Supporting Information

Theoretical design of homojunction solar cells based on chalcopyrite AgInSe₂: a combined study of first-principles calculations and device simulations

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S1. The AgInSe₂ crystal structures with different space groups



Figure S1. The crystal structures of AgInSe₂ are presented for different space groups: I⁴2d (a), Pna2₁ (b), R3m (c), I4₁/amd (d), Fdd2 (e), and P4/mmm (f). Here, the blue, orange, and yellow spheres represent the Ag, In, and Se atoms, respectively.

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We have calculated the total energies of $AgInSe_2$ structures in different space groups (I⁴2d, Pna21, R3m, I41/amd, Fdd2, and P4/mmm) from the Materials Project crystal database,^[1] as shown in Figure S1. The lattice constants and total energies of the structures with the six space groups are listed in Table S1. The calculation results show that the structure in the $I^{\bar{4}}2d$ space group has the lowest total energy. In addition, AgInSe₂ with the $I^{\bar{4}}2d$ space group has been widely reported in experiment.^[2,3] Therefore, we adopt AgInSe₂ with the $I^{\bar{4}}2d$ space group to carry out subsequent calculations on electronic structure, optical and defect properties, as well as device simulation studies.

space group. The total energy of 1-20 is set as zero.						
	Lattice constants			$\Delta E_{\rm tot} ({\rm meV/f.u.})$		
Phase	<i>a</i> (Å)	<i>b</i> (Å)	<i>c</i> (Å)			
I ⁴ 2d	6.16	6.16	11.90	0		
Pna2 ₁	7.06	7.06	8.66	0.04		
R3m	2.16	2.16	21.01	1.59		
I4 ₁ /amd	5.65	5.65	11.41	2.24		
Fdd2	5.65	5.65	5.71	2.23		
P4/mmm	3.94	3.94	5.69	3.51		

Table S1. The calculated lattice constants and total energy differences (ΔE_{tot}) of AgInSe₂ in six space group. The total energy of I⁴2d is set as zero.

S2. Calculation details of defect formation energy

In this work, the used the following equation to calculate the defect formation energy⁴⁻⁶:

$$\Delta E(\alpha,q) = E(\alpha,q) - E(host) + \Sigma_i n_i (\Delta \mu_i + \mu_i) + q(E_v + E_F + \Delta V) + E_{corr}, \#(S.1)$$

where $E(\alpha,q)$ is the total energy of the supercell with a defect α in the charge state q, while E(host) is the corresponding energy for a defect-free system. Further, n_i is the number of atoms added into $(n_i = -1)$ or removed from $(n_i = +1)$ the supercell for the element i during the formation of the defect. $\Delta \mu_i$ is the relative chemical potential to correct the computed chemical potential μ_i of the elemental solid/gas. E_v is the valence band maximum (VBM) of the defect-free system, E_F is the Fermi level, referenced to the VBM, and ΔV is the potential correction between the perfect crystal and the defect system. In addition, we have accounted for the image charge correction term in our calculations. The image charge correction term is calculated by the following equation⁷:

$$E_{corr} = \frac{2}{3}n^2 \frac{E_{wald}}{\varepsilon_0} \#(S.2)$$

S3. Stability of AgInSe₂

Currently, there is no direct research on the stability of AgInSe₂ solar cells. Therefore, we derive relevant conclusions by investigating the more mature CuInSe₂, MAPbI₃, and FAPbI₃ systems. Analysis of the chemical potential regions of AgInSe₂, CuInSe₂, MAPbI₃, and FAPbI₃ via first-principles calculations indicates that the chemical potential region of AgInSe₂ is similar in size to that of CuInSe₂ and significantly larger than those of MAPbI₃ and FAPbI₃.⁸⁻¹⁰ This implies that the stability of AgInSe₂ is comparable to that of CuInSe₂ and superior to those of MAPbI₃ and FAPbI₃. In addition, both AgInSe₂ and CuInSe₂ adopt a diamond-like structure, which is associated with notable thermodynamic and structural stability.^{11,12} Given that CuInSe₂ solar cells have demonstrated reliable operational stability in long-term tests,^{13,14} it is reasonable to infer that AgInSe₂ solar cells would exhibit comparable stability under practical conditions.

S4. Regulation of chemical potentials for AgInSe₂ preparation

We deeply appreciate your constructive suggestions, as the perspective you proposed has significantly enhanced the experimental guidance value of the manuscript. By integrating the mature experience from systems such as $Cu_2ZnSnSe_4$ (CZTSe) and $Cu(In,Ga)(S,Se)_2$ (CIGSSe), we elaborate on the experimental implementation mechanism of combining chemical potential control with annealing in AgInSe₂ from the perspective of synergistic regulation of preparation atmosphere and annealing process.

