

Supporting Information:

Kinetics of the reaction of OH with methyl nitrate (232–343 K)

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1 Assessment of impurities in the CH₃ONO₂ sample

The concentrations of potential impurities HNO₃ and CH₃OH were assessed using FTIR spectroscopy. The spectrum of CH₃ONO₂ along with vibrational assignments is displayed below. The insets show absorption features of HNO₃ and CH₃OH from reference spectra obtained using the same apparatus. The commercial methanol sample was purified by freeze-pump-thaw cycles and HNO₃ was synthesised by mixing potassium nitrate and 98 % sulphuric acid. As described in the manuscript, neither HNO₃ nor CH₃OH were detected, enabling upper limits to their concentrations to be derived.

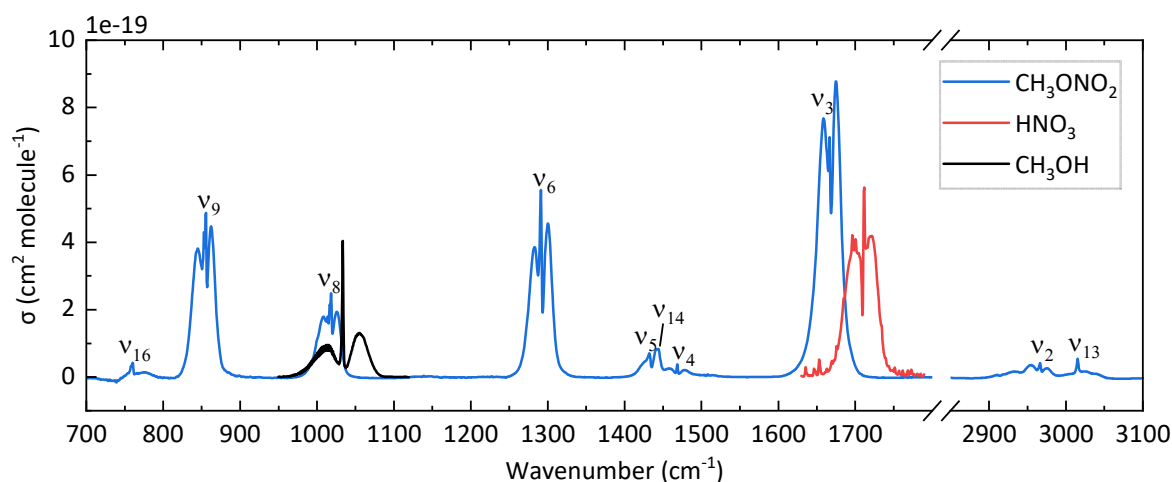


Figure S1: Gas-phase FTIR spectrum of methyl nitrate (blue) in the range 700-3100 cm⁻¹. The vibrational frequencies are assigned according to the literature⁴. Absorption features of HNO₃ (1.27×10^{15} molecule cm⁻³, scaled by factor 0.5, red) and CH₃OH (1.14×10^{15} molecule cm⁻³, scaled by factor 0.025, black), respectively. The concentrations of CH₃OH and HNO₃ were derived from simultaneously measured optical extinction at 185 nm.

