

Intermediate-spin iron(IV)-oxido species with record reactivity

Supporting Information

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Supporting Information

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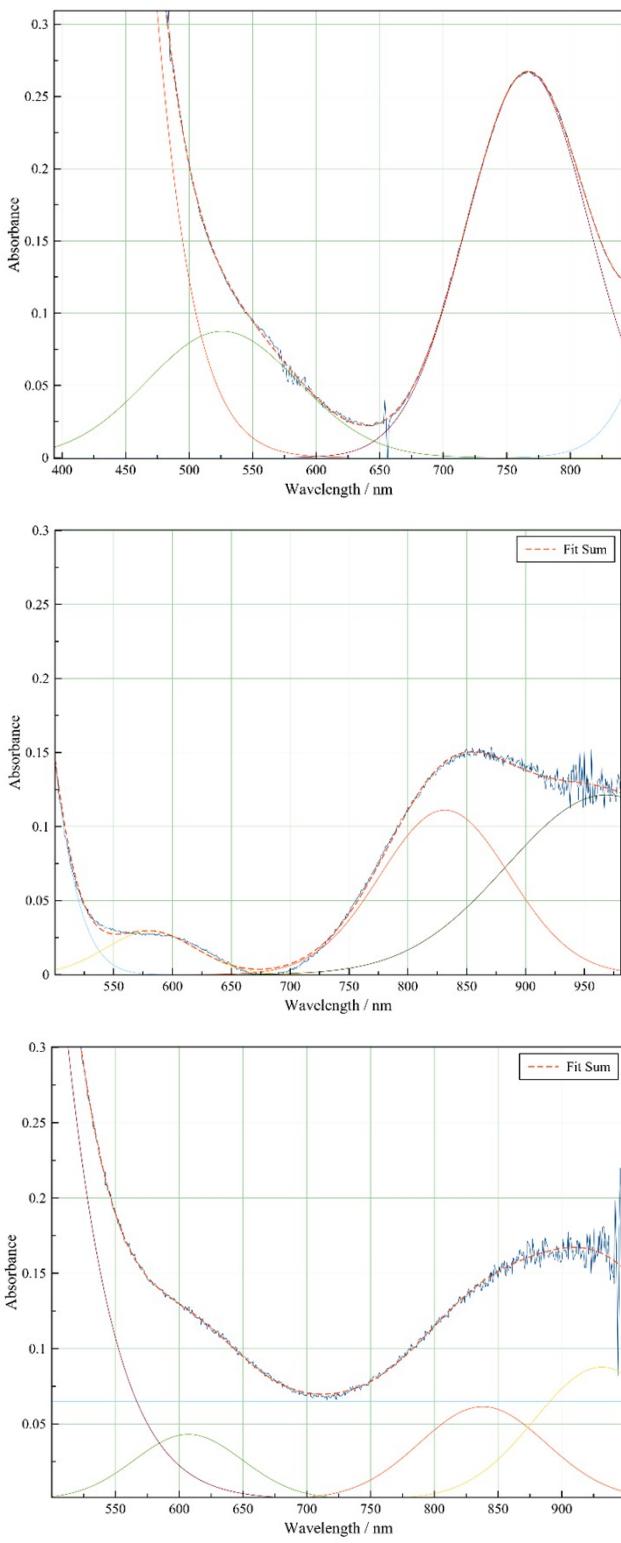
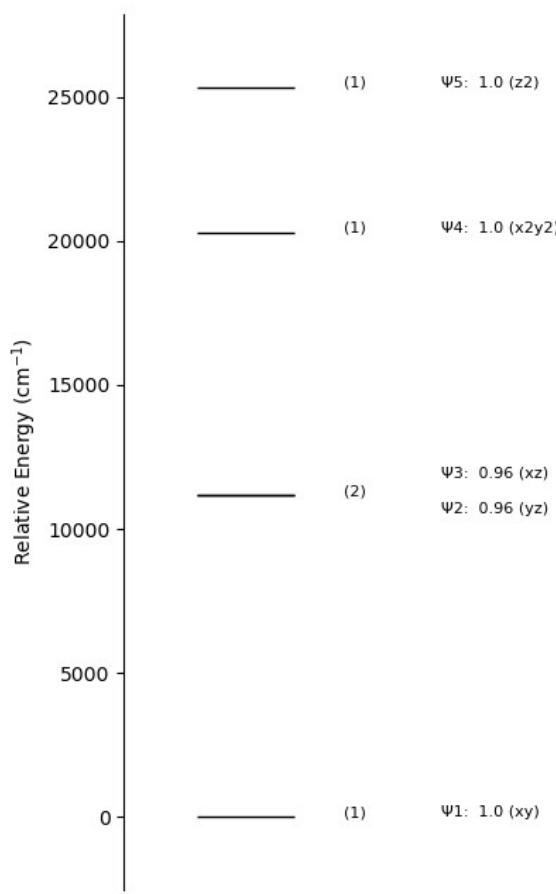


