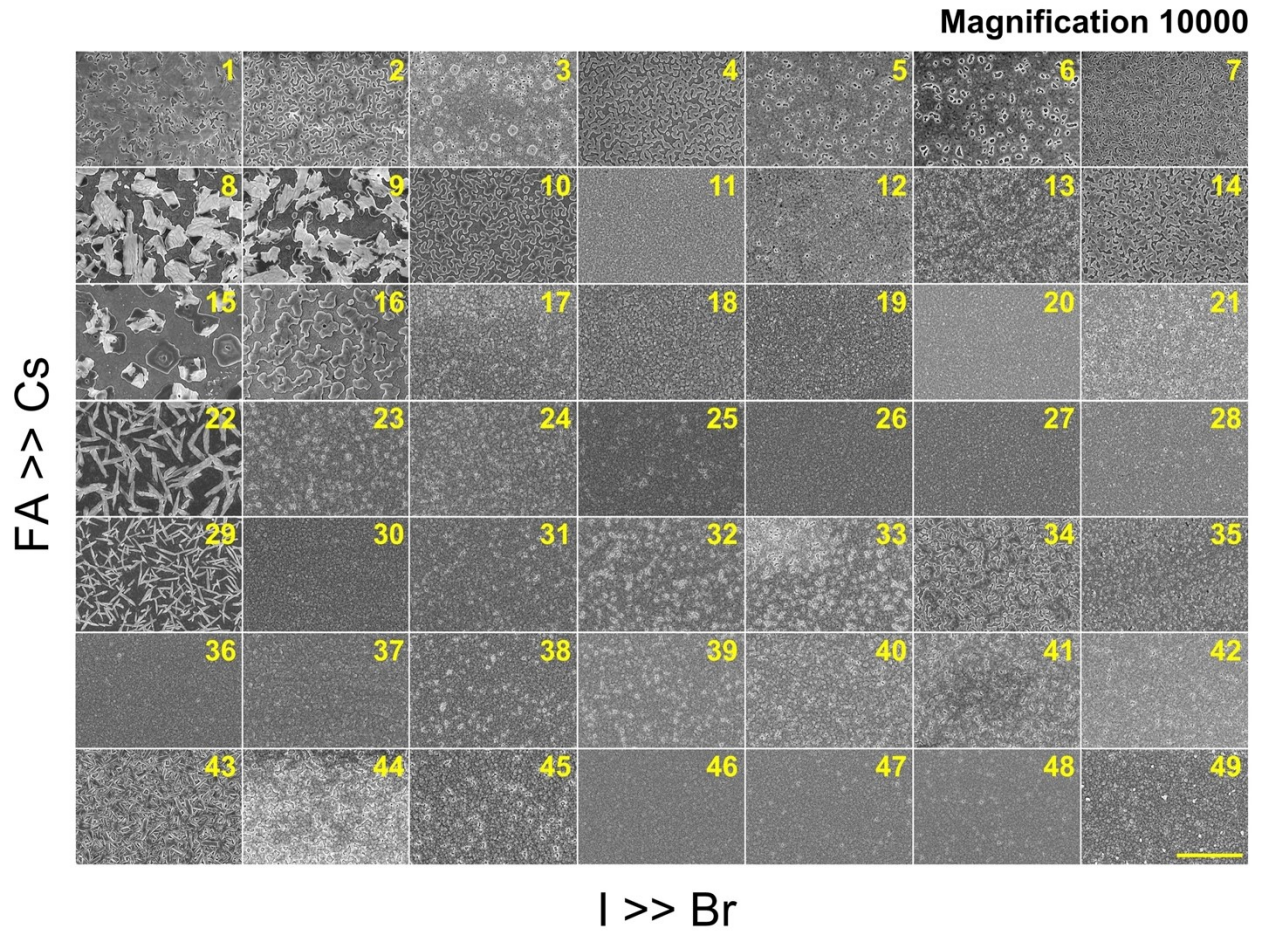


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213 Fig. S13 SEM images of the perovskite films in the $\text{Cs}_y\text{FA}_{1-y}\text{Pb}(\text{Br}_x\text{I}_{1-x})_3$ compositional space. The
 214 panels are arranged in the same order as in the composition matrix of Fig. 1A in the main text. The
 215 yellow scale bar on the right bottom is 2 μm .

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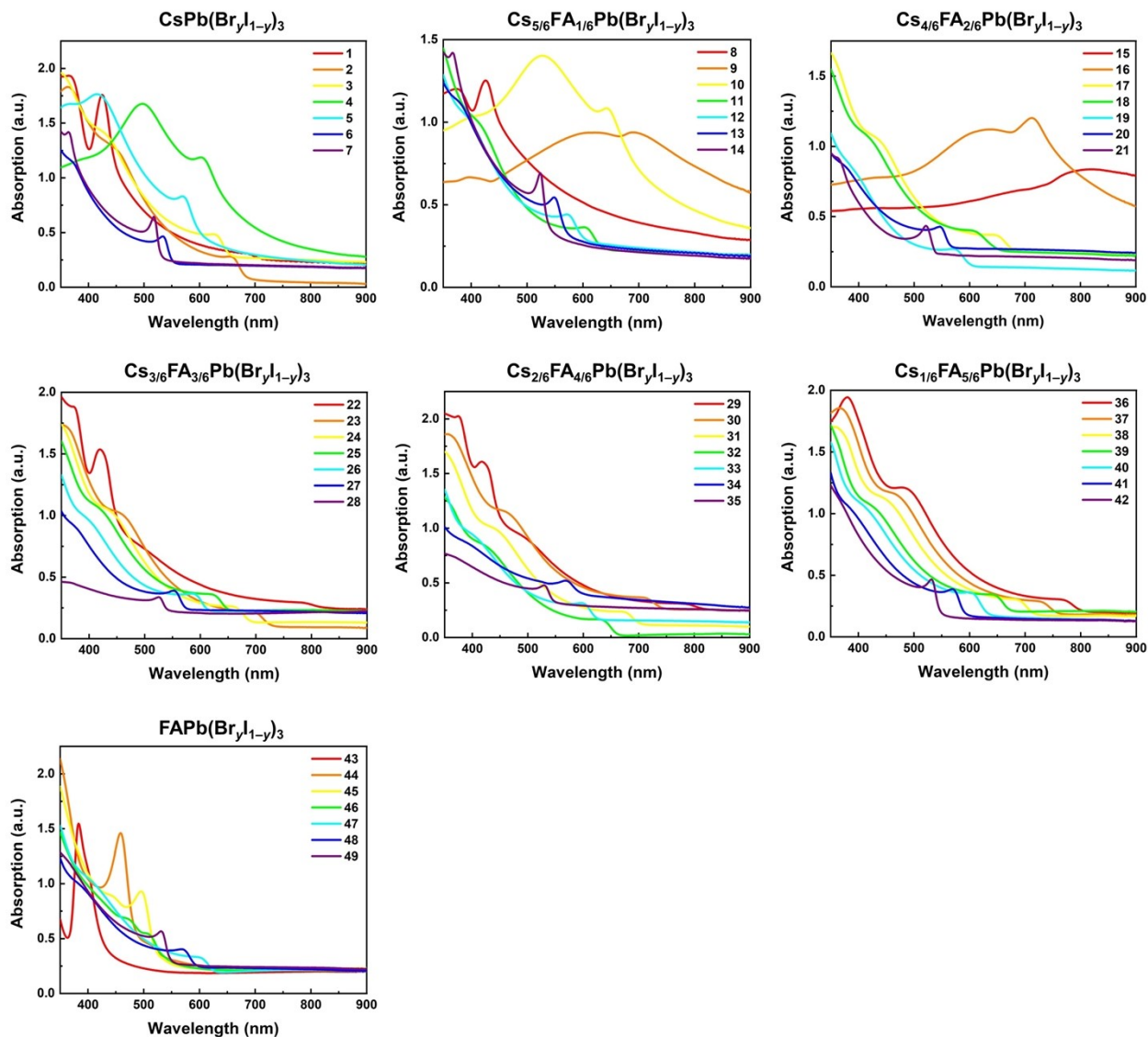