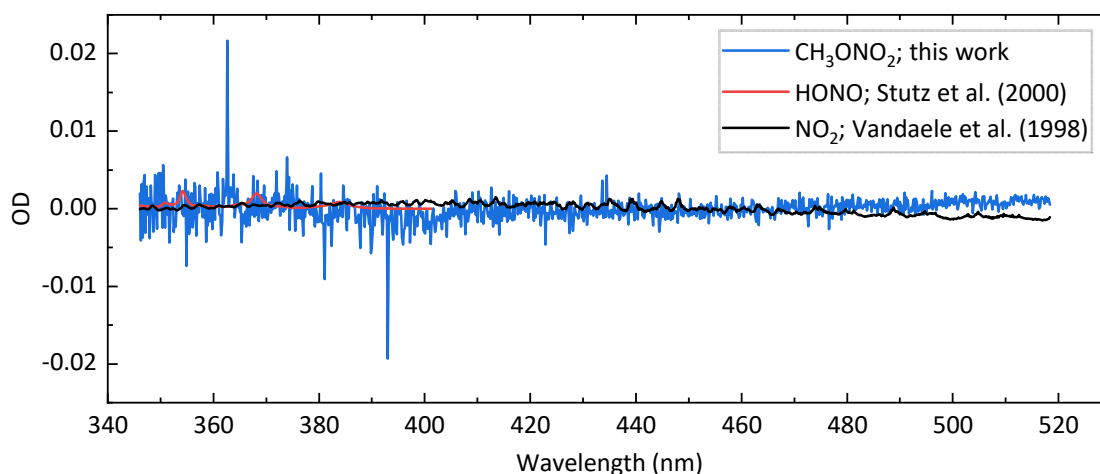


Figure S2: Optical density (blue) of a flowing sample of CH₃ONO₂ (1.86×10^{16} molecule cm⁻³) in regions where HONO and NO₂ absorb. The red and black lines are calculated optical densities, based on measurements by Stutz et al.³ and Vandaele et al.⁵, for HONO (5×10^{12} molecule cm⁻³) and NO₂ (5×10^{12} molecule cm⁻³).

2 Use of CH₃ONO₂ photolysis to generate OH

As described in the manuscript, the 248 nm photolysis of CH₃ONO₂ to generate OH results in bi-exponential OH profiles (Figure 5), the analysis of which does not lead to the same rate coefficient k_4 as derived from experiments in which HONO was photolyzed at 351 nm to generate OH. This behaviour contradicts the results of Talukdar et al.⁷, who were able to derive a reasonably accurate rate coefficient with this method. We identify two reasons why, in our study, the photolysis of CH₃ONO₂ does not lead to the correct value of k_4 .

1) Co-detection of OH and CH₃O

A potential explanation for a bias in k_4 is the undesired co-detection of CH₃O as well as OH. As CH₃O is formed directly at high yield whereas OH is formed after conversion of an H-atom (itself formed at low yield) to OH via reaction with CH₃ONO₂, CH₃O is expected to be present at a much larger concentration than OH. Figure S3 (green trace) shows the LIF-excitation spectra obtained at discrete steps when CH₃ONO₂ is photolyzed at 248 nm. Clearly, the excitation spectrum obtained is a composite of OH and CH₃O, which will preclude the determination of k_4 based on the assumption that the LIF signal obtained with excitation at 282 nm is due to OH only.

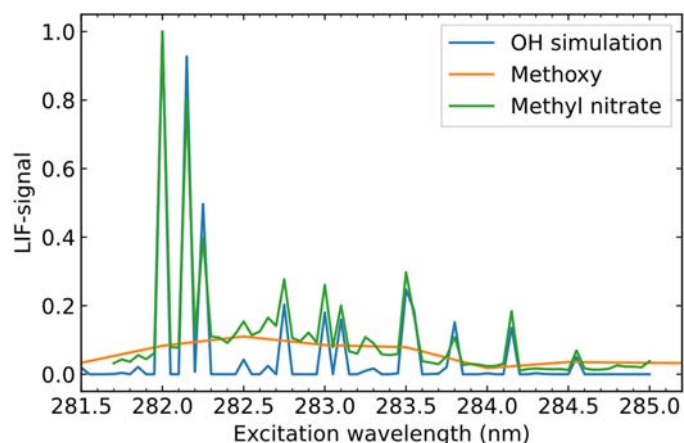


Figure S3: Green-trace: The fluorescence excitation spectrum (measured at discrete steps of 0.05 nm) observed when photolyzing methyl nitrate at 248 nm. Blue-trace: The simulated fluorescence excitation spectrum of OH (LIFBASE 2.1²) with a resolution of 0.125 Å plotted at the same discrete steps. Orange-trace: fluorescence excitation spectrum of CH₃O with a step size of 0.5 nm, obtained by photolyzing CH₃ONO at 248 nm.