Figure S1. Gaussian deconvolution of the UV-vis spectra of $[(L^1)\text{Fe}^{IV}=\text{O}(X)]^{n+}$ (from top to bottom: $X = \text{MeCN}, \text{Cl}^-, \text{Br}^-$; note that the resolution of the experimental spectra (cryo-stopped-flow) is limited; in the main text we have adopted the values determined here except for the Br^- complex, where the low energy transition was set to 900 nm).

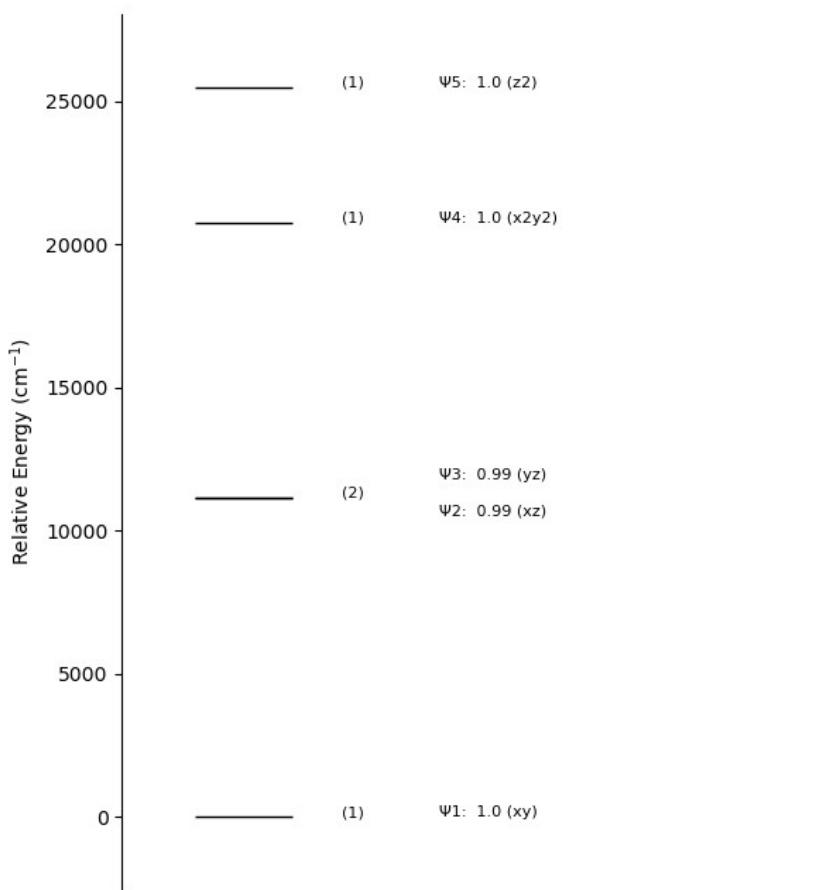
The axial AOM parameters from $[(L^2)Fe^{IV}=O(MeCN)]^{2+}$ were transferred to the octahedral geometry and were fixed. The Racah C/B ratio was fixed at 5.0. The equatorial e_{σ} values were varied along with the Racah B parameter.