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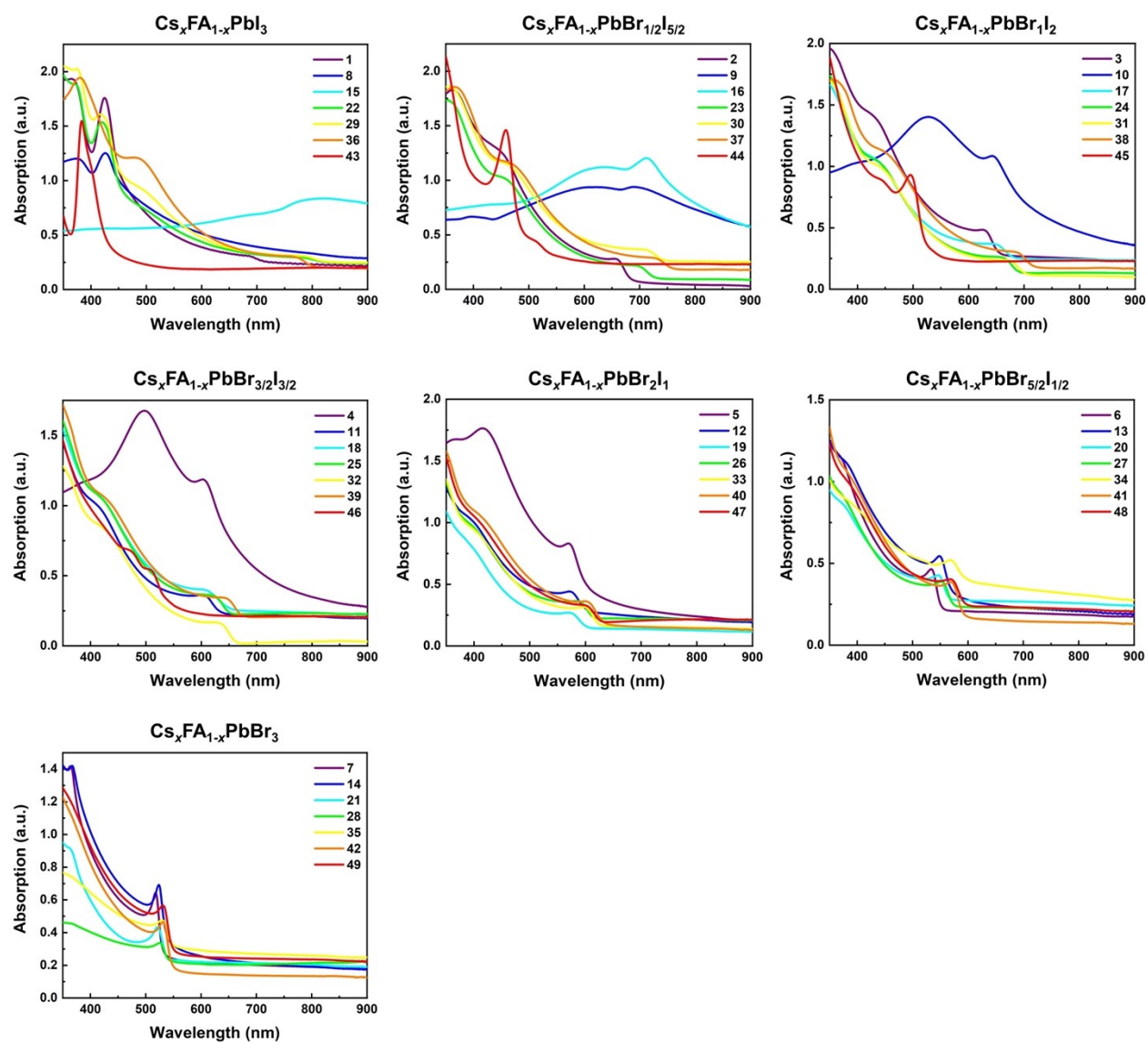
219 Fig. S14 SEM images of the perovskite films in the $\text{Cs}_y\text{FA}_{1-y}\text{Pb}(\text{Br}_x\text{I}_{1-x})_3$ compositional space. The
 220 panels are arranged in the same order as in the composition matrix of Fig. 1A in the main text. The
 221 yellow scale bar on the right bottom is 5 μm .

222

223 **Optical properties**

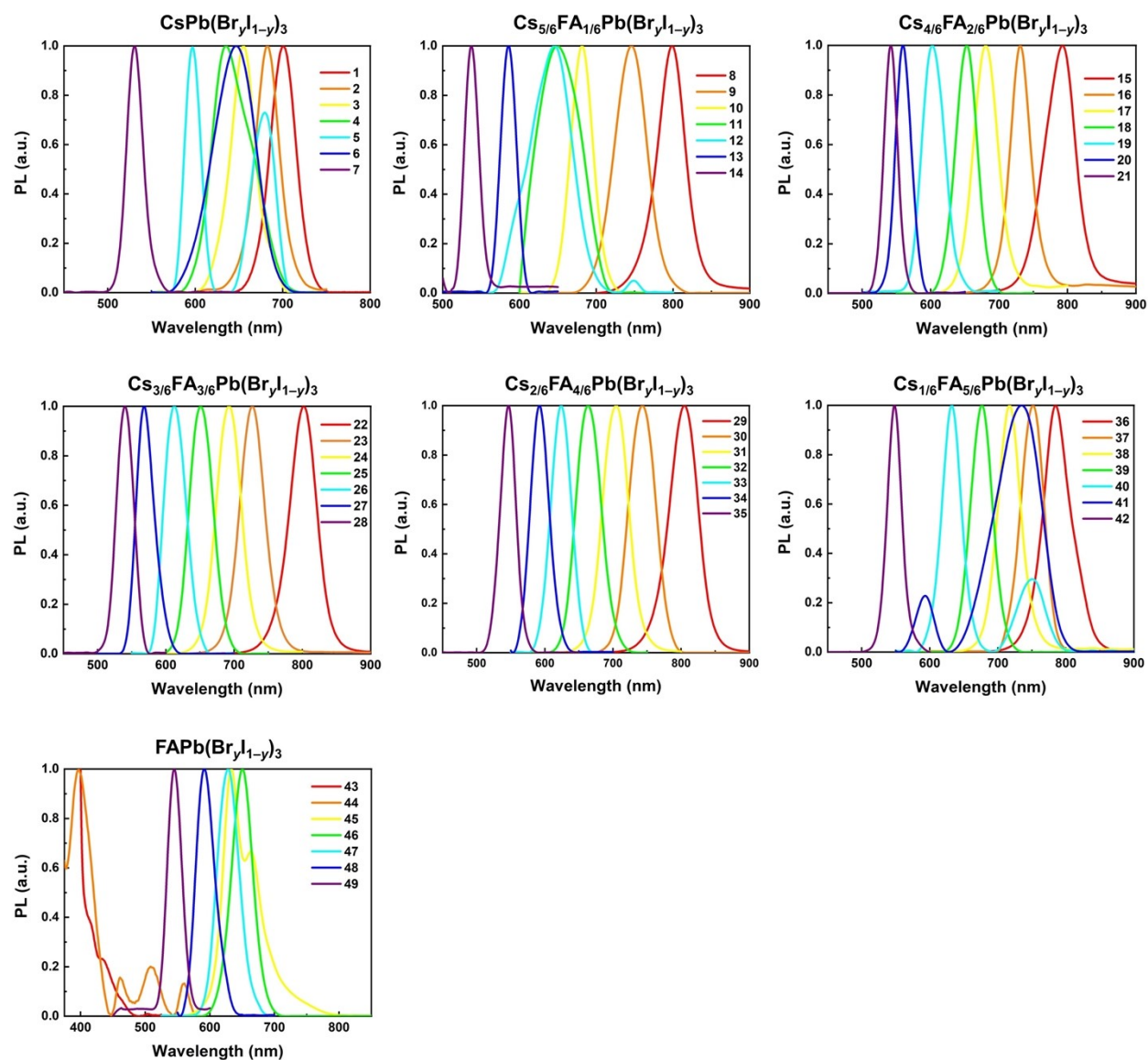


224
225 Fig. S15 Ultraviolet-visible absorbance as a function of wavelength. Each figure represents a fixed
226 Cs/FA ratio and a changing Br/I ratio, i.e. every row of the composition matrix.
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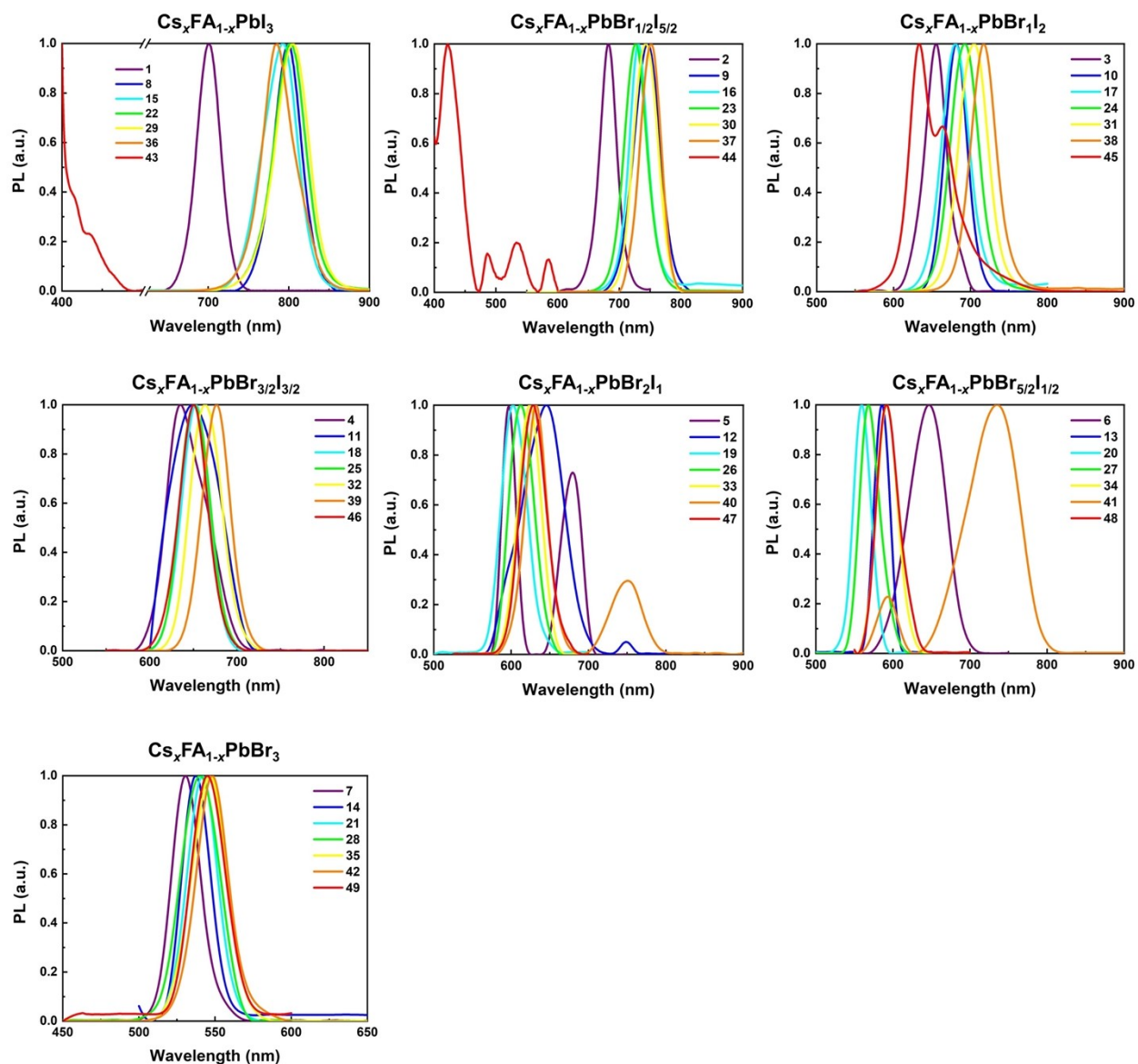
229 Fig. S16 Ultraviolet-visible absorbance as a function of wavelength. Each figure represents a fixed
 230 Br/I ratio and a changing Cs/FA ratio, i.e. every column of the composition matrix.