2) Reaction of OH with CH₃O and NO₂ and further secondary chemistry

The CH₃O and NO₂ photo-fragments formed when CH₃ONO₂ is photolyzed at 248 nm is expected to result in a dependence of the OH decay constant on the energy of the photolysis laser. This was explored using numerical simulation Facsimile¹ of an assumed reaction scheme (see below). Figure S4 shows that taking into account the side reactions of OH with CH₃O, NO₂ and CH₂O results in a deviation (a bias to larger values) compared to the expected rate constant at various photolysis energies (up to 10mJ / pulse).

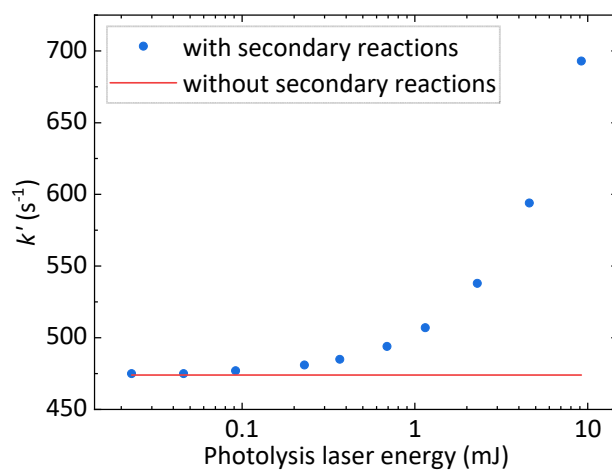


Figure S4: First-order decay coefficients k' plotted against photolysis laser energy. In blue, k' determined via numerical simulation¹, accounting for secondary reactions of OH with CH₃O, NO₂, and CH₂O. The reactions considered in the simulation are shown in Table S1. In red, OH decay constant, calculated from the CH₃ONO₂ concentration used in the simulation and the rate coefficient determined in this work (i.e. without secondary losses of OH).

Table S1: Reactions considered in the simulation. Rate coefficients correspond to 298 K and 100 Torr N₂.

Reaction	Rate coefficient	source
$\text{H}_2\text{O}_2 + h\nu \rightarrow \text{OH} + \text{OH}$	a	IUPAC ⁸
$\text{CH}_3\text{ONO}_2 + h\nu \rightarrow \text{CH}_3\text{O} + \text{NO}_2$	a	Talukdar et al. ⁷
$\text{CH}_3\text{ONO}_2 + h\nu \rightarrow \text{H} + \text{CH}_2\text{O} + \text{NO}_2$	a	Talukdar et al. ⁷
$\text{CH}_3\text{O} + \text{NO}_2 + \text{M} \rightarrow \text{CH}_3\text{ONO}_2 + \text{M}$	b1	IUPAC ⁸
$\text{OH} + \text{CH}_3\text{ONO}_2 \rightarrow \text{H}_2\text{O} + \text{CH}_2\text{O} + \text{NO}_2$	2.3e-13	This work
$\text{OH} + \text{NO}_2 + \text{M} \rightarrow \text{HNO}_3 + \text{M}$	b2	IUPAC ⁸
$\text{OH} + \text{H}_2\text{O}_2 \rightarrow \text{HO}_2 + \text{H}_2\text{O}$	3.5e-12	IUPAC ⁸
$\text{OH} + \text{HO}_2 \rightarrow \text{H}_2\text{O} + \text{O}_2$	1.1e-10	IUPAC ⁸
$\text{OH} + \text{CH}_3\text{O} \rightarrow \text{CH}_2\text{O} + \text{H}_2\text{O}$	3.0e-11	Tsang and Hampson ⁹
$\text{OH} + \text{CH}_2\text{O} \rightarrow \text{H}_2\text{O} + \text{HCO}$	8.5e-12	IUPAC ⁸
$\text{HO}_2 + \text{HO}_2 \rightarrow \text{H}_2\text{O}_2 + \text{O}_2$	1.7e-13	IUPAC ⁸
$\text{H} + \text{CH}_3\text{ONO}_2 \rightarrow \text{OH} + \text{CH}_3\text{ONO}$	1.6e-13	Talukdar et al. ⁷
$\text{H} + \text{NO}_2 \rightarrow \text{OH} + \text{NO}$	1.3e-10	Burkholder et al. ¹⁰
$\text{HO}_2 + \text{CH}_3\text{O} \rightarrow \text{CH}_3\text{OH} + \text{O}_2$	4.7e-11	Assaf et al. ¹¹
$\text{NO} + \text{HO}_2 \rightarrow \text{OH} + \text{NO}_2$	8.5e-12	IUPAC ⁸