Chen et al.'s¹⁵ computational investigations on CZTSe reveal that under Cu-poor and Zn-rich conditions (Cu/(Zn+Sn) \approx 0.8, Zn/Sn \approx 1.2), the concentration of the detrimental defect cluster [2Cu_{Zn}+Sn_{Zn}] decreases significantly. Under such conditions, Cu vacancies (V_{Cu}) emerge as the dominant acceptor defects, capable of providing moderate hole concentrations while mitigating excessive carrier recombination. Li et al.¹⁶ further validated this theoretical prediction experimentally: they first deposited a Cu/Zn/Sn metallic precursor by sputtering (controlling Cu/(Zn+Sn) = 0.70-0.75 and Zn/Sn = 1.10-1.15), followed by pre-alloying at 250 °C, soft selenization pretreatment at 280 °C, and finally annealing at 550 °C. This process promoted grain growth and improved the open - circuit voltage, increasing the photoelectric conversion efficiency (PCE) of CZTSe solar cells from 9.2% to 12.5%. This improvement validates that the combination of "precursor ratio + atmosphere + annealing" serves as an effective approach to integrate chemical potential control with annealing.

For the AgInSe₂ system, the research by Panda et al.¹⁷ provides direct references for its preparation. They prepared n-type AgInSe₂ films by adjusting the Ag/In target ratio, confirming that regulating the target composition is a direct means of chemical potential control. Meanwhile, the initial phase was formed via selenization at 250 °C, followed by high-temperature annealing at 500 °C to promote the decomposition of impurity phases and element diffusion, ultimately optimizing the crystallinity and optoelectronic properties of n-AgInSe₂ films. Similarly, p-AgInSe₂ can be prepared by adjusting the target ratio. In specific experimental operations, the chemical potential regulation of AgInSe₂ can be achieved through two core methods: first, adjusting the Ag/In ratio during sputtering to correspond to Ag-rich or Ag-poor growth points; second, regulating the Se partial pressure during the soft selenization stage to create a Se-rich environment. These methods can meet the chemical potential requirements for n-type (Point E: Ag-rich, In-poor, Se-rich) and p-type (Point C: Ag-poor, In-poor, Se-rich) materials. Meanwhile, the annealing temperature should be controlled at 600–800 K to promote crystallization and suppress the generation of harmful defects.

In summary, the experimental framework of "chemical potential regulation via atmosphere control and equilibrium promotion via annealing" in CZTSe, integrated with the specific methodologies of Panda et al. in AgInSe₂, establishes a comprehensive paradigm for combining chemical potential engineering with annealing processes. Additionally, Yang et al.¹⁸ employed DC sputtering to form a Cu-deficient ordered vacancy compound layer in the CIGSSe system, creating a high-performance buried homojunction with CIGSSe to fabricate homojunction structures. This achievement validates the feasibility of preparing AgInSe₂ homojunction solar cells via sputtering.

S5 Interface states of AgInSe₂ homojunction solar cell

The key factor influencing interface states is the interfacial defect concentration, as it directly determines the strength of interfacial recombination. The interfacial defect concentration is positively correlated with the degree of lattice mismatch.^{19,20} The advantage of homojunction solar cells over heterojunction solar cells lies in the lattice matching at the junction interface,^{21,22} so interfacial recombination at the junction can be neglected. The interfacial defect concentration between n-AgInSe₂ and ITO can be studied using the ElectronLatticeMatch (ELM) code,²³ and the resulting lattice mismatch is listed in Table S2. The lattice mismatch between the (001) plane of AgInSe₂ and the (001) plane of ITO is only 0.49%, outperforming the 2.70% and 2.10% mismatches at the interfaces of Cu₂ZnSnS₄/CdS and Cu(In,Ga)Se₂/CdS, respectively.⁴ Additionally, the lattice mismatch between the (100) plane of AgInSe₂ and the (110) plane of ITO is 3.35%, which is also within an acceptable range. This indicates that n-AgInSe₂ and ITO should exhibit low interface state density and low contact resistance.

Table S2. Lattice mismatch between the absorber and adjacent layers from the ELM screening.

absorber	adjacent layer	lattice mismatch (%)
AgInSe ₂ (001)	ITO (001)	0.49
AgInSe ₂ (100)	ITO (110)	3.35
Cu_2ZnSnS_4 (110)	CdS (110)	2.70
Cu(In,Ga)Se ₂ (001)	CdS (001)	2.10

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