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233 Fig. S17 Normalized photoluminescence spectra as a function of wavelength. Each figure
 234 represents a fixed Cs/FA ratio and a changing Br/I ratio, i.e. every row of the composition matrix.

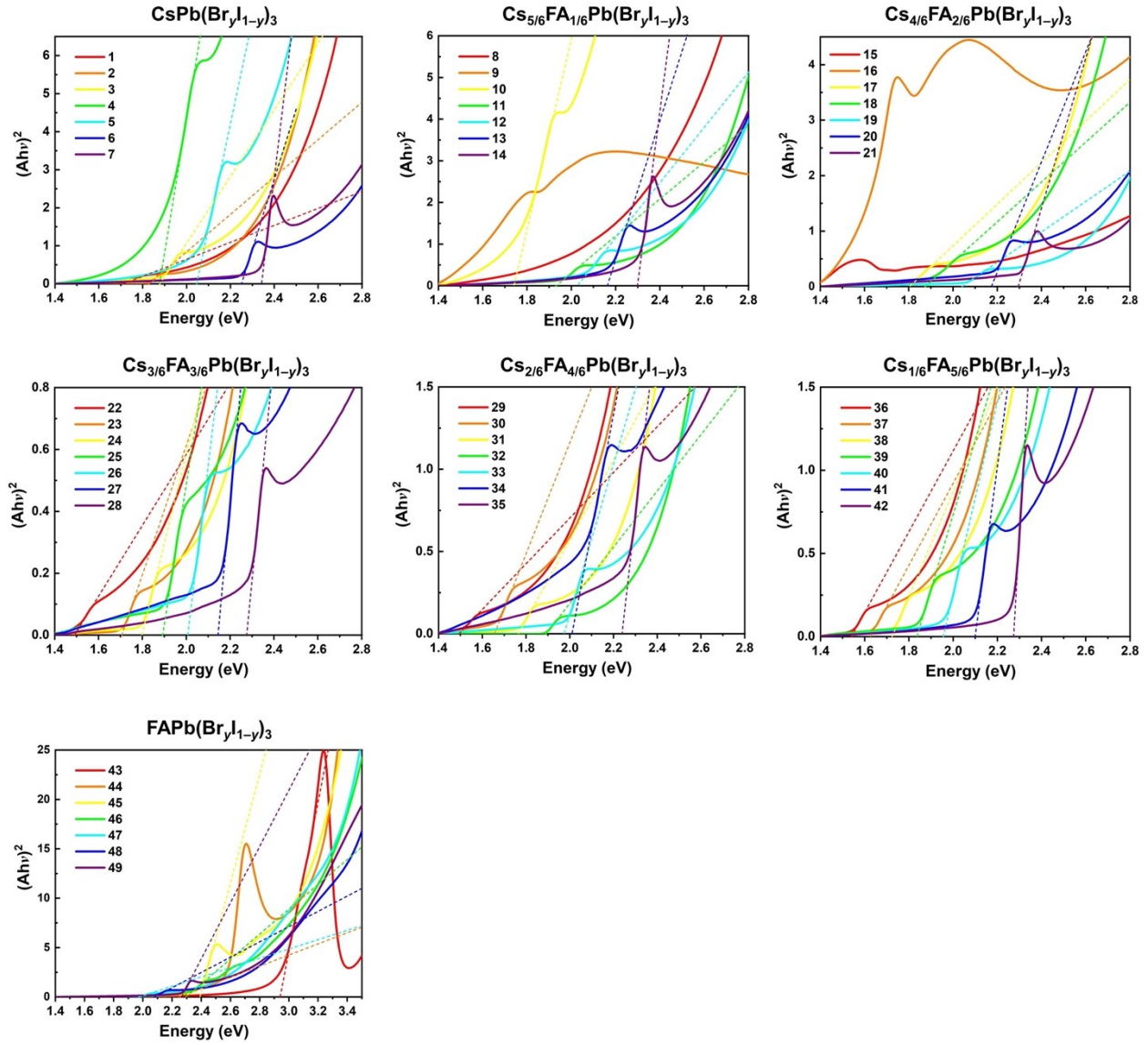
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237 Fig. S18 Normalized photoluminescence spectra as a function of wavelength. Each figure
 238 represents a fixed Br/I ratio and a changing Cs/FA ratio, i.e. every column of the composition
 239 matrix.

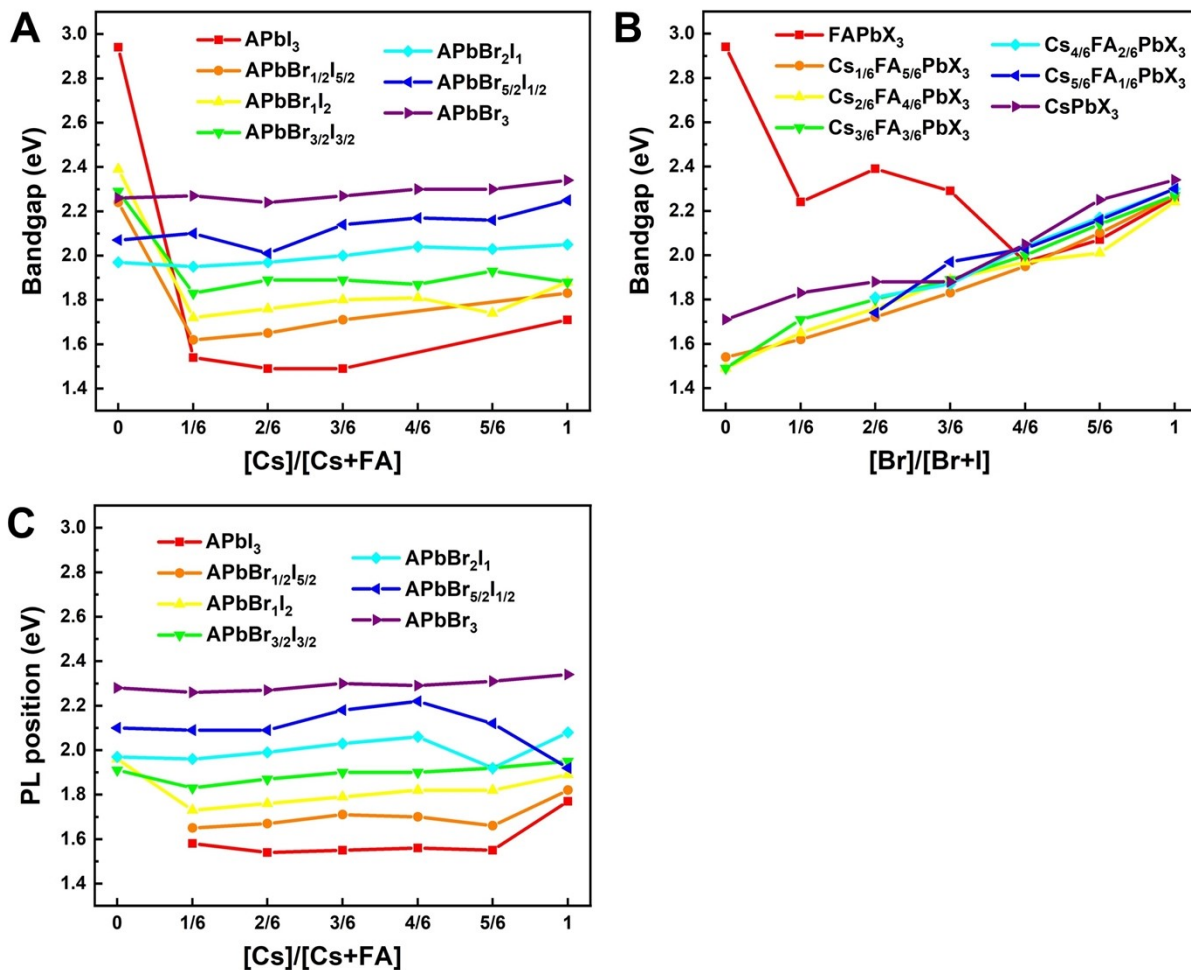
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241

242 Fig. S19 Tauc plots based on the square of absorption as a function of photon energy. The bandgaps
 243 are determined by the intercept between the linear regions close to the absorption onset and the
 244 intercept. Each figure represents a fixed Cs/FA ratio and a changing Br/I ratio, i.e. every row of
 245 the composition matrix.