A value of Racah B at ~ 600 cm $^{-1}$ was enough to correlate the band gap with the e_{σ} (equatorial) parameter. It was found that the average e_{σ} decreased going from MeCN > Cl $^-$ > Br $^-$. The shoulders were positioned approximately within range of the deconvoluted band (although there is a larger error on this).

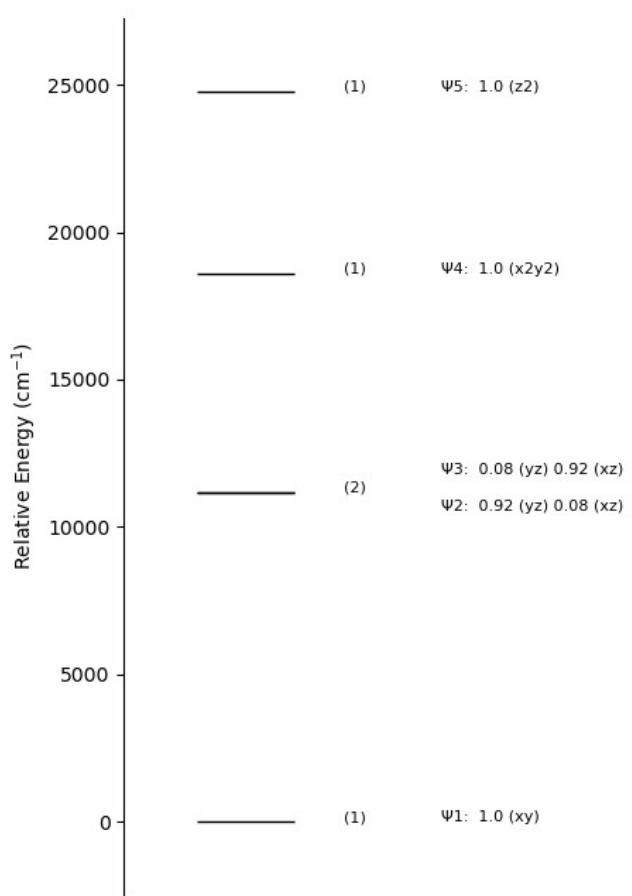
Orbital diagrams



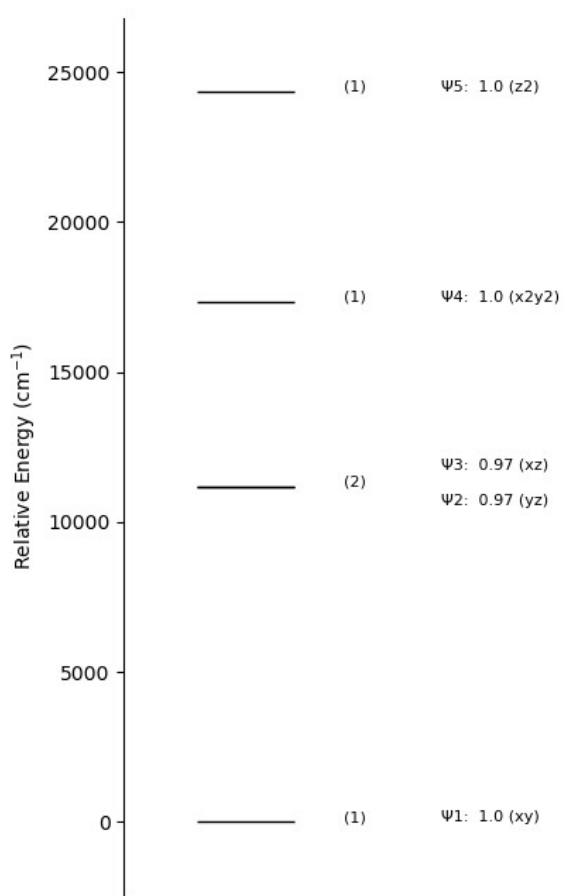
Scheme S1. L² / MeCN computed orbital energies.



Scheme S2. L^1 / MeCN computed orbital energies.



Scheme S3. L^1 / Cl^- computed orbital energies.



Scheme S4. L^1 / Br^- computed orbital energies.