Listed rate coefficients are in cm³ molecule⁻¹ s⁻¹. ^aRelative concentrations of OH, H and CH₃O + NO₂ were based on known cross-sections of H₂O₂ and CH₃ONO₂ and photolysis quantum yields at 248 nm. ^{b1, b2}Parameterisation with “Troe” expression using parameters listed by IUPAC to calculate the bimolecular rate coefficient $k(p, T)$. b1: $k_0 = 8.1 \times 10^{-29} (\text{T}/300)^{4.5} [\text{N}_2] \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$, $k_{\text{inf}} = 2.1 \times 10^{-11} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$, $F_c = 0.44$. b2: $k_0 = 3.2 \times 10^{-30} (\text{T}/300)^{4.5} [\text{N}_2] \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$, $k_{\text{inf}} = 3.0 \times 10^{-11} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$, $F_c = 0.41$. The “Troe” expression is:

$$k(p, T) = \frac{k_0 \left(\frac{T}{300}\right)^{-m} M k_{\text{inf}} \left(\frac{T}{300}\right)^{-n}}{k_0 \left(\frac{T}{300}\right)^{-m} M + k_{\text{inf}} \left(\frac{T}{300}\right)^{-n}} \log F$$

where

$$\log F = \frac{\log F_c}{1 + \left[\log \left(\frac{k_0 \left(\frac{T}{300}\right)^{-m} M}{k_{\text{inf}} \left(\frac{T}{300}\right)^{-n}} \right) / N \right]^2}$$

3 Atmospheric lifetime of CH₃ONO₂

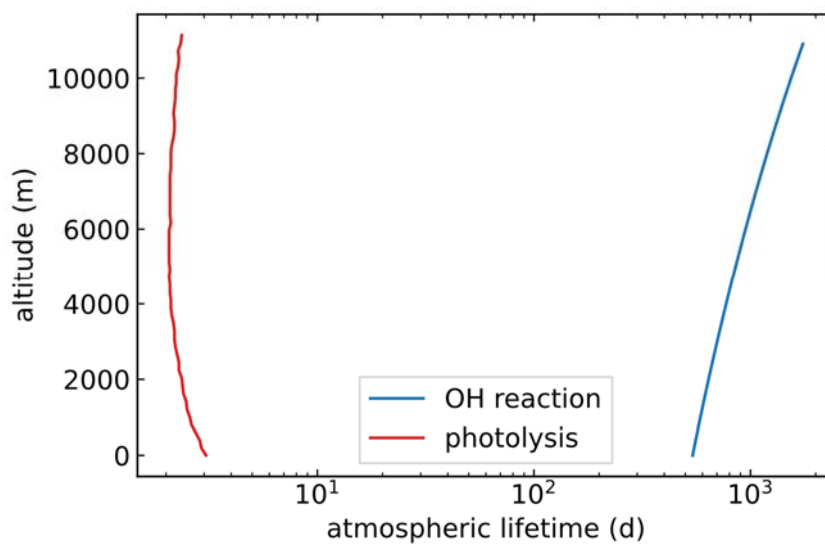


Figure S5: Altitude dependent atmospheric lifetimes (in days) of CH₃ONO₂ with respect to reaction with OH (blue, assuming a global mean concentration of 1×10^6 , molecule cm⁻³ 6). In red, the lifetime considering the photolysis, based on J-values from Talukdar et al.⁷.