246



247

248 Fig. S20 Bandgaps as a function of Cs content (A) and Br content (B). Bandgap values were
 249 calculated from Tauc plots in Fig. S18. (C) The PL energy extracted from the maximum emission
 250 wavelength of PL as a function of Cs content.

251

252 Table. S1 Values of the coefficients in eqn. (2) of main text with 95% confidence bounds

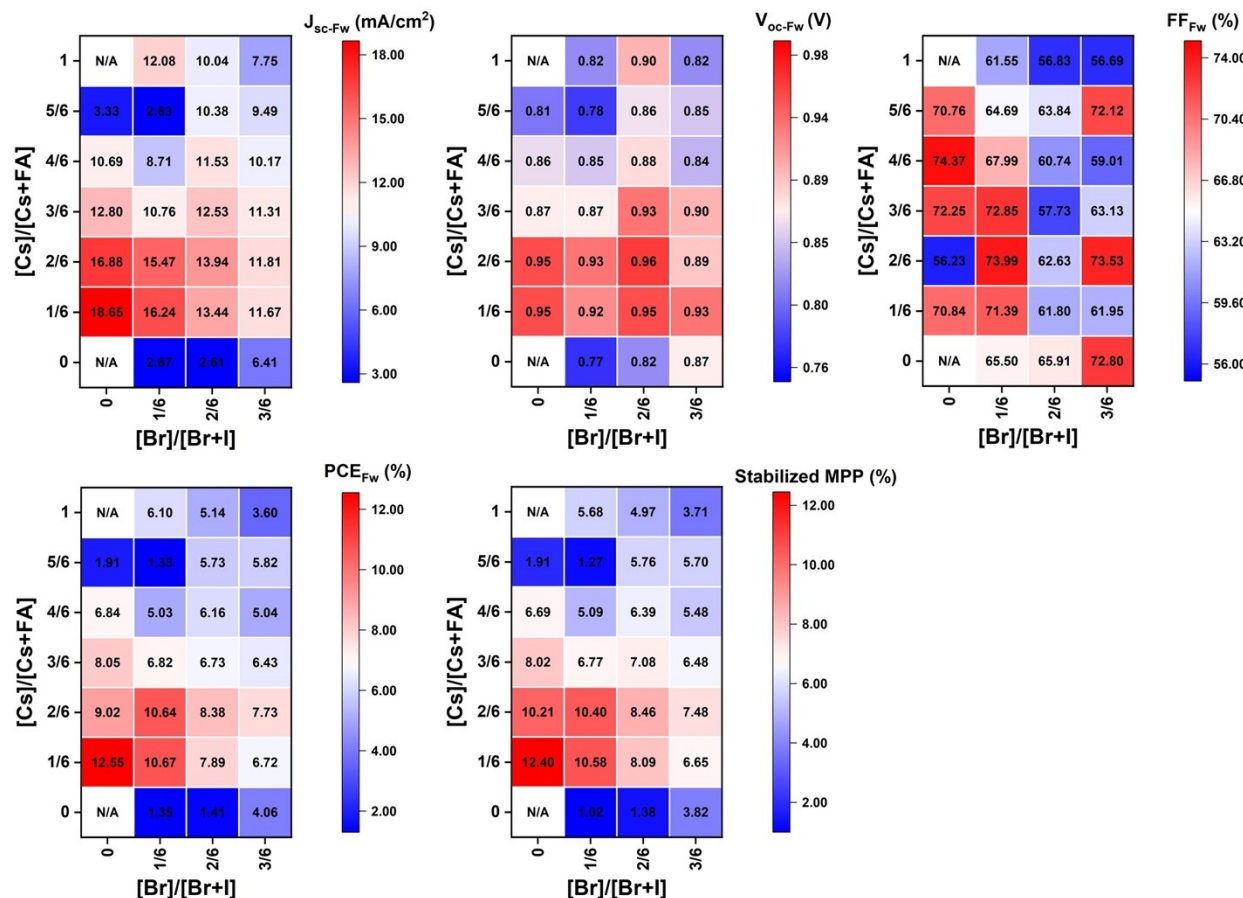
Coefficient	Value	Standard error	Correlation
C_1	1.502	0.033	0.965
C_2	0.560	0.076	0.985
C_3	0.080	0.096	0.989
C_4	0.185	0.065	0.970
C_5	0.106	0.071	0.972
C_6	-0.097	0.063	0.939

253 Note:

- 254 1. The model of the fitting second-order polynomial is $E_g(x, y) = C_1 + C_2 \cdot x + C_3 \cdot y + C_4 \cdot x^2 + C_5 \cdot$
255 $y^2 + C_6 \cdot x \cdot y$. The eqn. (2) was fitted by Origin software.
- 256 2. R-square value is also known as the coefficient of determination. The closer the fit is to the
257 data points, the closer r-square will be to the value of 1. A larger value of r-square does not
258 necessarily mean a better fit because the degree of freedom can also affect the value. Thus if
259 more parameters are introduced, the r-square value will rise, but this does not imply a better
260 fit. The adjusted r-square value accounts for the degrees of freedom to provide a better measure
261 of the goodness of fit.
- 262

264 Device performance

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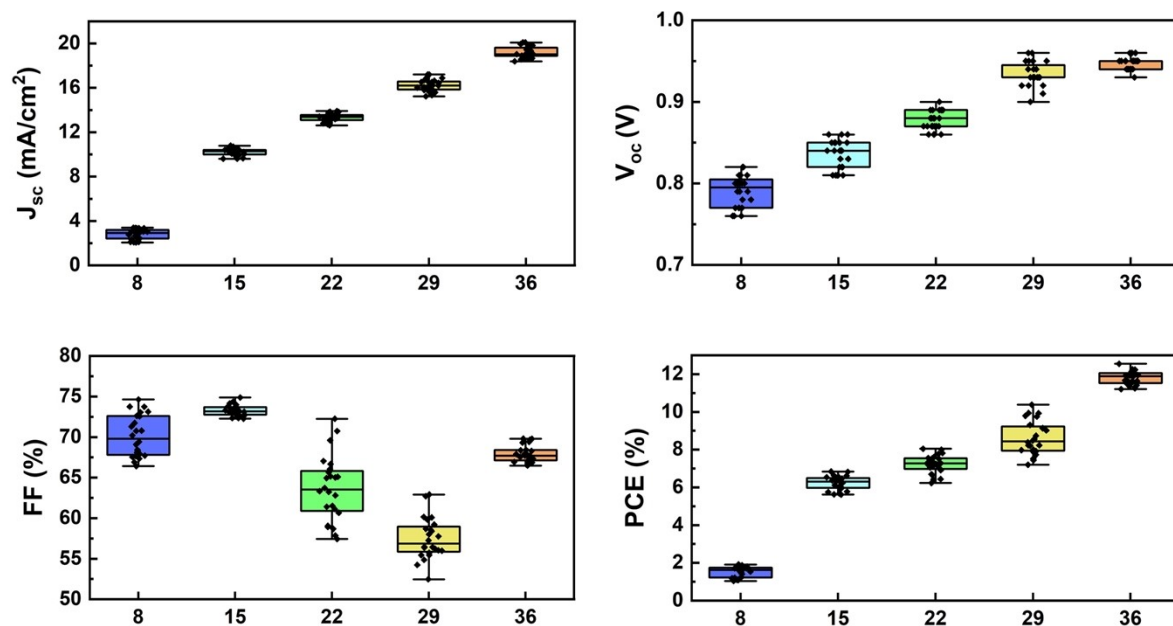


266

267 Fig.S21 The photovoltaic parameters of J_{sc} , V_{oc} , FF, and PCE for forward scan and stabilized power
268 output at MPP attained from best devices of I-rich compositions using 0.4 M perovskite precursor
269 solutions.

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271

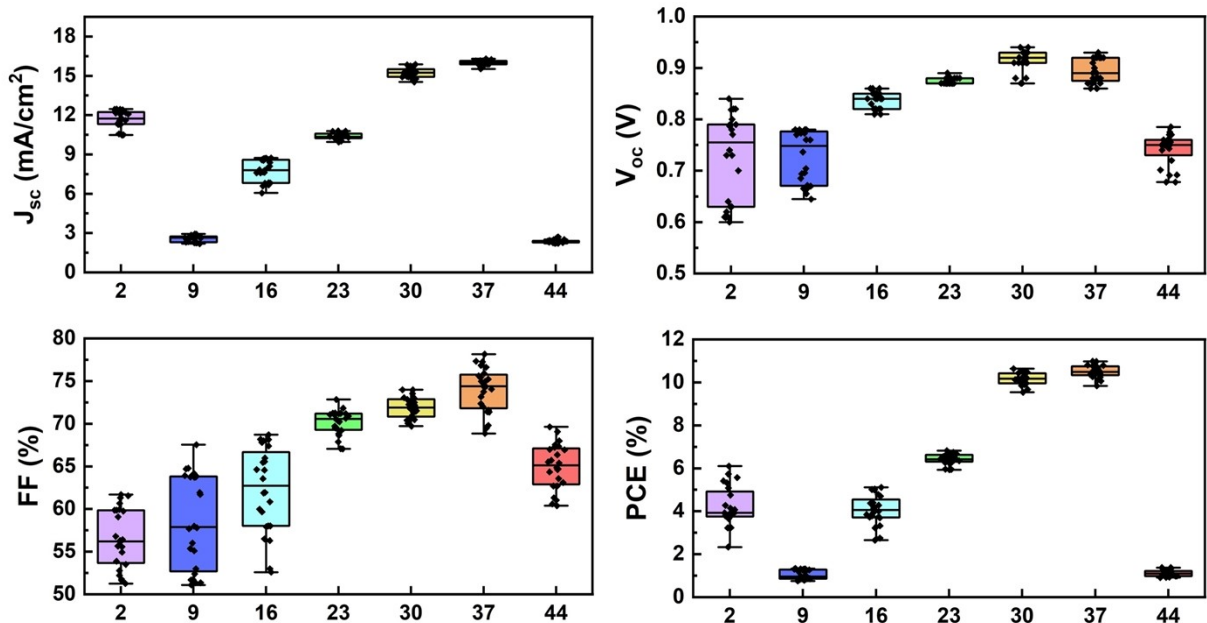


272

273 Fig.S22 Box plots of the photovoltaic parameters of the perovskite solar cells with composition

274 $Cs_yFA_{1-y}PbI_3$.

275

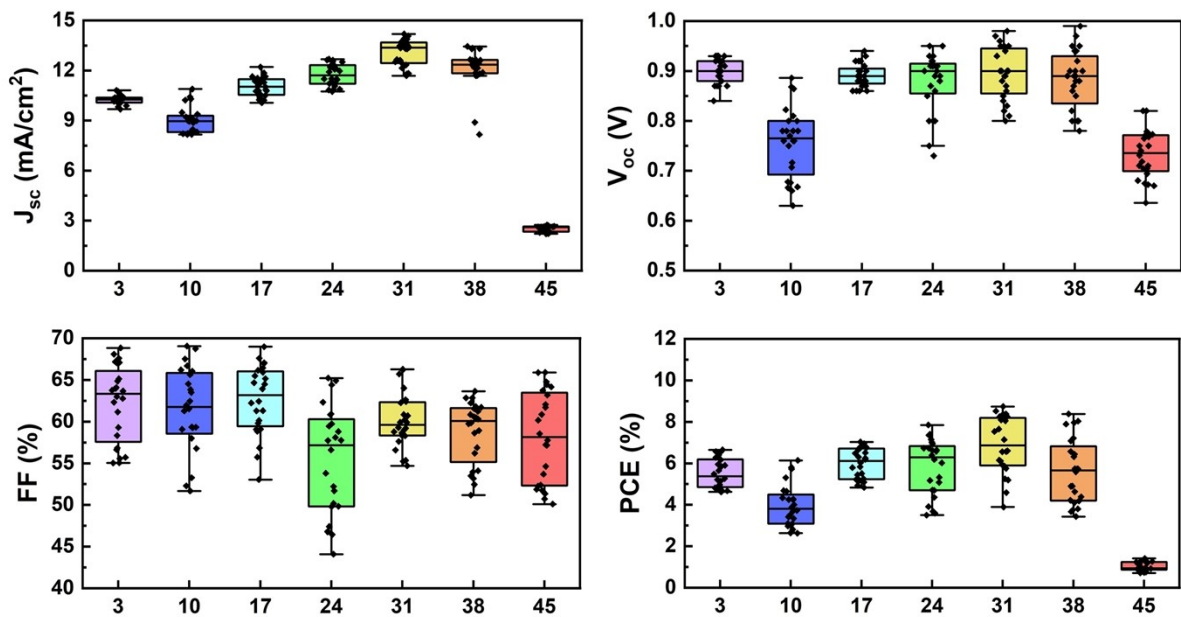


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277 Fig.S23 Box plots of the photovoltaic parameters of the perovskite solar cells with composition

278 $\text{Cs}_y\text{FA}_{1-y}\text{PbBr}_{1/2}\text{I}_{5/2}$.

279



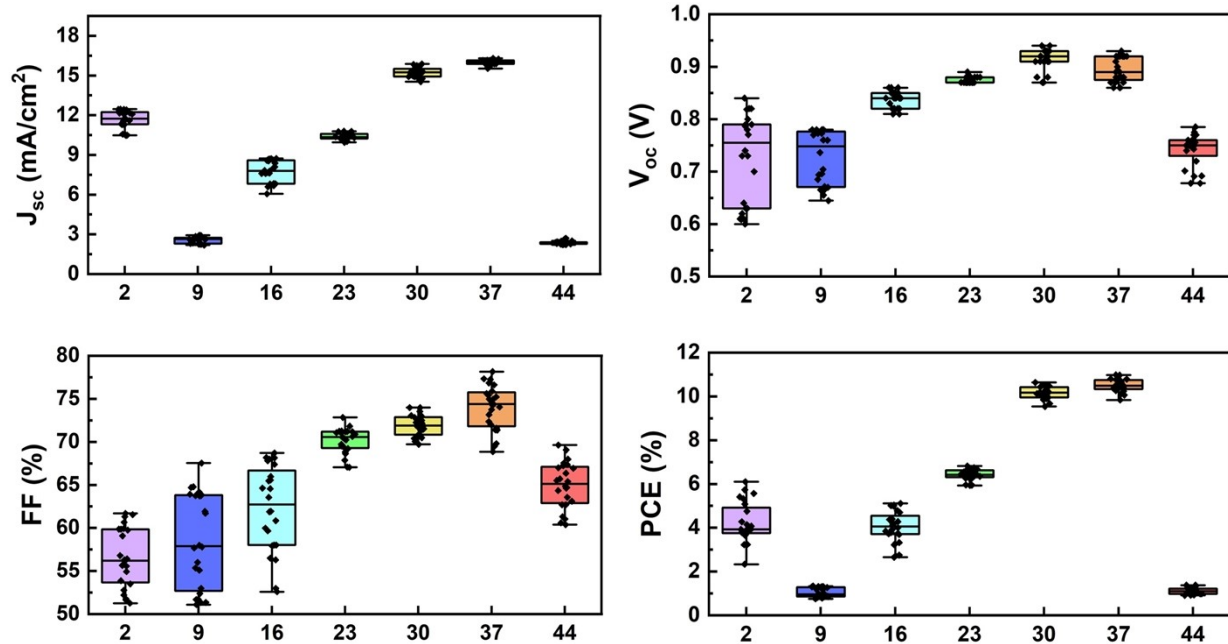
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281 Fig.S24 Box plots of the photovoltaic parameters of the perovskite solar cells with composition

282 $\text{Cs}_y\text{FA}_{1-y}\text{PbBr}_1\text{I}_2$.

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286 Fig.S25 Box plots of the photovoltaic parameters of the perovskite solar cells with composition

287 $\text{Cs}_y\text{FA}_{1-y}\text{PbBr}_{3/2}\text{I}_{3/2}$.