4 Calculation of uncertainties in the rate coefficient

The potential systematic bias of 5 % is an estimation based on the multiple determination of the absorption cross-section (Figure 1). To determine the absorption cross-section, a weighted linear fit of OD versus concentration is used and the resulting error is given by 2σ . For each point in Figure 1, the error for the optical density (ΔOD) is derived by averaging the measured OD and calculating the standard deviation of 2σ . The error in the x-direction, $\Delta(c \cdot l)$ is determined by an error propagation of the following (estimated) errors $\Delta T = 1$ K, $\Delta p = 0$ Torr (better than 1%) and $\Delta c\% = 2\%$ ($c \times 0.02$). Following the equation for the error propagation:

$$\Delta(c \cdot l) = \pm 3.24 \times 10^{16} \times 298\text{K} \times \sqrt{\left(\frac{l p}{T} \Delta c\%\right)^2 + \left(\frac{l c\%}{T} \Delta p\right)^2 + \left(-\frac{l c\% p}{T^2} \Delta T\right)^2 + \left(\frac{c\% p}{T} \Delta l\right)^2}$$

The following fit equations are obtained from the respective weighted fit (orthogonal distance regression) for the two individual measurement series:

$$y = [(1.69 \pm 0.07) \times 10^{-18}]x + (0.004 \pm 0.009) \text{ and } y = [(1.67 \pm 0.14) \times 10^{-18}]x + (-0.012 \pm 0.031),$$

where the slope corresponds to the absorption cross-section. The differences here result in a deviation of 1.2%. Both measurements together result in the following fit equation: $y = [(1.65 \pm 0.09) \times 10^{-18}]x + (0.002 \pm 0.015)$. By this measurement, we estimate the total error of the experiment to be 5%. Added to this is the statistical error from the various fits. A monoexponential fit (least squares) was used to determine k' and the standard deviation (2σ) was obtained with the value. These can be seen as error bars in Figure 3. Here, a weighted linear fit (least squares) was performed to determine k , which again outputs an error (2σ). This in turn can be seen as an error in the Arrhenius plot, where a weighted monoexponential fit (least squares) was also performed. As is often the case in such experiments, the scatter in the data is greater than the total uncertainty estimated as described above, which may indicate unidentified (and thus not assessable) sources of uncertainty.

References

1. A. R. Curtis and W. P. Sweetenham, *Facsimile, Atomic Energy Research Establishment, Report R-12805, 1987, 1987.*
2. J. Luque and D. R. Crosley, *LIFBASE (version 1.5) <http://www.sri.com/cem/lifbase>, 1999.*
3. J. Stutz, E. S. Kim, U. Platt, P. Bruno, C. Perrino and A. Febo, *Journal of Geophysical Research: Atmospheres*, 2000, **105**, 14585.
4. J. F. Stanton, B. A. Flowers, D. A. Matthews, A. F. Ware and G. B. Ellison, *J. Mol. Spectrosc.*, 2008, **251**, 384.
5. A. C. Vandaele, C. Hermans, P. C. Simon, M. Carleer, R. Colin, S. Fally, M. F. Merienne, A. Jenouvrier and B. Coquart, *J. Quant. Spectrosc. Radiat. Transfer*, 1998, **59**, 171.
6. J. Lelieveld, S. Gromov, A. Pozzer and D. Taraborrelli, *Atmos. Chem. Phys.*, 2016, **16**, 12477.
7. R. K. Talukdar, J. B. Burkholder, M. Hunter, M. K. Gilles, J. M. Roberts and A. R. Ravishankara, *J. Chem. Soc., Faraday Trans.*, 1997, **93**, 2797.
8. IUPAC, Task Group on Atmospheric Chemical Kinetic Data Evaluation, (Ammann, M., Cox, R.A., Crowley, J.N., Herrmann, H., Jenkin, M.E., McNeill, V.F., Mellouki, A., Rossi, M. J., Troe, J. and Wallington, T. J.). Last access Sept. 2023, <https://iupac.aeris-data.fr/>).
9. W. Tsang and R. F. Hampson, *J. Phys. Chem. Ref. Data*, 1986, **15**, 1087.
10. J. Burkholder, S. Sander, J. Abbatt, J. Barker, C. Cappa, J. Crouse, T. Dibble, R. Huie, C. Kolb and M. Kurylo, *Chemical kinetics and photochemical data for use in atmospheric studies; evaluation number 19*, Jet Propulsion Laboratory, National Aeronautics and Space Administration, 2020.
11. E. Assaf, C. Schoemaeker, L. Vereecken and C. Fittschen, *Phys. Chem. Chem. Phys.*, 2018, **20**, 10660.