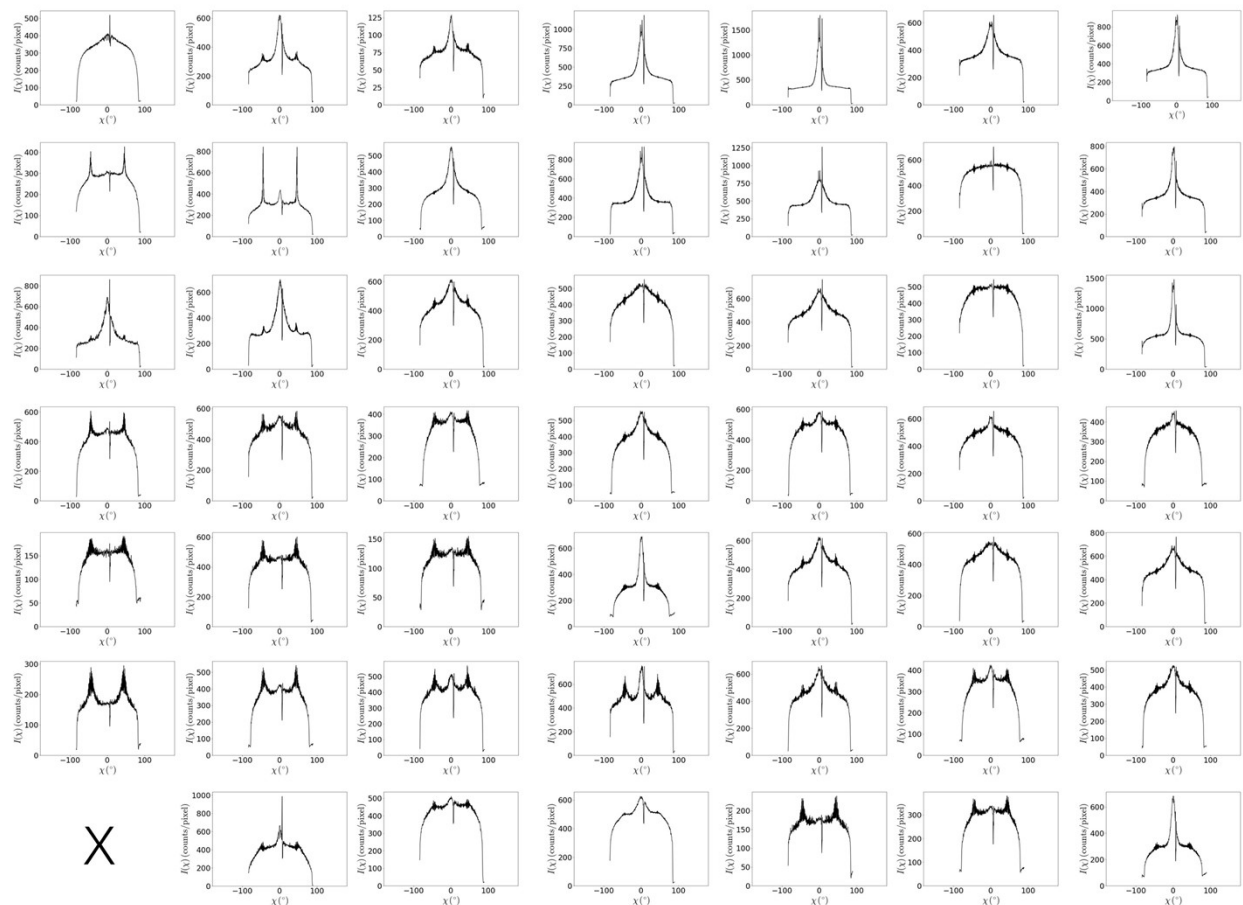
288

289 **Note 1: Additional discussions on texture:**

290 The texture of perovskite crystallites within a film have been proposed to affect the charge carrier
291 transport and collection properties across the layer, thus influencing the overall photovoltaic
292 performance of the solar cell stack ^{14,15}. XRD highlighted a high level of texturing in our films, as
293 mentioned above, resulting in prominent peaks at $\sim 14^\circ$ and $\sim 28^\circ$. The reason for the orientation
294 can be traced back to different surface energies for different planes ¹⁴. The texture of a film can
295 also be monitored looking at the GIWAXS diffraction patterns, tracking the intensity variations of
296 a specific plane with varying azimuthal angle. The ring at $q = 1.035 \pm 0.09 \text{ \AA}$ in GIWAXS
297 corresponds to the peak at $\sim 14^\circ$ in XRD, and is assigned to the $(100)_C$ plane for a cubic structure,
298 to the $(110)_T$, $(001)_T$ reflections for the tetragonal phase, and to the $(110)_O$, $(002)_O$ reflections for
299 the orthorhombic phase. The peak is integrated and plotted as function of the azimuthal angle as
300 detailed in Fig. S26. Fig. S27 summarizes the preferred facet orientation, as extracted from Fig. S26.
301 All the 26 I-rich compositions show preferred facet orientation of crystal plane at peak $q = 1.035$
302 $\pm 0.09 \text{ \AA}$, parallel to the substrate (0°) or/and stacked along the 45° direction with respect to the
303 substrate (Fig. S26-S27). This indicates the ordered and long-range-orientated crystallization of
304 the nineteen I-rich films, which is beneficial for charge transport within the perovskite films, and
305 consequently device performance ¹⁶. In addition, Fig. S28 schematically demonstrates $(100)_C$,
306 $(110)_T$, $(001)_T$, $(110)_O$, and $(002)_O$ crystal planes, involve only one kind of stacking layer,
307 composed of Pb and X atoms without A, regardless of the crystal phase. Therefore, the nineteen
308 I-rich films are long-range-orientated and have same atomic stacking structure for the plane at
309 peak $q = 1.035 \pm 0.09 \text{ \AA}$. The azimuthal angle integration also shows the perovskite films for Cs-
310 rich compositions are more oriented at 0° than that of FA-rich samples, suggesting higher texture
311 of Cs-rich compositions, as illustrated in Fig. S26. However, the PCE of Cs-rich compositions is
312 significantly lower than that of FA-rich compositions. Hence, the impact of the texture on the
313 device performance is negligible in the face of other factors, e.g., bandgap, surface morphology,
314 and the existence of δ -phase.

315

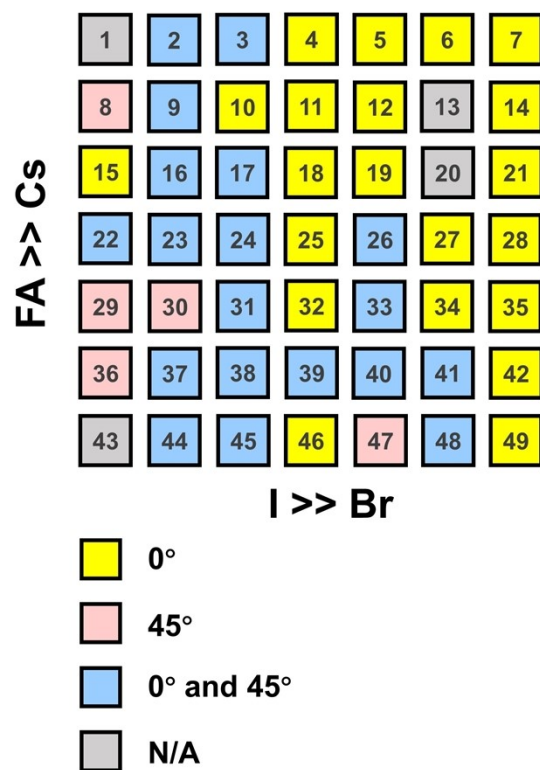
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318 Fig. S26 Integrated azimuthal angle along the ring q ($q = 1.035 \pm 0.09 \text{ \AA}$) at an incident angle of
 319 0.5° corresponding to the location $\sim 14^\circ$ in XRD, assigned to (100) plane for a cubic structure,
 320 which plane splits into the (110), (001) reflections and (110), (002) reflections in the tetragonal
 321 phase and orthorhombic phase, respectively¹⁷.

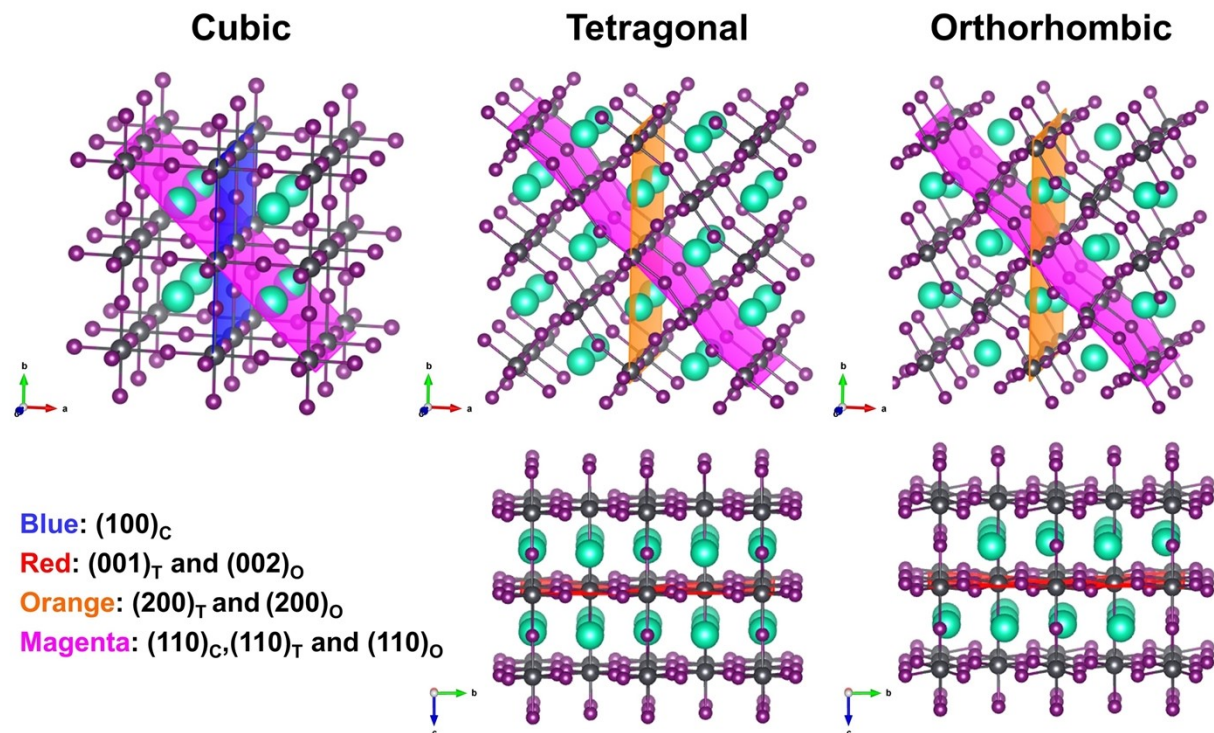
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324 Fig. S27 Color map of the preferential integrated azimuthal angle along the ring q ($q = 1.035 \pm$
 325 0.09 \AA) to summarize Fig. S25.

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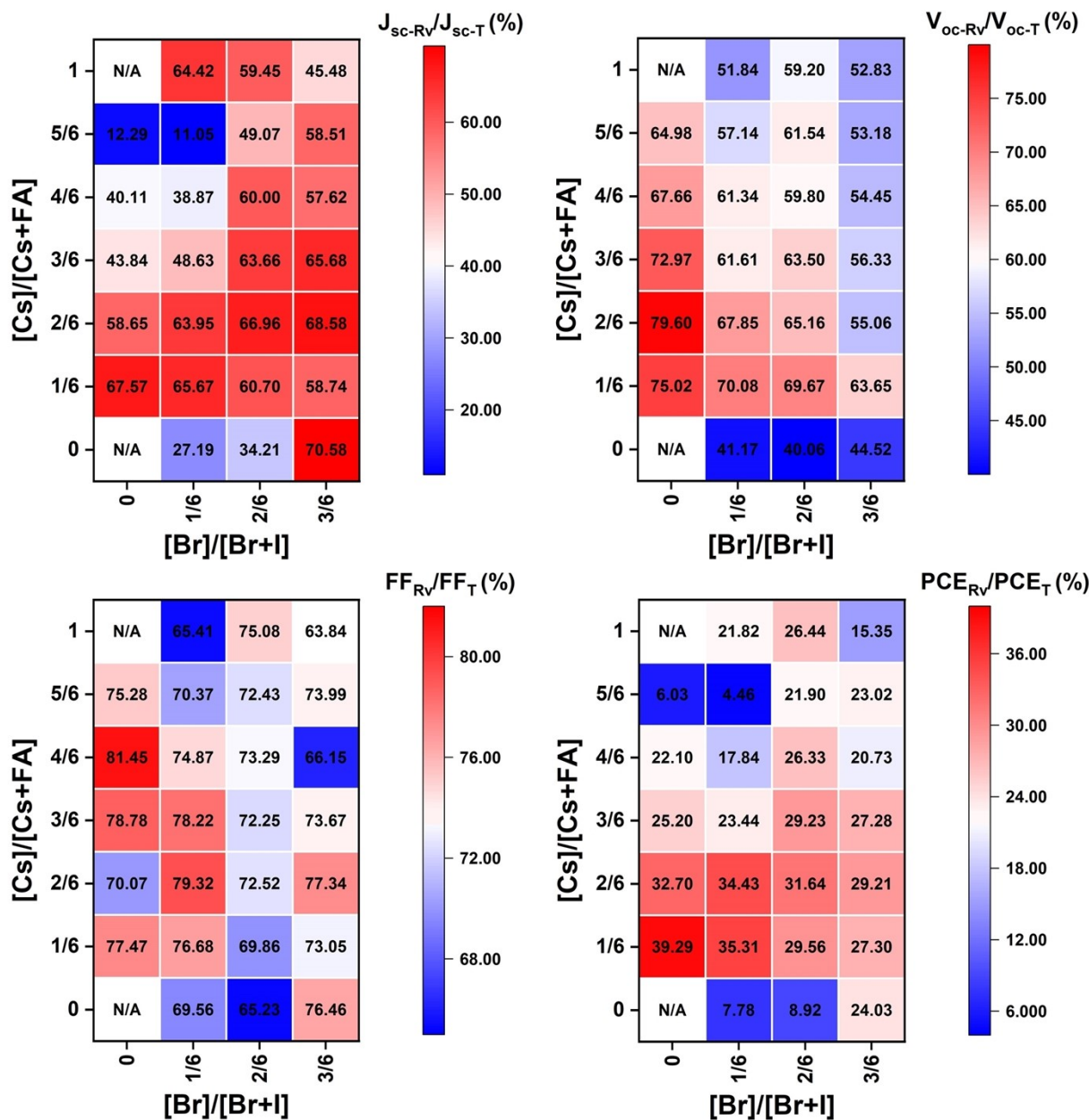
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328 Fig. S28 Crystal structure models of three perovskite phases, including cubic, tetragonal and
 329 orthorhombic structure for different planes.

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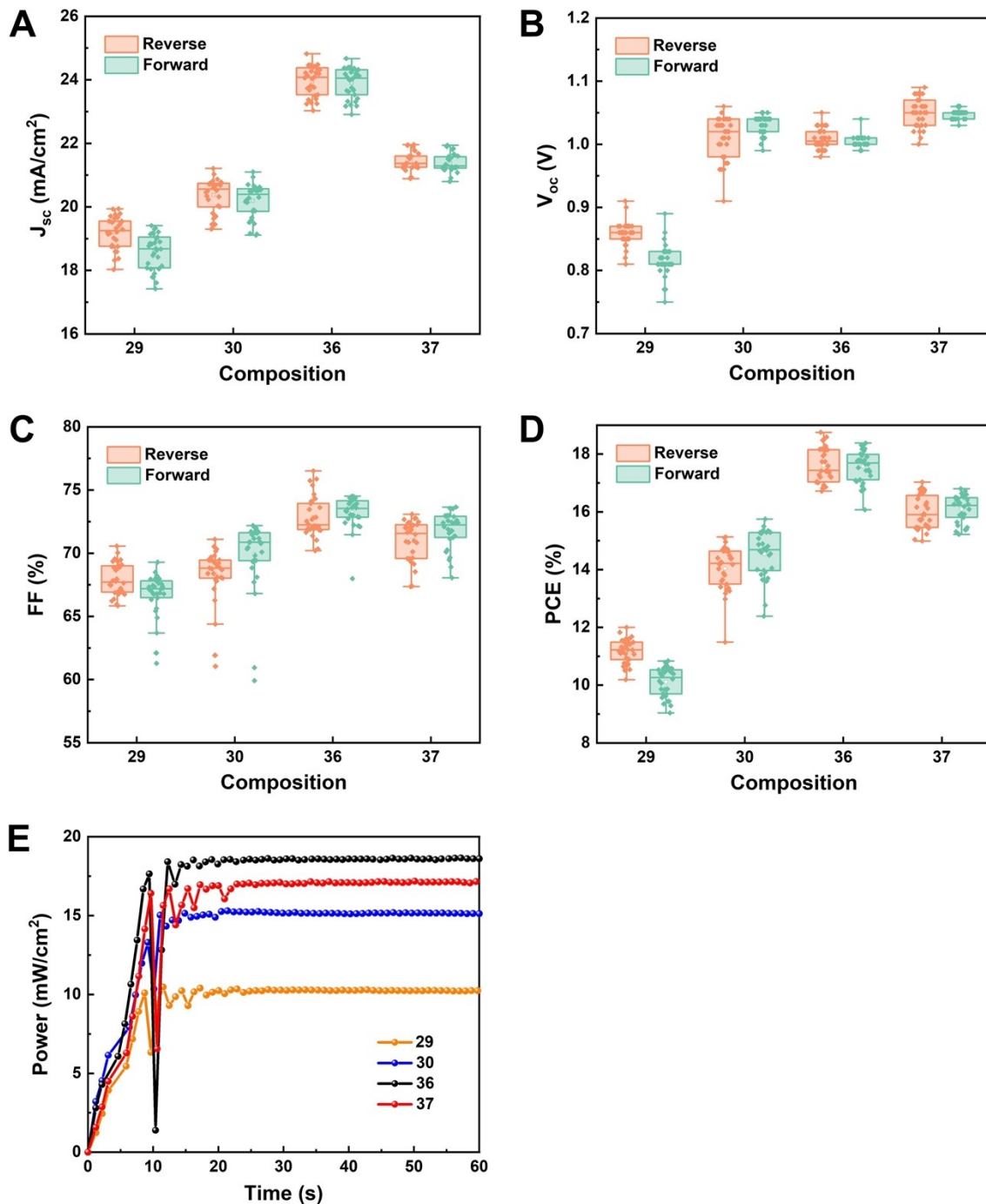
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334 Fig.S29 The ratio of photovoltaic parameters of J_{sc} , V_{oc} , FF, and PCE for reverse scan from best
 335 devices of I-rich compositions using 0.4 M perovskite precursor solutions to the corresponding
 336 theoretical Shockley-Queisser limit¹⁸.

337

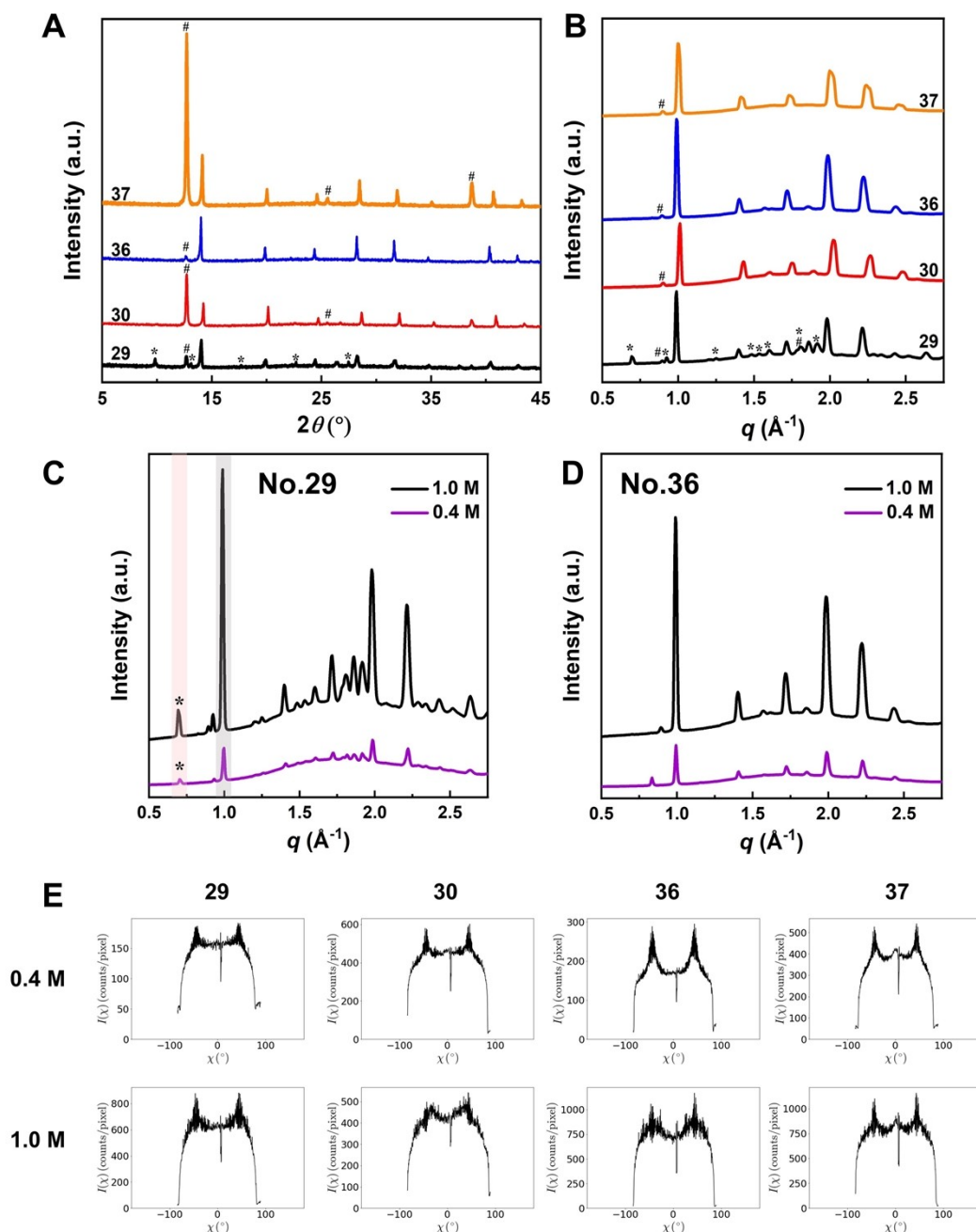


338

339 Fig.S30 Box plots for photovoltaic parameters of J_{sc} (A), V_{oc} (B), FF (C), and PCE (D) using 1.0
 340 M perovskite precursor solutions for composition No 29, No.30, No.36, and No37. (E) The MPP
 341 tracking for the champion devices.

342

343



345

Fig.S31 XRD (A) and circular average of GIWAXS (B) patterns for composition No.29, No.30, No.36, and No.37 using 1.0 M perovskite precursor solutions along with the 150 $^{\circ}\text{C}$ annealing. The # and * denote δ -CsPbI₃ and excess PbI₂, respectively. The GIWAXS comparison of the films with composition No.29 (C) and No. 36 (D) deposited by 0.4 M precursors solution annealing at 65 $^{\circ}\text{C}$ and by 1.0 M precursors solution annealing at 150 $^{\circ}\text{C}$. (E) Integrated azimuthal angle for

351 composition No.29, No.30, No.36, and No37, deposited by 0.4 M precursors solution annealing at
352 65°C and by 1.0 M precursors solution annealing at 150°C, along the ring q ($q = 1.035 \pm 0.09$ Å)
353 at an incident angle of 0.5° corresponding to the location $\sim 14^\circ$ in XRD, assigned to (100) plane
354 for a cubic structure, which plane splits into the (110), (001) reflections in the tetragonal phase.
355

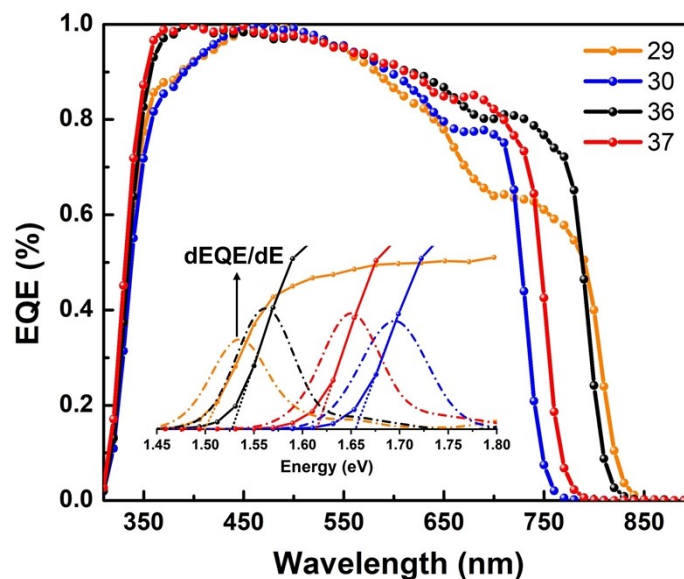
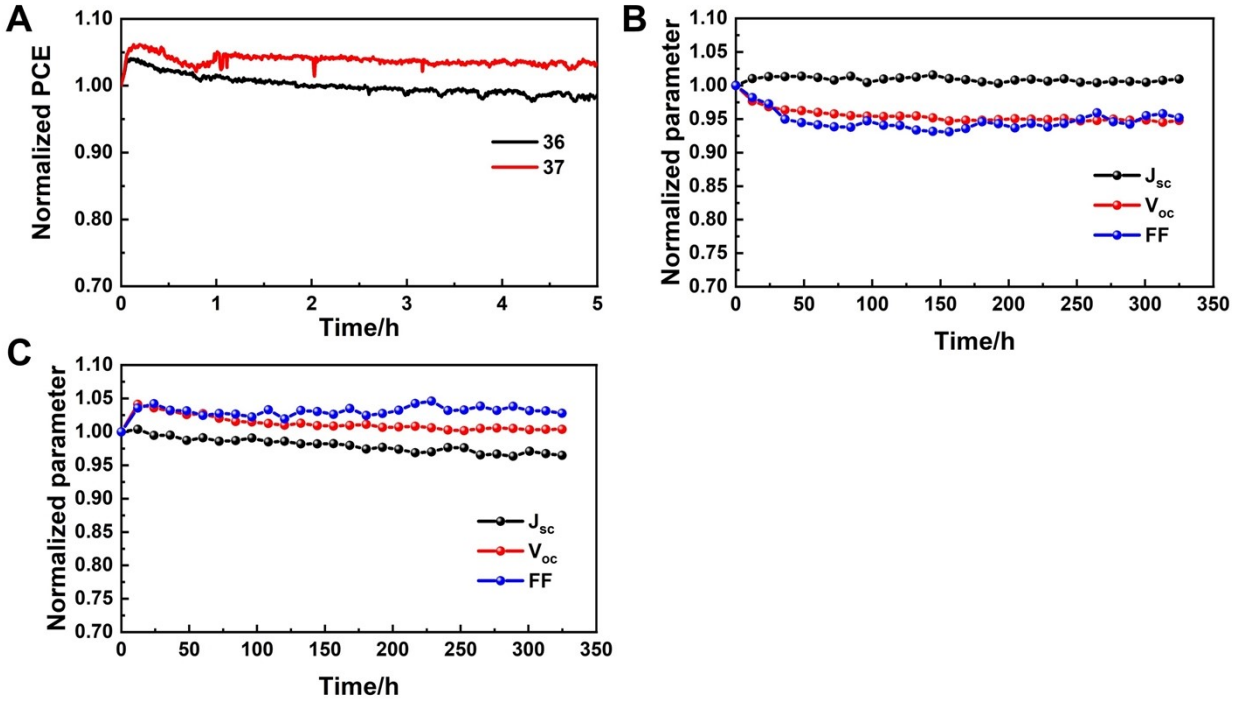


Fig. S32 EQE spectra of the perovskite solar cell for No.29, No.30, No.36, and No.37. The insert showed the photovoltaic bandgap extracted from extrapolation to 0 of the EQE curve or the inflection point of the derivative of the EQE curve.

Table S2. Optical bandgap extracted from absorbance (E_g) and PL (E_{PL}) as well as the photovoltaic bandgap extracted from EQE spectra measured by solar cells using the extrapolation to 0 of the EQE curve (E_0) or the inflection point of the derivative of the EQE curve (E_{IP}).

Composition	29	30	36	37
E_g /eV	1.49	1.65	1.54	1.62
E_{PL} /eV	1.54	1.67	1.58	1.65
E_0 /eV	1.50	1.65	1.53	1.61
E_{IP} /eV	1.53	1.70	1.56	1.65

Note: For the mismatch between optical bandgap and photovoltaic bandgap, the optical bandgap determined by the Tauc method represents an internal property of the photovoltaic material while the photovoltaic bandgap determined by the EQE represents an external property of the solar cells. In practice, the device structure and fabrication methods may modify the optoelectronic properties of the device, producing a mismatch between the optical value and photovoltaic bandgap^{19,20}.



371

372 Fig.S33 (A) Continuous MPP tracking results under 1 sun irradiance at N_2 atmosphere of devices
 373 of composition 36 and 37 for first 5 hours, which was zoomed from the Fig 10 F in the main text.

374 The evolution of the normalized V_{oc} , J_{sc} , and FF of PSCs for composition No.36 (B) and No.37
 375 (C) that aged under at MPP with 1 sun irradiation in N_2 atmosphere at room temperature.

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